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NUCLEAR LEVELS OF ²²⁸Th POPULATED IN THE DECAY OF ²²⁸Pa

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Abstract, — The decay of ²²⁸Pa to the levels of ²²⁸Th has been extensively studied by various techniques. Single spectra of γ -and K X-rays have been measured with Si(Li) and Ge(Li) detectors. A set of two Ge(Li) detectors has been used to study γ - γ and γ -K X-ray coincidences. Measurements of γ -spectra in coincidence with the selected internal-conversion lines have been carried out using a Ge(Li) detector and a six-gap magnetic β -spectrometer. Finally, spectra of internal-conversion electrons have been studied in a β -spectrometer which uses a Si(Li) detector placed together with a system of diaphragms in a homogeneous magnetic field. The decay scheme has been constructed including 39 levels of ²²⁸Th. It accounts for 111 of 160 transitions ascribed to the ²²⁸Pa activity. The electron-capture decay energy has been determined to be 2 103 $\begin{cases} +16 \\ -12 \end{cases}$ keV. The strength distribution for the electron-capture feeding of the ²²⁸Th levels is analysed in terms of nuclear models. The Coriolis mixing of four octupole bands (K = 0, 1, 2 and 3) is studied in some detail. The coupling matrix elements deduced from the experiment are compared with the results of the microscopic-model calculations.

1. Introduction. — Many features of the ${}^{228}Pa \rightarrow {}^{228}Th$ decay scheme were established and discussed by Arbman *et al.* [1] already in 1960. However, a decade later the available experimental techniques were much improved, mostly owing to the development of semiconductor detectors, and it seemed reasonable to reinvestigate this decay. The

present paper describes such new extensive studies which led to a more complete knowledge of properties of the ²²⁸Th levels. Among other results, the existence of K = 0, 2 and 3 octupole bands is confirmed and evidence is given tentatively for the previously unobserved K = 1 octupole band. The Coriolis coupling of these bands is analysed in some detail, reference being made to the microscopic-model calculations. Also the ²²⁸Pa decay energy and branching ratios for the electron-capture feeding of the ²²⁸Th levels are determined, which allows to calculate the distribution of the beta strength and to study this distribution in terms of nuclear models.

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2. Source preparations. — The 228 Pa 22 h activity was produced in the 232 Th (p, 5 n) reaction. About 1 g of metallic thorium foil was bombarded for 3 to 6 h with 100 MeV protons in the JINR synchrocyclotron at Dubna. Protactinium samples were prepared by the chemical procedure briefly described by Kurcewicz *et al.* [2]. During the measurements, apart from the 228 Pa activity and traces of its decay products, contributions from 229 Pa, 230 Pa, 232 Pa and 233 Pa were observed.

3. Singles spectra of γ -rays and internal-conversion electrons. — The low-energy part of the γ -ray spectrum was measured with a 2.5 mm thick and 5 mm in diameter Si(Li) detector, having at 60 keV a resolution (FWHM) of 1.4 keV. The energy and intensity calibration was performed using the 57 Co, 109 Cd, 169 Yb and 241 Am sources.

The spectrum in the energy range of 100 to 2 100 keV was measured with several Ge(Li) detectors. The main results were obtained using a 5.6 cm³ detector with a 1 600 channel analyzer and a 33 cm³ detector with a 4 096 channel analyzer (Fig. 1 and 2). The resolution (FWHM) of these detectors at 1 332 keV was 3.0 and 4.0 keV, respectively. Energies of the intense γ -lines were determined from those of the standard lines by counting the ²²⁸Pa source and standard sources simultaneously. In several runs, different sets of ²²Na, ⁶⁰Co, ⁸⁸Y, ^{110m}Ag, ²⁰⁷Bi or ²²⁶Ra standard sources were used. Also the energies of γ -lines of ²³⁰Pa (Kurcewicz *et al.* [2]), ²³²Pa (Kaczorowski *et al.* [3])

and ²³³Pa were used as internal-calibration standards. The energies of ²²⁸Pa lines with the assigned uncertainties of 0.2 keV or less were determined in this way. In the next step, these precisely determined energies were considered as secondary standards when finding the energies of weaker lines. For calibrating the efficiency of the detectors such sources as ⁵⁶Co, ^{110m}Ag and ²²⁶Ra were used, which have several γ -lines with relative intensities accurately known. The intensity of some ²²⁸Pa lines had to be corrected for the contribution from γ -lines of ²³⁰Pa, ²³²Pa or ²³³Pa. Some of the γ -spectra were analysed using the GIER computer code.

The energy and intensity data for γ -rays of ²²⁸Pa are listed in columns 1 and 2 of table I.

Column 3 of table I lists the data on relative intensities of the lines of internal-conversion electrons. A β -spectrometer with a 3 mm thick Si(Li) detector in a homogeneous magnetic field, described by Płochocki *et al.* [4], was used for detection of these lines in the energy range above 300 keV. Two parts of this spectrum are shown in figures 3 and 4. Since the detector was not thick enough to stop high-energy electrons completely, it was necessary to correct the line intensities for the detection efficiency. The efficiency curve was based on the data reported by Amov *et al.* [5] who studied the most intense ²²⁸Pa conversion lines, in the energy range above 800 keV, using a high resolution magnetic β -spectrometer.

To calculate the internal-conversion coefficients, the y-ray and conversion-line intensities were normalized

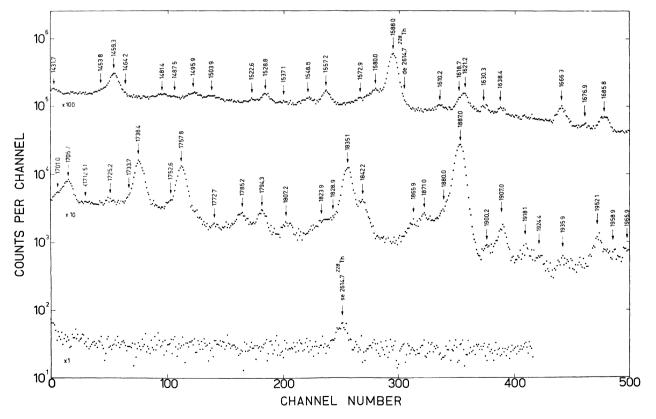


FIG. 1. — Singles γ -ray spectrum in the energy range above 1 400 keV taken using a 33 cm³ Ge(Li) detector.

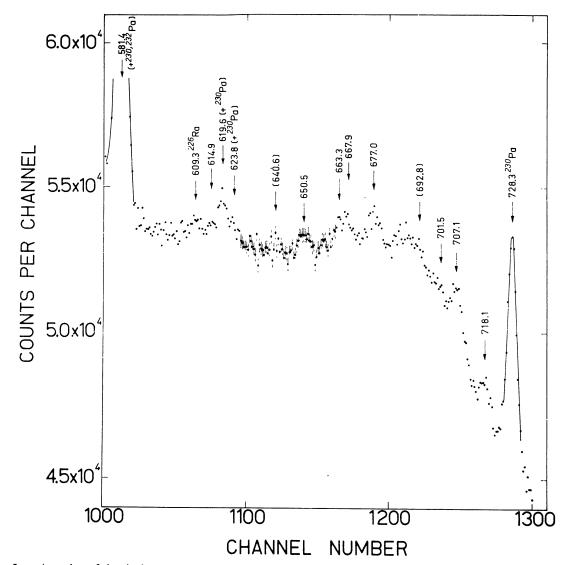


FIG. 2. — A section of the singles y-ray spectrum showing several weak lines in the energy range from 600 keV to 720 keV.

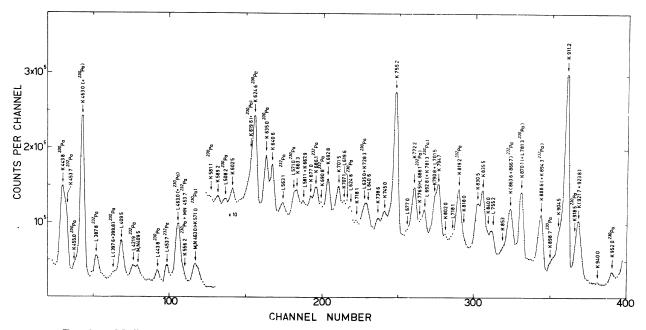


FIG. 3. — Medium-energy part of the spectrum of internal-conversion electrons measured using the β spectrometer with a Si(Li) detector in a homogeneous magnetic field.

TABLE I

Data on internal transitions in ²²⁸Th

Energy (keV)	Into ۲ rays	ensity K electrons	$\alpha_{K} \times 10^{3}$	Multipolarity	Initial level energy (^e) (keV)
$57.70 \pm 0.10 \\ 99.7 \pm 0.3 \\ 129.22 \pm 0.10 \\ 132.0 \pm 0.6 \\ 138.3 \pm 0.2$	$\begin{array}{ccccccc} & & & & & & & & & & \\ 87 & \pm & & 14 \\ 36 & \pm & 15 \\ 475 & \pm & 25 \\ 475 & \pm & 25 \\ 77 & \pm & 12 \\ 67 & \pm & 10 \end{array}$			E 2 (a) M 1 (a) E 2 (a)	57.70 1 531.9 186.92
$\begin{array}{r} 136.3 \pm 0.2 \\ 146.1 \pm 0.3 \\ 153.9 \pm 0.3 \\ 178.0 \pm 0.2 \ (^{a}) \\ 184.5 \pm 0.2 \ (^{a}) \\ 191.2 \pm 0.2 \end{array}$	57 ± 10 54 ± 6 60 ± 8 ≤ 14 ≤ 30 45 ± 7	≈ 200 (ª) 520 ± 26 (ª)	≥ 3 400≥ 4 100	M 1 E 0 + (M 1 + E 2)	1 168.18 1 122.50 1 200.4 1 153.6 378.1
$\begin{array}{c} 199.7 \pm 0.2 \\ 204.4 \pm 0.2 \\ 209.28 \pm 0.10 \\ 210.0 \pm 0.8 \ (^{b}) \\ 216.15 \pm 0.10 \end{array}$	$\begin{array}{rrrrr} 49 & \pm & 6 \\ 80 & \pm & 7 \\ 278 & \pm & 25 \\ \approx & 60 \ (^{\flat}) \\ 145 & \pm & 15 \end{array}$	88 ± 5 (ª)	75 ± 9	E 1	1 168.18 1 226.56 396.09 1 642.5 1 168.18
$\begin{array}{l} (219.8 \ \pm \ 0.6) \\ 223.61 \ \pm \ 0.10 \\ (240.2 \ \pm \ 0.8) \ (^{\flat}) \\ 255 \ \pm \ 1 \ (^{\flat}) \\ 270.23 \ \pm \ 0.10 \end{array}$	$pprox 30 \ 153 \ \pm \ 13 \ pprox 17 \ (^b) \ pprox 50 \ (^b) \ 350 \ \pm \ 17 \ $	$\begin{array}{rrrr} 780 & \pm \ 40 \ (^{a}) \\ \\ 48 & \pm \ 3 \ (^{a}) \end{array}$	$1\ 210 \pm 120$ 33 ± 3	M 1 + E 2	1 944.5 1 450.14 618.3 1 687.3 327.74
$\begin{array}{c} 278 \ (^{a})\\ 281.87 \ \pm \ 0.10\\ 327.64 \ \pm \ 0.10\\ 332.36 \ \pm \ 0.10\end{array}$	$ \begin{array}{c} \leqslant 20 \\ 205 \\ \pm 11 \\ 660 \\ \pm 50 \end{array} \\ 262 \\ \pm 24 \end{array} $	$\begin{array}{rrrr} 24 & \pm & 1 \ (^a) \\ 512 & \pm & 26 \ (^a) \\ 116 & \pm & 6 \ (^a) \end{array}$		(M 1) M 1 + E 2 Ej1 and E 2	1 450.14 327.74 and 1 450.14 519.28
$\begin{array}{c} 338.32 \pm 0.10 \\ 341.1 \pm 0.3 \\ 409.51 \pm 0.10 \\ 449.3 \pm 0.6 \\ (461 \pm 1) \ (^{\circ}) \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	E 1 E 2 E 2	396.06 1 432.00 1 432.00 1 900.00 1 892.6
$\begin{array}{c} 463.00 \pm 0.10 \\ 481.3 \pm 0.8 \\ (498 \pm 1) \ (^{\flat}) \\ 525.0 \pm 0.6 \\ 547.5 \pm 0.6 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	394 ± 80	42 ± 9	E 2	1 432.00 1 450.14 1 450.14
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$				951.9 1 531.9 969 1 645.7
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\leqslant 13 \\ \leqslant 10 \\ 16 \pm 4 \\ 50 \pm 20 \\ 13 \pm 2$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$		M 1 + E 2 M 1 M 1 + E 2	1 015.6 951.9
$\begin{array}{rrrr} 640.6 & \pm \ 0.5 \\ 650.5 & \pm \ 0.4 \\ 663.3 & \pm \ 0.6 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	260 ± 140 30 ± 10	M 1 M 1 + E 2	and/or 1 645.7 969 1 064.1
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 E 2 E 0 + $(M_k 1 + E)$ M 1 M 1 + E 2 (E 2)	1 645.7 1 892.6 1 724.1 1 226.56 1 687.3
$\begin{array}{rrr} 726.2 & \pm \ 0.8 \ (^{b}) \\ 738.6 & \pm \ 0.6 \\ 745.0 & \pm \ 0.6 \\ 750.5 & \pm \ 0.5 \end{array}$	$\begin{array}{c} \approx 80 \ (^{\flat}) \\ \leqslant 15 \\ \\ \leqslant 20 \\ 35 \ \pm 7 \end{array}$	$\begin{array}{rrrr} 4.5 & \pm & 1.0 \\ 8.6 & \pm & 1.3 \end{array}$	≥ 71 ≥ 90	M 1 M 1	1 122.50 1 892.6 and/or 1 938.6 1 944.5 1 925.3
$\begin{array}{c} 755.18 \pm 0.10 \\ 772.17 \pm 0.10 \\ 776.5 \pm 0.2 \\ (782.0 \pm 0.6) (^{b}) \\ 790.8 \pm 0.3 \end{array}$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	M 1 E 2 + M 1 E 2 M 1	1 724.1 1 168.18 969.05 1 944.5

TABLE I (continued)

	TABLE I (commune)					
Energy (keV)	Inte γ rays	nsity K electrons	$\alpha_K \times 10^3$	Multipolarity	Initial level energy (^e) (keV)	
$794.7 \pm 0.2 $	334 ± 15	$21.3 ~\pm~ 3.4$	15 ± 3	E 2 + M 1	1 122.50	
$(796 \pm 1) (^{b}) \\ 802.0 \pm 0.5$	≈ 20 (^b) ≤ 15	$2.8 ~\pm~ 0.7$	> 44	M 1	1 174.7	
$\begin{array}{r} 802.0 \\ \pm 0.3 \\ 818.0 \\ \pm 0.8 \end{array}$	100 ± 50	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\geqslant 44$ 8.8 \pm 5.1	M 1 E 2		
$823.5 \pm 1.0 $	≈ 40					
$\begin{array}{rrrr} 830.5 & \pm \ 0.3 \\ 835.5 & \pm \ 0.3 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 E 2	1 226.56 1 022.36	
$840.0 \pm 0.4 $	$170~\pm~10$	6.5 ± 1.6	9.1 ± 2.4	E 2 E 2	1 168.18	
853 ± 1	≤ 10 17(+ 10	2.0 ± 0.9	≥ 47	M 1	1 944.5	
$\begin{array}{rrrr} 870.1 & \pm \ 0.4 \\ 884.2 & \pm \ 0.5 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	26.8 ± 2.7	36 ± 4	M 1	1 892.6 1 900.0	
$888.6 \hspace{0.2in} \pm \hspace{0.2in} 0.5$	130 ± 30	$2.9 ~\pm~ 0.9$	5.3 ± 2.0	E 1		
$\begin{array}{rrr} 894.3 & \pm \ 0.5 \\ 904.5 & \pm \ 0.3 \end{array}$	$\begin{array}{rrrr} 440 & \pm \ 150 \\ 480 & \pm \ 40 \end{array}$	14.4 ± 2.2	7.1 ± 1.3	E 2	951.9	
911.23 \pm 0.10	2670 ± 110	14.4 ± 2.2 100	$7.1 \pm 1.3 \\ 8.9 (^{d})$	E 2 E 2	1 091.4 969.05	
921.7 \pm 0.3 (°)	≤ 100	24.1 ± 3.2 (°)	≥ 50	M 1	1 944.5	
$\begin{array}{rrr} 923.8 & \pm \ 0.5 \ (^{c}) \\ 940.0 & \pm \ 0.8 \end{array}$	$\begin{array}{ccc} \approx & 60 \ (^{b}) \\ 100 & \pm & 50 \end{array}$	$\begin{array}{rrrr} 10.8 & \pm & 3.1 \ (^{o}) \\ 1.0 & \pm & 0.2 \end{array}$	pprox 43 2.3 \pm 1.2	M 1 E 1	1 892.6 1 892.6	
$945.6 \pm \ 0.8$	300 ± 100	1.0 ± 0.2	2.5 1.2		1 092.0	
957.8 \pm 0.8	≈ 100	(9 27		F 2	1 015.6	
$\begin{array}{rrr} 964.6 & \pm \ 0.3 \\ 969.11 & \pm \ 0.10 \end{array}$	$\begin{array}{rrrr} 1 \ 680 & \pm \ 200 \\ 2 \ 200 & \pm \ 400 \end{array}$	$\begin{array}{ccc} 68 & \pm 27 \\ 50 & \pm 20 \end{array}$	$\begin{array}{rrrr} 9.6 \ \pm & 4.0 \\ 5.4 \ \pm & 2.5 \end{array}$	E 2 E 2	1 022.36 969.05	
					and 1 938.6	
$\begin{array}{rrr} 975.0 & \pm \ 0.3 \\ 987.8 & \pm \ 0.2 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	15 ± 3	14 ± 3	E 2 + M 1	1 944.5	
907.0 ± 0.2	40 ± 2				1 174.7 and 2 010.0	
$1\ 018.6\ \pm\ 0.3$	35 ± 5					
$\begin{array}{rrrr} 1 \ 033.2 & \pm \ 0.3 \\ 1 \ 039.9 & \pm \ 0.3 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$				1 091.4 1 226.56	
$1\ 046.1\ \pm\ 0.8$	6 ± 2				1 220.50	
$\begin{array}{rrrr} 1 \ 054.4 & \pm \ 0.5 \\ 1 \ 065.2 & \pm \ 0.5 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$				1 450.14	
$1\ 005.2\ \pm\ 0.5$ $1\ 070.2\ \pm\ 0.5$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$				1 122.50	
$1\ 096.0\ \pm\ 0.8$	$egin{array}{ccccc} 19 & \pm & 4 \ 5 & \pm & 2 \ 3 & \pm & 1 \end{array}$				1 153.6	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$egin{array}{cccc} 3 & \pm & 1 \ 72 & \pm & 3 \end{array}$	$0.55~\pm~~0.11$	1.8 ± 0.4	E 1	1 168.18	
$1\ 118.6\ \pm\ 0.6$	7 ± 1					
$\begin{array}{rrrr} 1 164.4 &\pm 0.6 \\ 1 184.4 &\pm 0.6 \end{array}$	$egin{array}{cccc} 12&\pm&1\ 4&\pm&2 \end{array}$	$\begin{array}{rrrr} 0.70 \ \pm & 0.07 \\ 0.22 \ \pm & 0.06 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 + M 1 E 2, M 1	1 580.3	
$1 \hspace{.1cm} 194.7 \hspace{.1cm} \pm \hspace{.1cm} 1.0 \hspace{.1cm}$	3 ± 1	0.122 ± 0.000 0.15 ± 0.05	12 ± 6	E 2, M 1	1 500.5	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	20 ± 0.2	16 05	E 2	1 (40 5	
1240.4 ± 0.2 1253.1 ± 0.6	3 + 1	$\begin{array}{rrrr} 2.9 & \pm & 0.3 \ (°) \\ 0.29 & \pm & 0.08 \ (°) \end{array}$	$\begin{array}{rrrr} 4.6~\pm&0.5\\ 23~\pm&10\end{array}$	E 2 M 1	1 642.5 1 580.3	
$1\ 273.0\ \pm\ 0.6$	13 ± 2					
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccc} 20 & \pm & 1 \\ 19 & \pm & 2 \end{array}$	$\begin{array}{rrrr} 0.72 \ \pm & 0.07 \\ 0.21 \ \pm & 0.07 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 + M 1 E 1 + M 2		
$1 \hspace{.1cm} 311.0 \hspace{.1cm} \pm \hspace{.1cm} 0.6 \hspace{.1cm}$	9 ± 3	0.26 ± 0.06	6.9 ± 2.7	E 2 + M 1		
$\begin{array}{rrrr} 1 \ 420.6 & \pm \ 0.6 \\ 1 \ 431.7 & \pm \ 0.6 \end{array}$	$egin{array}{cccc} 16 &\pm &1\ 23 &\pm &2 \end{array}$	$\begin{array}{rrrr} 0.34 \ \pm & 0.08 \\ 0.51 \ \pm & 0.09 \end{array}$	5.1 ± 1.2	E2		
1451.7 ± 0.0 1453.8 ± 0.6	$egin{array}{cccc} 23&\pm&2\\ 20&\pm&1 \end{array}$	0.31 ± 0.09 0.13 ± 0.06	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 E 1		
1459.3 ± 0.2	120 ± 5	$1.73~\pm~0.30$ (°)	3.4 ± 0.6	E 2	1 645.7	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$egin{array}{cccc} 10&\pm&5\\ 15&\pm&1 \end{array}$	$0.30~\pm~~0.10$	4.7 ± 1.6	E 2		
$1\ 487.5\ \pm\ 0.6$	≈ 9					
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	28 ± 2	0.35 ± 0.10		E 1, E 2	1 000 0	
1503.9 ± 0.0 1522.6 ± 0.6	$egin{array}{cccc} 16 &\pm &1 \ 8 &\pm &1 \end{array}$	0.16 ± 0.07	2.4 ± 1.1	E 1	1 900.0 1 580.3	
$1 \hspace{.1cm} 528.8 \hspace{.1cm} \pm \hspace{.1cm} 0.4$	29 \pm 2	0.40 ± 0.12	3.3 ± 1.1	E 1, E 2	1 925.3	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccc} 2 & \pm & 1 \\ 16 & \pm & 1 \end{array}$	$0.26~\pm~0.07$	2.9 ± 1.5	E 2	1 724.1 1 944.5	
$1\ 557.2\ \pm\ 0.3$	46 ± 2	1.1 ± 0.2		E 2 + M 1		
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrr} 0.34 \ \pm & 0.11 \\ 0.45 \ \pm & 0.14 \end{array}$	$\begin{array}{rrrr} \textbf{2.8} \ \pm & \textbf{1.0} \\ \textbf{1.7} \ \pm & \textbf{0.5} \end{array}$	E 1, E 2 E 1	1 900.0 1 580.3	
$1\ 588.0\ \pm\ 0.2$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$		(E 2)	1 580.3 1 645.7	
$1\ 610.2\ \pm\ 0.4$	12 ± 1	0.58 ± 0.13	11 ± 3	M 1		
$\begin{array}{rrrr} 1 \ 618.7 & \pm \ 0.4 \\ 1 \ 621.2 & \pm \ 0.4 \end{array}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrr} \textbf{0.46} \ \pm & \textbf{0.15} \\ \textbf{1.3} \ \pm & \textbf{0.3} \end{array}$	$\begin{array}{rrrr} 4.7 \pm & 1.6 \\ 7.2 \pm & 1.7 \end{array}$	E 2 + M 1 M 1	1 676.3	
	·					

Nº.	2-3
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IABLE I (continued)						
Energy (keV)	Inter y rays	nsity K electrons	$\alpha_K \times 10^3$	Multipolarity	Initial level energy (^e) (keV)	
$\begin{array}{rrrr} 1 \ 630.3 & \pm \ 0.4 \\ 1 \ 638.4 & \pm \ 0.4 \\ 1 \ 666.3 & \pm \ 0.2 \\ 1 \ 676.9 & \pm \ 0.6 \end{array}$	$egin{array}{ccccc} 17 & \pm & 2 \ 16 & \pm & 2 \ 31 & \pm & 2 \ 5 & \pm & 1 \end{array}$	$\begin{array}{c}$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	E 2 M 1	1 687.3 2 016.4 1 724.1 1 676.3	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 24 & \pm & 2 \\ 10 & \pm & 1 \end{array}$	$0.33~\pm~0.12$	3.2 ± 1.2	E 2	1 01 012	
$\begin{array}{rrrr} 1 \ 705.7 & \pm \ 0.4 \\ (1 \ 712.5 & \pm \ 0.6) \\ 1 \ 725.2 & \pm \ 0.6 \end{array}$	$\begin{array}{cccc} 36 & \pm & 2 \\ pprox 2 \end{array}$	0.9 ± 0.2	5.9 ± 1.3	E 2 + M 1	1 892.6 1 900.0	
$egin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	106 ± 5	$2.60~\pm~~0.26$ (°)	5.8 ± 0.7	E 2 + M 1	1 925.3 1 938.6	
$\begin{array}{rrrrr} 1 \ 757.8 & \pm \ 0.2 \\ 1 \ 772.7 & \pm \ 0.6 \end{array}$	$egin{array}{ccccc} 5 & \pm & 1 \ 90 & \pm & 5 \ 5 & \pm & 1 \end{array}$	$1.50~\pm~~0.30$ (°)	3.9 ± 0.8	E 2 + M 1	1 944.5	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$\begin{array}{cccc} 14 & \pm & 1 \\ 15 & \pm & 1 \end{array}$	$0.28~\pm~~0.14$	4.7 ± 2.4	E 2 + M 1	1 842.9	
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$5.4 \pm 0.4 \\ 6 \pm 1 \\ 9 \pm 2$	$0.16~\pm~~0.08$	7.0 ± 3.6	E 2 + M 1	1 993.9 2 010.0 2 016.4	
$\begin{array}{ccc} 1 \ 835.1 & \pm \ 0.2 \\ 1 \ 842.2 & \pm \ 0.3 \end{array}$	$\begin{array}{cccc} 106 & \pm & 6 \\ 27 & \pm & 2 \end{array}$	$\begin{array}{rrr} 1.56 \ \pm & 0.23 \ (^{e}) \\ 0.60 \ \pm & 0.20 \ (^{e}) \end{array}$	$\begin{array}{rrrr} 3.5 \ \pm & 0.6 \\ 5.2 \ \pm & 1.8 \end{array}$	E 2 + M 1 E 2 + M 1	1 892.6 1 900.0 and/or 1 842.9	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{cccccccccccccccccccccccccccccccccccc$	$0.18~\pm~0.09$	2.8 ± 1.4	(E 2)	1 000 (
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$2.97~{\pm}~~0.45$ (°)	2.7 ± 0.4	E 2	1 938.6 1 944.5 1 900.0 1 965.0	
$\begin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$	$egin{array}{rrrrrrrrrrrrrrrrrrrrrrrrrrrrrrrr$				1 993.9 2 010.0 2 016.4 1 965.0	

TABLE I (continued)

(a) Data from reference [1]; K-electron intensities are renormalized by the present authors and assumed to be accurate within about 5 % (see caption to table I in ref. [1]).
(b) Data from our coincidence measurements.

(c) Data from reference [5].

(d) Based on assumption of E 2 multipolarity for the 911.23 keV transition.

(*) For the transitions placed in the decay scheme, see Fig. 10*a* and 10*b*. (*f*) The intensity of 1 000 units corresponds to 7.4 % of the ²²⁸Pa decay rate.

by assuming the theoretical (Hager and Seltzer [6]) E 2 internal-conversion coefficient of 8.9×10^{-3} for the 911.23 keV transition. The columns 4 and 5 in table I list, respectively, the values of the internalconversion coefficients and the multipolarity assignments, deduced by comparing these values with the theoretical ones [6].

4. Coincidence measurements. — The γ -spectra in coincidence with internal-conversion electrons were measured using a 5.6 cm³ Ge(Li) detector and a six-gap magnetic β -spectrometer. The conditions of these measurements were identical to those in ²³⁰Pa studies of Kurcewicz et al. [2]. The results are presented in figures 5 and 6, and in table II.

TABLE II

	Results of the ey coincidence studies					
Selected conversion line	Energies (keV) and intensities (in brackets) of coincident γ lines					
L 57.7	209 (250 \pm 50), 270 (254 \pm 57), 282 (160 \pm 60), 332 (290 \pm 70), 338 (890 \pm 100), 410 (950 \pm 110), 463 (1 330 \pm 150), 498 ? (\approx 100), 830 (\approx 250), 835 and/or 840 (\approx 390), 894 ? (\approx 150), 905 (\approx 420), 911 (\equiv 2 670), 965 (1 780 \pm 330)					
L 129.3	191 (41 \pm 15), 209 (264 \pm 23), 224 (39 \pm 12), 240 ? (\approx 17), 255 (32 \pm 11), 282 (42 \pm 14), 332 (240 \pm 25), 341 (205 \pm 24), 410 (360 \pm 41), 707 (105 \pm 47), 772 ? (\approx 110), 782 ? (\approx 105), 796 ? (\approx 20), 830 (180 \pm 85), 835 (\equiv 450), 870 (\approx 50), 905 (390 \pm 65), 988 ? (\approx 90)					

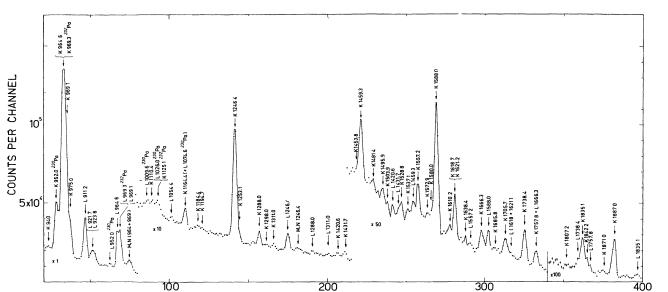


FIG. 4. — High-energy part of the spectrum of internal-conversion electrons measured using the β spectrometer with a Si(Li) detector in a homogeneous magnetic field.

NUMBER

CHANNEL

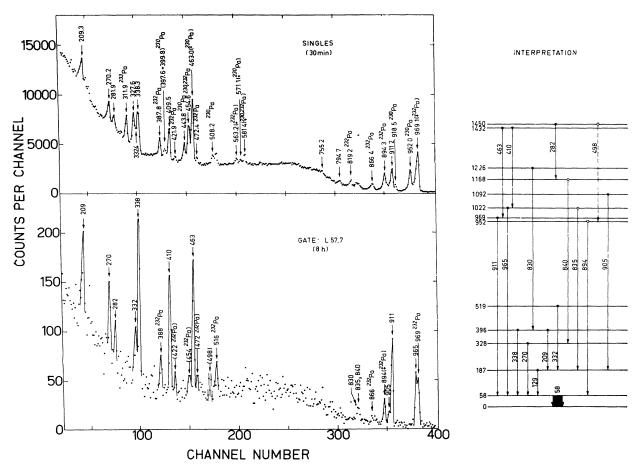


FIG. 5. — Gamma-ray spectrum coincident with the L 57.7 internal-conversion line and its interpretation.

The measurements of γ - γ coincidences were performed using a spectrometer with two Ge(Li) detectors. These detectors were placed at an angle of about 60° with regard to the source position. Absorbers were used to stop γ -rays scattered from one detector in the

direction of the other one. A fast-slow coincidence circuit with a time-to-amplitude converter was applied. Two coincidence spectra were recorded simultaneously. One was gated by a selected γ -line and a portion of the Compton continuum, and the second one by a

400

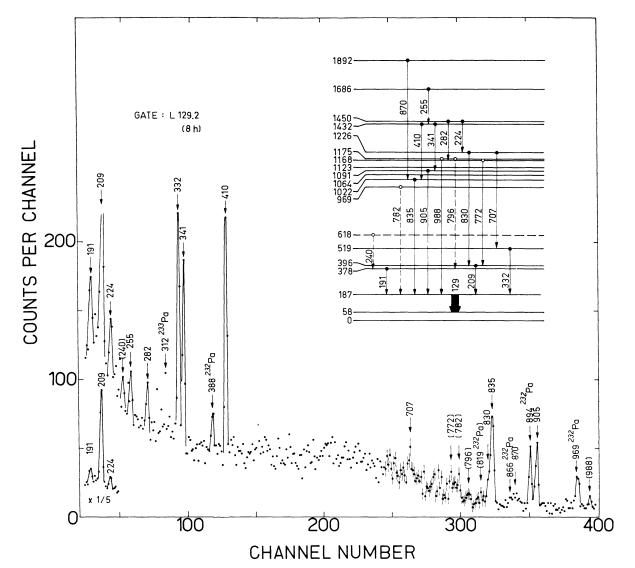


FIG. 6. - Gamma-ray spectrum coincident with the L 129.2 internal-conversion line and its interpretation.

section of this continuum above the line. Thus the pure coincidence effect due to the selected transition could be deduced. In figure 7 a typical pair of coincidence spectra is shown by way of example. As it may be seen from table III, the coincidence data were obtained for five gating lines.

TABLE III

Results of the yy coincidence studies

energy interval	ed γ lines	
(keV)	(keV)	Energies (keV) and intensities (in brackets) of coincident γ lines
260-280	270	282 (71 ± 27), 328 (312 ± 62), 795 (\equiv 334), 840 (153 ± 44)
330-350	338	224 (40 \pm 8), 282 (35 \pm 12), 328 (26 \pm 13), 480 (36 \pm 17), 556 ? (\approx 50), 573 (74 \pm 20), 619 (31 \pm 16), 668 ? (\approx 50), 726 (51 \pm 20), 772 (\equiv 198), 830 (250 \pm 40), 870 (24 \pm 14), 1 246 (100 \pm 36) 205 ($=$ 480), 1 232 (120 \pm 50)
	341	905 (\equiv 480), 1 033 (120 \pm 50)
400-420	410	129 (174 \pm 34), 210 (65 \pm 20), 461 ? (145 \pm 50), 520 (60 \pm 36), 835 (470 \pm 95), 965 (\equiv 1 680)
790-810	795	270 (\equiv 350), 328 (820 \pm 250)
900-920	905 912	341 (124 \pm 30). 154 (53 \pm 21), 328 (47 \pm 25), 463 (\equiv 2 200), 481 ? (\approx 60), 707 (\approx 20), 755 (290 \pm 70), 923 (\approx 60), 975 (180 \pm 60)

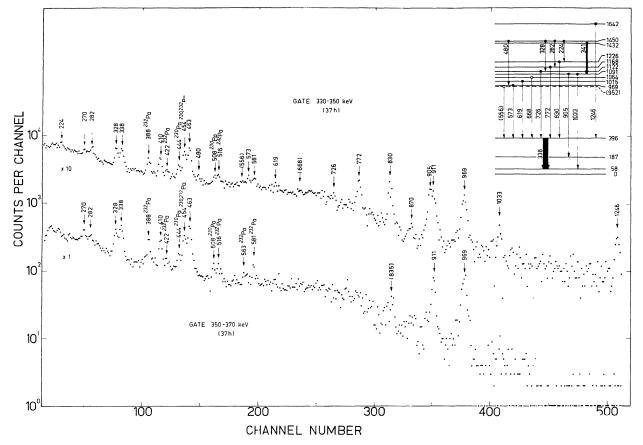


FIG. 7. — Gamma-ray spectrum measured in coincidence with the 338.3 keV γ -line and its interpretation. The spectrum coincident with the fraction of the Compton distribution is shown or comparison. The gating transition was taken using a 4.5 cm³ Ge(Li) detector, and the coincident spectrum-using a 13 cm³ Ge(Li) detector.

5. Determination of the electron-capture decay energy. — The method used to determine the ²²⁸Pa decay energy $Q_{\rm EC}$ was similar to that described by Kurcewicz *et al.* [2] for the decay of ²³⁰Pa. It takes account of the well known dependence of the relative *K*-capture probability P_K upon the electron-capture transition energy Q.

The ratio of the intensities was determined experimentally for the 1 588 and 1 887 keV γ -lines from the singles γ -ray spectrum and the spectrum coincident with the K X-rays (cf. Fig. 8). From these data it was possible to calculate the ratio of the P_K values for the electron-capture transitions to the levels at 1 646 and 1 945 keV (cf. Fig. 9). This ratio was found to be 0.47 \pm 0.11, where from the energy for the transition to the 1 945 keV level is equal to 158 $\begin{cases} +16\\ -12 \end{cases}$ keV. Thus, the decay energy is $Q_{\rm EC} = 2103 \begin{cases} +16\\ -12 \end{cases}$ keV.

6. The decay scheme. — The 228 Pa decay scheme shown in figures 10*a* and 10*b* is an extension of that published by Arbman *et al.* [1]. The twenty one levels of 228 Th reported in reference [1] have been confirmed, and new levels at 618, 952, 969 (2⁻), 1 016, 1 064, 1 175, 1 200, 1 580, 1 642, 1 676, 1 843, 1 900, 1 925, 1 939, 1 965, 1 994, 2 010 and 2 016 keV have been

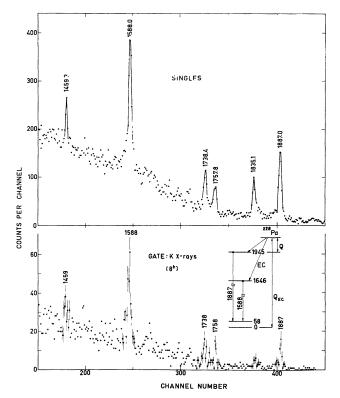


Fig. 8. — Section of the singles γ -ray spectrum and of the spectrum coincident with K X-rays. In the insert a fragment of the decay scheme is shown.

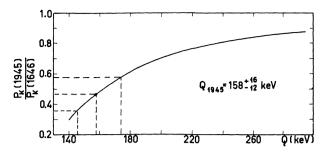


FIG. 9. — Dependence of the P_K (1945)/ P_K (1646) ratio as a function of the energy Q of the EC transition to the 1945 keV level. The experimental value of this ratio, obtained in the coincidence measurements (see Fig. 8), is also presented.

found. It has recently been shown that many of these new levels are populated also in the β^- -decay of ²²⁸Ac (Dalmasso and Maria [7], Herment and Vieu [8], Herment [9]).

The construction of the ²²⁸Th level scheme is based on the transition energy fits, searched with the use of a special computer program, and on the results of the coincidence experiments. The decay scheme includes 111 of the total number of 160 transitions ascribed to the ²²⁸Pa activity. The multipole character established for numerous transitions allows to define the parity for the majority of the ²²⁸Th levels and to assign spin values to many of them. With the knowledge

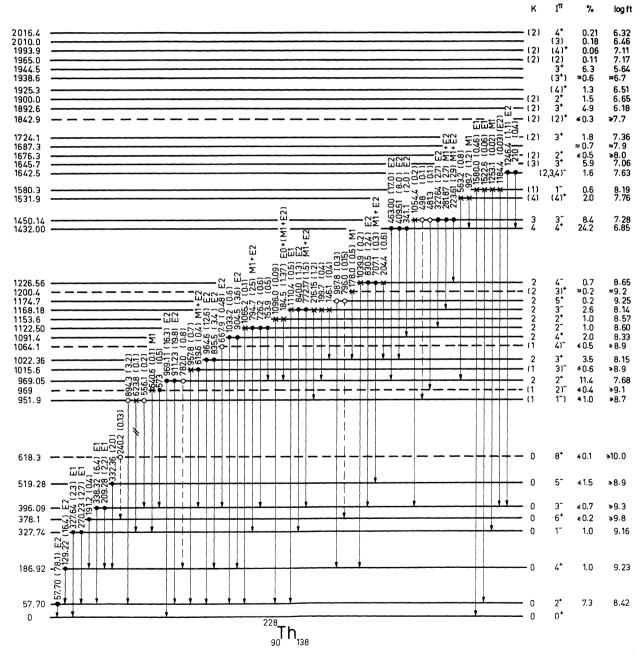


FIG. 10a. — Scheme of the ²²⁸Pa decay to ²²⁸Th levels. The spacing of close-lying levels is not up to scale. The internal transitions, whose position in the decay scheme has been established or suggested by coincidence measurements, are marked with full and open circles, respectively. Crosses refer to the transitions placed on the basis of the energy fit alone. Transitions placed in two alternative positions are marked with two bars. The intensities (in parentheses) of the transitions are given in per cent of the ²²⁸Pa decays. All energies are in keV.

Nº 2-3

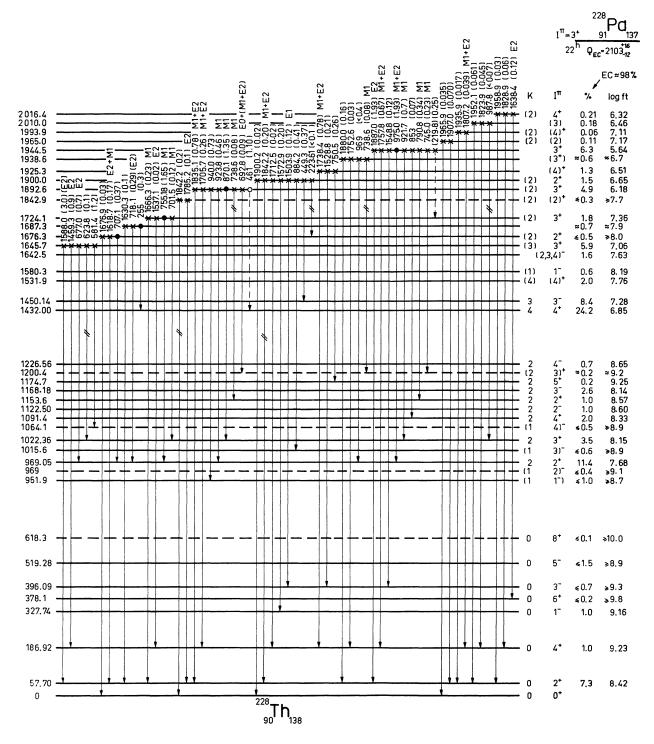


FIG. 10b. — Scheme of the ²²⁸Pa decay to ²²⁸Th levels (for comments cf. caption to Fig. 10a).

of the $Q_{\rm EC}$ value, of the ²²⁸Pa half-life and of the EC branching ratios it was possible to calculate log ft values. The low log ft value for the transition to the 1944 keV 3⁺ level and the direct EC feeding of the 2⁺ and 4⁺ levels indicates the spin and parity 3⁺ for the ²²⁸Pa ground state. This is in agreement with the assignment proposed by Arbman *et al.* [1]. The assignments of the K quantum numbers to the levels at lower excitation energy results from the interpre-

tation of these levels in terms of nuclear models (cf. section 7).

The arguments taken into account when constructing the decay scheme can be easily reproduced if use is made of the information contained in tables I-III. It has been decided, therefore, to omit in this section any comments on the existence of individual levels in ²²⁸Th and on the spin-parity assignment. In the next section, however, brief comments can be

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found on several, mostly tentative, levels whose identification is important for the verification of the applicability of the deformed-nuclei theory to low-energy excitations in ²²⁸Th.

The balance of the intensities for the decay scheme is based on the assumption that the EC process occurs in 98 % of the ²²⁸Pa decays (see tables by Lederer et al. [10]) and that there is no EC feeding of the ²²⁸Th ground state. The total transition intensities have been calculated with the use of the theoretical internal conversion coefficients of Hager and Seltzer [6]. The intensity of the 49 transitions not included in the decay scheme corresponds to that of about 11 % of the total EC decays. Including of these transitions in the decay scheme would result in a change of the EC branchings and log ft values with respect to those given in figures 10a and 10b. This, however, could hardly affect the spin and parity assignment to the ²²⁸Pa ground state. We believe also that the qualitative conclusions of section 7.3 on the beta-strength distribution would not be changed.

7. **Discussion.** — The properties of the 228 Th levels are discussed in this section in terms of the models developed for the deformed nuclei. For a general presentation of these models the reader may refer to Nathan and Nilsson [11] and Soloviev [12].

7.1 POSITIVE-PARITY STATES BELOW 1 500 keV. — The interpretation of the low-energy 228 Th levels of positive parity is illustrated in figure 11.

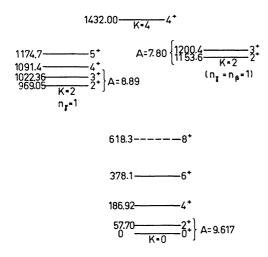


FIG. 11. — Interpretation of the 228 Th positive-parity states below 1 500 keV. Notation : A — moment-of-inertia parameter in keV, n β and n γ — quantum numbers of the β and γ oscillations, respectively.

The ground-state band is shown with four rotational levels. The 6^+ level introduced by Arbman *et al.* [1] as uncertain is now well proved by coincidence data. The evidence for the 8^+ level is only tentative. The calculations based on the rotational formula, with

three parameters determined from the position of the 2^+ , 4^+ and 6^+ levels, for the 8^+ level yield the energy of 624 keV. This is not far from the tentatively given experimental value.

The new level at 1 174.5 keV, also observed in the decay of ²²⁸Ac (Herment [9]), is interpreted as the spin-parity 5⁺ member of the γ -vibrational band. The β -vibrational 0⁺ level, introduced by Lederer *et al.* [13] at 0.83 MeV, has not been found to be fed in the decay of ²²⁸Pa.

The 1 153.6 keV level and the γ -vibrational bandhead state are linked by the E 0 transition (¹). Hence, for the 1 153.6 keV level we have $KI^{\pi} = 22^+$, and therefore this level could be interpreted as a two-phonon $(\beta + \gamma)$ -state. Its energy is, however, significantly lower than the sum of the energies of the β -and γ -vibrational levels. Similar $K^{\pi} = 2^+$ levels have been observed in ²³⁰Th (see ref. [2] and earlier papers quoted there) and in ²³⁴U (Bjørnholm *et al.* [15]).

The assignment of $KI^{\pi} = 44^+$ to the 1 432.0 keV level has been concluded from the ratios of the E 2 reduced probabilities of the transitions to the γ -band. The experimental ratios

B (E 2,
$$44^+ \rightarrow 22^+$$
) : B (E 2, $44^+ \rightarrow 23^+$) :
B (E 2, $44^+ \rightarrow 24^+$) =
 $(1.12 \pm 0.08) : 1 : (0.64 \pm 0.05)$

are in agreement with the theoretical ones, 1.04:1:0.61, obtained from the Mihailov formula [16] with the parameter a = 0.030. No agreement is achieved when $KI^{\pi} = 33^+$ is assumed for the decaying state. A possible two-quasiparticle configuration of this state is discussed in section 7.3.

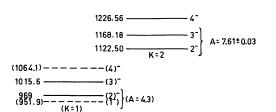
7.2 OCTUPOLE BANDS. — The negative-parity states observed in 228 Th below 1 500 keV are interpreted as members of the octupole bands (cf. Fig. 12).

The existence of the K = 0, 2 and 3 octupole bands observed by Arbman *et al.* [1] has been confirmed in the present study. The reduced branching ratios for the transitions deexciting these bands are found to be consistent with the adopted interpretation, provided that the analysis is carried out with the use of the Mihailov formula [16].

A new octupole band with K = 1 is proposed to have its first four levels at 951.9, 969, 1015.6 and 1064.1 keV.

The 951.9 keV level. — To this level a spin and parity 1⁻ may be assigned only tentatively. The 1⁻ level may be expected to decay to the 0⁺ ground state. The fact that such a transition has not been observed in the γ -ray spectrum does not necessarily contradict this assignment, since it can be masked by the strong

⁽¹⁾ The EO character of the 184.5 keV transition was earlier noticed by Bjørnholm [14]. Reference should also be made to the publication by Herment and Vieu [8].



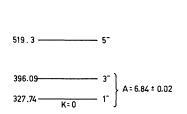


FIG. 12. — Octupole bands in ²²⁸Th. A — moment-of-inertia parameter in keV.

951.92 keV line of ²³⁰Pa. This level may be interpreted as the head of a new band.

The 969 keV level. — The $KI^{\pi} = 12^{-1}$ level at 969 keV is proposed mainly because of its decay to the 396.09 keV $KI^{\pi} = 03^{-1}$ and 327.74 keV $KI^{\pi} = 01^{-1}$ levels. The 573 keV transition appears in coincidence with the 338.32 keV line. We believe that this is a M 1 or E 2 transition from the 969 keV 12⁻¹ level rather than a K-forbidden E 1 transition from the 969.05 keV 22⁺ level. The 640.6 keV transition is placed between the hypothetical 2⁻¹ and the 327.74 keV 1⁻¹ levels for the energy-fit reason, without any coincidence support. However, its M 1 character is consistent with our interpretation.

The 1 015.6 keV level. — The existence of this level is well proved by coincidence data. Also it is shown clearly that its parity is negative. This level is a candidate to be interpreted as a 3⁻ member of the $K^{\pi} = 1^{-}$ band.

The 1064.1 keV level. — This level could be the 4⁻ state of the $K^{\pi} = 1^{-}$ band, but its existence should be considered as tentative.

If the $K^{\pi} = 1^{-}$ band has the level energies as suggested here, it is for the first time possible, anyway to the present author's knowledge, to get some information on the Coriolis interaction of all four one-phonon octupole bands directly from experiment. In the calculations performed it has been necessary to consider several quantities as free parameters : (i) three coupling matrix elements A_{01} , A_{12} and A_{23} (the subscripts

referring to the K values) and (ii) the unperturbed energy of the $KI^{\pi} = 33^{-}$ level. The moment-of-inertia parameter A has been assumed to be the same for all bands. For the sake of simplicity, the relations between A, A_{01} and A_{12} , derived from the known positions of the I = 1 and I = 2 levels in the K = 1 and 2 bands, have been used in the calculations, the small experimental errors being neglected. A fit of the calculated energies to the experimental ones has been performed for the four I = 3 levels. The values of the matrix elements found in this way are listed in table IV, and compared with the theoretical expectations.

TABLE IV

Matrix elements of the Coriolis interaction between octupole bands (in keV)

$A_{K,K+1}$	Experiment (^a)	(^b)	Theory (°)	(^d)
A ₀₁	31.5	32.0	42.3	7.2
A_{12}	21.5	29.2	41.9	44.1
A ₂₃	35.5	22.6		60.4

(a) From the analysis of the level-energy spacings. The inertial parameter A = 8.34 keV.

(^b) The spherical-limit values calculated according to the formula given by Neergård and Vogel [18].

(^c) Based on microscopic calculations by Zheleznova *et al.* [17].

(^d) Based on microscopic calculations by Błocki [19].

For the parameter values resulting from the bestfit procedure adopted here, the energies of the 1⁻ and 2⁻ levels are reproduced exactly and those of the 3⁻ levels — within ± 0.6 keV. The agreement between the calculated and experimental energies for the 4⁻ levels is worse. For the K = 1 and K =2 bands, the calculated energies of the 4⁻ levels are 1 039.4 and 1 234.2 keV, respectively, which is pretty far from the experimental values.

The experimental energies of the levels of the ²²⁸Th octupole bands are compared in table V with the theoretical results based on the microscopic-model calculations carried out by different authors.

7.3 DISTRIBUTION OF THE BETA STRENGTH IN THE 228 Pa $\rightarrow ^{228}$ Th DECAY. — In the considerations of the 228 Pa EC decay given below, the most probable configuration of the Nilsson-model orbitals, p 530 \uparrow and n 752 \uparrow , has been assumed for the ground state of this nucleus, in agreement with Arbman *et al.* [1].

The rather low probability (low strength) observed for the $\Delta K = 3$ EC decay to the levels of the groundstate band and of the $K^{\pi} = 0^{-}$ octupole band of ²²⁸Th is related to the effect of K forbiddenness. Also, the $\Delta K = 2$ value for the 1st forbidden transitions to the levels of the possible $K^{\pi} = 1^{-}$ octupole band eliminates all matrix elements, except the unique

		Experiment	Energy (keV)			
K	Ι ^π	Experiment this work	Ref. [17]	Ref. [20]	Ref. [18]	Ref. [21]
0	1 ⁻ 3 ⁻	328 396	350	620	360 400	460
	5-	519			400	
1	1-	952	1 110	1 020	880	1 100
	2-	969			890	
	3- 4-	1 016 1 064			980	
2	2-	1 122	1 600	1 160	1 130	1 400
2	3-	1 122	1 000	1 100	1 200	1 400
	4 ⁻	1 227			1 200	
3	3-	1 450		1 530	1 420	

TABLE V

Energy levels of octupole bands in ²²⁸Th

one, which is compatible with the high limits set for $\log ft$ values as given in the decay scheme.

The allowed transitions to the gamma-vibrational band, as well as the 1st-forbidden non-unique transitions to the $K^{\pi} = 2^{-}$ and 3^{-} octupole bands, are not hindered by the K selection rule. To explain the low rate of transitions to the first two of these bands qualitatively, we may refer to the microscopic-model calculations performed by Zheleznova et al. [17]. They show for both collective wave functions a very low contribution of those two-quasiparticle configurations which can presumably participate in the EC transformation. The transition to the $KI^{\pi} = 33^{-}$ state at 1 450 keV is faster (log ft = 7.28). This fact seems to be in a strong disagreement with the results of calculations performed by Błocki [19] who has found that the 3⁻ state in question has an almost pure (98.5 %) two-proton configuration $(541 \downarrow + 642 \uparrow)$, which cannot be fed in the EC decay of ²²⁸Pa. Zheleznova et al. [17] do not give any explicit information on the structure of the 3^- states.

Table VI contains a list of those pure two-quasiparticle states, predicted by the superconductivity model, which have proper configurations from the point of view of their direct EC feeding from the ²²⁸Pa ground state, and which are theoretically predicted at energies below 2.4 MeV.

The transitions occurring in this case are either allowed hindered or 1 st-forbidden unhindered, for which we assume $\log ft$ values of 6.5 and 7.0, respectively. It should be realized that the experimental information about the beta transitions between the quasiparticle states of odd-A actinide nuclei is very scarce and, therefore, our estimate of the log ft values contains a large uncertainty.

The $(752\uparrow + 761\uparrow)$ configuration may possibly be ascribed to the 4⁺ level observed at 1 432 keV. However, identification of other two-quasiparticle states,

TABLE	11	
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Two-quasiparticle levels in ²²⁸Th fed by allowed and first forbidden EC decay of ²²⁸Pa

Two-quasiparticle states						
Configuration	Energy	Spin	log <i>ft</i>			
	(MeV)	and parity	(assumed)			
nn $(752 \uparrow + 631 \uparrow)$	1.31	4-	7.0			
nn $(752 \uparrow + 761 \uparrow)$	1.44	4+	6.5			
nn (752 ↑ ± 770 ↑)	1.71	2+, 3+	6.5			
nn (752 $\uparrow \pm 640 \uparrow$)	1.76	2-, 3-	7.0			
nn (752 $\uparrow \pm 501 \downarrow$)	1.97	2+, 3+	6.5			
nn $(752 \uparrow \pm 631 \downarrow)$	2.11	2-, 3-	7.0			
pp $(530 \uparrow + 651 \uparrow)$	1.12	2-	7.0			
pp $(530 \uparrow + 532 \downarrow)$	1.36	2+	6.5			
pp $(530 \uparrow + 642 \uparrow)$	1.40	3-	7.0			
pp $(530 \uparrow + 523 \downarrow)$	1.41	3+	6.5			
pp $(530 \uparrow + 521 \uparrow)$	2.00	2+	6.5			
pp (530 † + 633 †)	2.10	4-	7.0			

Energies calculated by Błocki and Kurcewicz (unpublished). The version of the superconductivity model used in these calculations has been earlier described in reference [19] and [20].

at higher energy, would be difficult not only because of the lack of complete experimental information on spins, parities, log *ft* values or deexcitation patterns, but also due to the expected level-mixing effects. It has been decided, therefore, to analyse the distribution of the average beta strength rather than probabilities of individual transitions. The energy range of the ²²⁸Th excitations defined by the value of $Q_{\rm EC}$ has been divided into $\Delta E = 0.3$ MeV intervals and for each interval the beta strength

$$S = \frac{1}{\Delta E} \sum \frac{1}{ft}$$

has been calculated. The results are shown in figure 13. Except for one energy interval, the experimental betastrength distribution is lower than the analogous distribution calculated on the basis of the table VI data. Some excess of the strength observed between 1.8 and 2.1 MeV could be perhaps an indication of the role of the four-quasiparticle configuration $n752\uparrow$, $n631\uparrow$, $p530\uparrow$, $p631\downarrow$. Such a state would be fed by a fast, allowed unhindered transition. Admixtures of this four-quasiparticle configuration to some of the even-parity states in the considered energy interval can explain the appearance of the enhancement of the beta decay to these states.

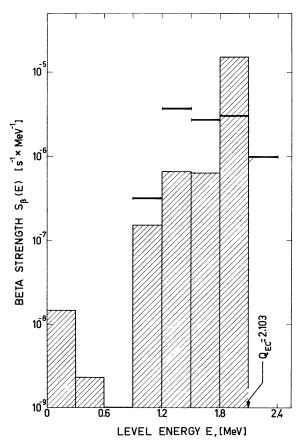


FIG. 13. — Experimental beta strength distribution for the 228 Pa $\rightarrow ^{228}$ Th decay. The bold line (1.2-2.4 MeV) shows the distribution calculated for two-quasiparticle states basing on the data of table VI.

8. Levels of ²²⁸Th populated in the decay of ²²⁸Ac. — Among the 39 levels of ²²⁸Th shown in the decay scheme of ²²⁸Pa (Fig. 10*a* and 10*b*), 27 were also found to be populated in the decay of ²²⁸Ac. The total number of the ²²⁸Th levels observed in the latter decay is 49 (Dalmasso and Maria [7], Herment and Vieu [8], Herment [9]). The few differences in the results obtained both from studies of the ²²⁸Pa and ²²⁸Ac decays, which are of importance from the point of view of the discussion developed in sections 7.1 and 7.2, are briefly commented below.

The 969 keV level. — In the ²²⁸Ac \rightarrow ²²⁸Th decay scheme, Herment [9] (see also ref. [7] and [8]) shows only one level at 969 keV and interprets it as the γ -vibrational state. There is no suggestion as to the existence of the negative-parity level at approximately the same energy. However, its existence may now be considered as certain, since the arguments given in section 7.2 (e. g. the M 1 character of the 640 keV transition) are now strengthened by the fact that Herment places the 640 keV transition between the 969 keV and 328 keV 3⁻ levels not only basing on the energy fit but also on coincidence studies.

The 1016 keV level. — From the M 1 + E 2 character of the 620 keV transition, placed on the basis of the coincidence experiment between the 1016 keV and 396 keV 3⁻ levels, the present authors deduce the negative parity for the 1016 keV state (section 7.2). On the other hand Herment, who places the 620 keV transition in the same way but having no evidence for its character, suggests that the 1154 keV (2)⁺ and 1026 keV levels may be connected by the 138 keV (M 1 + E 2) transition. This would indicate the positive parity of the 1016 keV state. Since there is no coincidence evidence for such a placing of the 138 keV transition, the assignment of the negative parity to the 1016 keV level seems more probable.

The 1 432 keV level. - The existence of the 308 keV (presumably E1) transition between the 1 432 keV and 1 123 keV 2⁻ levels, as shown in the $^{228}Ac \rightarrow ^{228}Th$ decay scheme by Herment, would mean that the present authors would have to abandon the $KI^{\pi} = 4,4^+$ assignment proposed in section 7.1 for the 1 432 keV state. The spin value could not be higher than 3, which is actually the assignment suggested by Herment. On the other hand it is difficult to learn from reference [9] how good is the evidence for the existence and placing of the 308 keV transition. This transition has not been observed in the present study of the ²²⁸Pa decay (the upper limit for the intensity of the 308 keV y-line is 30 units in the scale adopted for table I). Thus, the problem of the spin of the 1 432 keV level seems toremain open.

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