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### ► To cite this version:

Amy Williams, Kathleen Craft, Maëva Millan, Sarah Stewart Johnson, Christine Knudson, et al.. Fatty Acid Preservation in Modern and Relict Hot-Spring Deposits in Iceland, with Implications for Organics Detection on Mars. *Astrobiology*, 2021, 21 (1), pp.60-82. 10.1089/ast.2019.2115 . hal-03754903

**HAL Id: hal-03754903**

**<https://hal.science/hal-03754903>**

Submitted on 20 Aug 2022

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# Astrobiology

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## Fatty Acid Preservation in Modern and Relict Hot Spring Deposits in Iceland, with Implications for Organics Detection on Mars

Journal:	<i>Astrobiology</i>
Manuscript ID	AST-2019-2115.R3
Manuscript Type:	Special Collection: Hot Springs Van Kronendonk Ed
Date Submitted by the Author:	n/a
Complete List of Authors:	Williams, Amy; University of Florida, Department of Geological Sciences; NASA Goddard Space Flight Center, Planetary Environments Laboratory Craft, Kathleen; Johns Hopkins University Applied Physics Laboratory Millan, Maeva; Georgetown University, Department of Biology; NASA Goddard Space Flight Center, Planetary Environments Laboratory Johnson, Sarah; Georgetown University, Department of Biology; Georgetown University, Science, Technology, and International Affairs Program Knudson, Christine; NASA Goddard Space Flight Center, Planetary Environments Laboratory; University of Maryland at College Park College of Computer Mathematical and Natural Sciences, Center for Research and Exploration in Space Science and Technology Juarez Rivera, Marisol; Arizona State University McAdam, Amy; NASA Goddard Space Flight Center, Planetary Environments Laboratory Tobler, Dominique; University of Copenhagen, Department of Chemistry Skok, J.R.; SETI Institute
Keyword:	Hot Springs, Organic Geochemistry, Mars, Analog, Biosignatures
Manuscript Keywords (Search Terms):	hot springs, sinter, FAME, MSL, Sample Analysis at Mars (SAM) instrument, TMAH

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3 **1 Fatty Acid Preservation in Modern and Relict Hot Spring Deposits in Iceland, with**  
4 **2 Implications for Organics Detection on Mars**

6  
7 **3 *Running Title: Organics Preservation in Icelandic Hot Springs***

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49 22 Keywords: Hot springs, Sinter, Fatty acids, TMAH, SAM instrument

## 26 **Abstract**

27 Hydrothermal spring deposits host unique microbial ecosystems and have the capacity to preserve  
28 microbial communities as biosignatures within siliceous sinter layers. This quality makes  
29 terrestrial hot springs appealing natural laboratories to study the preservation of both organic and  
30 morphologic biosignatures. The discovery of hydrothermal deposits on Mars has called attention  
31 to these hot springs as Mars-analog environments, driving forward the study of biosignature  
32 preservation in these settings to help prepare future missions targeting the recovery of  
33 biosignatures from Martian hot spring deposits. This study quantifies the fatty acid load in three  
34 Icelandic hot spring deposits ranging from modern and inactive to relict. Samples were collected  
35 from both the surface and 2 to 18 cm in depth to approximate the drilling capabilities of current  
36 and upcoming Mars rovers. To determine the preservation potential of organics in siliceous sinter  
37 deposits, fatty acid analyses were performed with pyrolysis-GC-MS utilizing thermochemolysis  
38 with tetramethylammonium hydroxide (TMAH). This technique is available on both current and  
39 upcoming Mars rovers. Results reveal that fatty acids are often degraded in the subsurface relative  
40 to surface samples but are preserved and detectable with the TMAH pyrolysis-GC-MS method.  
41 Hot spring mid-to-distal aprons are often the only texturally and geomorphically definable feature  
42 in older, degraded terrestrial sinter systems and are therefore most readily detectable on Mars from  
43 orbital images. These findings have implications for the detection of organics in Martian  
44 hydrothermal systems as they suggest that organics might be detectable on Mars in relatively  
45 recent hot spring deposits, but preservation likely deteriorates over geologic time scales. Rovers  
46 with thermochemolysis pyrolysis-GC-MS instrumentation may be able to detect fatty acids in hot  
47 spring deposits if the organics are relatively young; therefore Martian landing site and sample  
48 selection are of paramount importance in the search for organics on Mars.

## 49 **1. Introduction**

50 Siliceous sinter are terrestrial geothermal spring deposits often formed in volcanic settings from  
51 silica-supersaturated fluids heated to high temperature by magmatic reservoirs. These fluids  
52 rapidly ascend to the surface where the water discharges at geysers and hot spring vents,  
53 precipitating silica as the pressure and temperature of the fluids rapidly decrease (Campbell et al.,  
54 2015; Renaut and Jones, 2011). The precipitating silica can nucleate and coat all surfaces in the  
55 spring discharge flow path, resulting in unique mineralogical textures that are dependent on

1  
2  
3 56 temperature and fluid composition when the silica precipitates. Microorganisms that thrive in these  
4  
5 57 extreme temperature, flow rate, and fluid compositions include bacteria, archaea, phages, and  
6  
7 58 eukaryotes (e.g. diatoms) (Breitbart et al., 2004; Leira et al., 2017; Pumas et al., 2018; Saini et al.,  
8  
9 59 2017; Skirnisdottir et al., 2000), and the cells of these organisms can also be entombed by silica  
10  
11 60 precipitation, preserving both the physical structure of the cell as well as cellular organic matter  
12  
13 61 (Cady and Farmer, 1996). These domains of life inhabit specific microenvironments at the hot  
14  
15 62 springs conducive to the organisms' ideal growth conditions. Hyperthermophiles and thermophiles  
16  
17 63 are found near the higher-temperature spring vent, mesophilic organisms such as cyanobacteria  
18  
19 64 inhabit fluids <73 °C between the vent, proximal slope, mid-apron, and distal-apron, and plants  
20  
21 65 and other mesophiles inhabit the lower temperature environments (commonly proximal slope,  
22  
23 66 mid-apron or distal-apron) (Figure 1).

24  
25 67 The preservation of organic biosignatures in geothermal spring deposits is of particular interest to  
26  
27 68 the astrobiology community, as several siliceous hot spring deposits have been identified on Mars  
28  
29 69 with both remote sensing and *in situ* exploration. Hydrothermally-driven siliceous sinter deposits  
30  
31 70 have been identified on Mars from orbital observations in Nili Patera (Skok et al., 2010) and Arabia  
32  
33 71 Terra (Allen and Oehler, 2008; Rossi et al., 2008). Hot spring opal-A deposits with nodular and  
34  
35 72 digitate stromatolitic morphologies have been identified in Columbia Hills, Gusev crater, by the  
36  
37 73 *Spirit* rover (Ruff et al., 2011; Squyres et al., 2008), which are strikingly similar to digitate  
38  
39 74 stromatolites with putative biogenic origins found at the El Tatio geothermal field in Chile (Ruff  
40  
41 75 and Farmer, 2016). Although not geothermal in nature, diagenetic silica enrichment is also  
42  
43 76 observed in Marias Pass, Gale crater, by the *Curiosity* rover, which implies groundwater  
44  
45 77 circulation in Gale crater long after lacustrine activity ceased (Frydenvang et al., 2017).

46  
47 78 The preservation of physical cells and cellular organic matter in hot springs both constitute  
48  
49 79 biosignatures. Biosignatures are defined as an object, pattern, or substance whose origin  
50  
51 80 specifically requires a biological agent (Hays et al., 2017). Extensive work has been performed  
52  
53 81 that documents the extraordinary preservation of microbial cell structure in sinter deposits by  
54  
55 82 entombment in a very fine matrix of silica particles (Brasier et al., 2011; Cady and Farmer, 1996;  
56  
57 83 Campbell et al., 2015; Guido et al., 2010; Jones et al., 2000; Konhauser et al., 2001; McCollom et  
58  
59 84 al., 2016; Munoz-Saez et al., 2016; Renaut et al., 1998; Tarhan et al., 2016; Walter, 1996; Walter  
60  
85 and Des Marais, 1993; Westall et al., 2015; Tobler et al., 2008). Few studies, however, have

86 focused on the preservation of organic matter in modern and relict (up to 900 year old) siliceous  
87 sinter deposits (Chaves Torres et al., 2019; Gibson et al., 2014; Kaur et al., 2015, 2011, 2008;  
88 Pancost et al., 2005; Pancost et al., 2006; Teece et al., 2019), and even fewer studies have explored  
89 organic preservation potential in Mars-analog hot spring environments (e.g. Williams et al., 2019).

### 90 **1.1. Martian *In Situ* Organics Detection**

91 The *in situ* search for organic molecules on Mars has been a priority of the scientific community  
92 for decades, and several missions have carried payloads with varying organics detection  
93 capabilities. The approaches that have been deployed and/or planned to detect organics on Mars  
94 missions include Raman spectroscopy (Hays et al., 2017 and references therein), deep UV laser  
95 excitation (Hays et al., 2015 and references therein), and gas-chromatography mass-spectrometry  
96 (GC-MS). GC-MS has the highest heritage, having flown in modified forms on the Mars Viking  
97 landers (Biemann et al., 1977), the Mars Phoenix polar lander (Boynton et al., 2001), and the  
98 Sample Analysis at Mars (SAM) instrument on the Mars *Curiosity* rover (Mahaffy et al., 2012). A  
99 GC-MS is also included as part of the Mars Organic Molecule Analyzer (MOMA) instrument on  
100 the upcoming European Space Agency's *Rosalind Franklin* (ExoMars) rover (Goesmann et al.,  
101 2017; Goetz et al., 2016).

102 The fundamentals of the GC-MS system for all Mars flight instruments are essentially the same:  
103 the goal of the approach is to detect and identify various molecular species and classes at low  
104 concentrations with high analytic specificity. Volatile compounds thermally evolved from solid  
105 samples in an oven may be directly heated (pyrolyzed) or subjected to a derivatization procedure.  
106 The evolved species are separated by the GC into a time sequence, then separated by mass and  
107 analyzed with the MS. Although pyrolysis GC-MS has the potential to liberate organics bound  
108 within mineral matrices, the detection of those organics can be compromised by the presence of  
109 perchlorate or other oxidizing salts. Perchlorate is present on the Martian surface and is known to  
110 combust organics during pyrolysis due to the evolution of O<sub>2</sub> from the perchlorate, effectively  
111 destroying the organic molecule in the process (Guzman et al., 2018). However, certain  
112 derivatization or thermochemolysis reactions can liberate and volatilize select organics in the  
113 presence of perchlorates without the organics combusting.

114 A derivatization or thermochemolysis reaction—referred to here, as in the literature, as a “wet  
115 chemistry experiment”—is a “one-step” reaction with a reagent that improves organic matter yield

1  
2  
3 116 and liberates organics bound within macromolecules. Wet chemistry experiments aim to transform  
4  
5 117 polar organic molecules (e.g. carboxylic acids) into volatile derivatives that are amenable and  
6  
7 118 detectable by GC-MS (del Rio et al., 1996; Metcalffe and Wang, 1981), as well as freeing bound  
8  
9 119 components of larger macromolecules otherwise undetectable with GC-MS (Grasset et al., 2002).

10  
11 120 Different wet chemistry reagents can be used based on the chemistry of the targeted labile group  
12  
13 121 and the required yield of the reaction, such as *N,N*-methyltert-butyl-  
14  
15 122 dimethylsilyltrifluoroacetamide (MTBSTFA; (Goesmann et al., 2017; Mahaffy et al., 2012), *N,N*-  
16  
17 123 dimethylformamide dimethyl acetal (DMF-DMA; Goesmann et al., 2017), trimethylsulfonium  
18  
19 124 hydroxide (TMSH; Ishida et al., 1999), and tetramethylammonium hydroxide (TMAH; Grasset et  
20  
21 125 al., 2002; Mahaffy et al., 2012), to name just a few. Thermochemolysis with TMAH in methanol  
22  
23 126 (MeOH, a solvent) allows access to refractory or insoluble organic material, including  
24  
25 127 macromolecular organics such as kerogen-like compounds. Simply pyrolyzing a sample at high  
26  
27 128 temperature (>650°C) results in fragmentation of the parent molecule due to thermal stress, which  
28  
29 129 limits molecule identification. To cleave these refractory molecules at a milder temperature (e.g.  
30  
31 130 ≤600°C), thermochemolysis with TMAH allows the analysis of refractory matter in complex  
32  
33 131 matrices containing low amounts of organic material with minimal destruction of the organic  
34  
35 132 material (Geffroy-Rodier et al., 2009; He et al., 2019; Williams et al., 2019). TMAH  
36  
37 133 thermochemolysis selectively cleaves ester and ether bonds, replacing a labile hydrogen atom with  
38  
39 134 a methyl group (Figure 2), rather than random and uncontrolled thermal decomposition and  
40  
41 135 fragmentation of the organic material (as during pyrolysis). TMAH has been used for several  
42  
43 136 decades in a variety of terrestrial fields of study, including petroleum geochemistry (Larter and  
44  
45 137 Horsfield, 1993), soil science (Chefetz et al., 2000; Deport et al., 2006; Schulten, 1996), and  
46  
47 138 sedimentology (Guignard et al., 2005; Pulchan et al., 1997). TMAH is currently available on the  
48  
49 139 *Curiosity* rover SAM instrument (Mahaffy et al., 2012; Williams et al., 2019) and will also be used  
50  
51 140 on the ExoMars rover MOMA instrument (Goesmann et al., 2017). Fortunately, perchlorate does  
52  
53 141 not substantially affect the recovery of fatty acids liberated with TMAH thermochemolysis (He et  
54  
55 142 al., *submitted*), making TMAH an ideal reagent to use in the search for select organics on Mars.

50 143 Although the SAM instrument TMAH experiment on the *Curiosity* rover has yet to be performed  
51  
52 144 (as of the time of this writing), several other lines of evidence already indicate that organics are  
53  
54 145 present on the surface and near-surface (<6 cm) of Mars. Chlorinated hydrocarbons such as  
55  
56 146 chlorobenzene and C<sub>2</sub> to C<sub>4</sub> dichloroalkanes were detected in the Sheepbed mudstone in

1  
2  
3 147 Yellowknife Bay, Gale crater (Freissinet et al., 2015), and thiophenic, aromatic, and aliphatic  
4  
5 148 organic compounds were found in the 3.5 Gy Murray Formation lacustrine mudstones of Pahrump  
6  
7 149 Hills, Gale crater (Eigenbrode et al., 2018). In more recent SAM experiments, preliminary analyses  
8  
9 150 suggest the presence of select medium-chain alkanes (Freissinet et al., 2019) and medium to high  
10  
11 151 molecular masses, including derivatized molecules (Millan et al., 2019), in samples exposed to  
12  
13 152 SAM's other wet chemistry experiment, derivatization with MTBSTFA (Buch et al., 2006).

14 153 These findings represent the first confirmed *in situ* organics detection on Mars, but the source of  
15  
16 154 these organic molecules is uncertain. Abiotic Martian sources (e.g., from atmospheric,  
17  
18 155 hydrothermal, or igneous processes), abiotic exogenous sources such as comets, interplanetary  
19  
20 156 dust particles, or meteorites (Eigenbrode et al., 2018; Freissinet et al., 2015), and life all produce  
21  
22 157 organic molecules. TMAH thermochemolysis, however, provides an unprecedented opportunity  
23  
24 158 to assess the biogenicity of Martian organics, specifically carboxylic and fatty acids. Fatty acids,  
25  
26 159 bound in phospholipids and glycolipids, are abundant and ubiquitous components of eukaryotic  
27  
28 160 and bacterial cellular membranes (Vestal and White, 1989). Carboxylic acids have also been  
29  
30 161 identified in exogenous carbonaceous material, such as in the Murchison meteorite (Cronin et al.,  
31  
32 162 1993). Biotic and abiotic-generated carboxylic acids produce distinctly different profiles in GC-  
33  
34 163 MS analyses. Carboxylic and fatty acids derived from metabolic processes can range from C<sub>2</sub> to  
35  
36 164 >C<sub>30</sub> in length have an even-over-odd carbon chain length preference due to enzymatically formed  
37  
38 165 acetyl (C<sub>2</sub>) units derived from glucose (Volkman, 2006). Abiotic fatty acid patterns will favor  
39  
40 166 shorter carbon chain lengths with no carbon preference (Bray and Evans, 1961; McCollom et al.,  
41  
42 167 1999), although it is important to note that very short chain carboxylic acids can represent  
43  
44 168 metabolic byproducts that were not incorporated into cellular membranes. Therefore, within a  
45  
46 169 certain level of uncertainty, characteristics of a FAME profile could be used to deconvolve the  
47  
48 170 origin of fatty acids detected on Mars (Williams et al., 2019).

49 171 The work presented here investigates the preservation and detection of organic molecules,  
50  
51 172 specifically fatty acids, in modern inactive and relict Icelandic hydrothermal spring systems that  
52  
53 173 precipitate siliceous sinter and travertine. Organics preservation in these Mars-analog hot spring  
54  
55 174 environments is compared between mostly inactive and relict spring deposits and include samples  
56  
57 175 from spring vent, mid-apron and distal apron facies, and within a depth range comparable to the  
58  
59 176 sampling depth of the *Curiosity* and future rover drill systems. We present data to assess the limits



177 of space-flight-like analyses and how returned *in situ* data from extraterrestrial hot springs may be  
178 interpreted.

## 179 2. Methods

### 180 2.1. Sample Locations

181 Samples were collected at three Icelandic hydrothermal sinter sites, at the surface and subsurface  
182 ranging from 2 to 18 cm in depth from spring vent to distal apron, to explore the coupling of system  
183 activity with potential preservation of biosignatures. In hydrothermal hot-spring systems there are  
184 three main regions: near-vent, the location of higher-temperature fluid discharge from the  
185 subsurface; the middle-apron, a moderate temperature region often with pooling water; and the  
186 distal-apron, areas of cooler fluid further from the vent (e.g. Figure 1). A currently active region,  
187 Hveravellir, sits in the central highlands area of Iceland (Figure 3B) and contains many hot springs  
188 and fumaroles. This area is protected as a monument in Iceland, with boardwalks and restricted  
189 sites that preserve the formations and minimize human alteration. At Hveravellir, we collected  
190 sample 160730.01 (mid-apron) near an active spring vent, but sampling site was only moist and  
191 not within the main flow region (Figure 4). We also sampled at a recently inactive vent (160730.09)  
192 and the distal-apron (160730.08) of an inactive vent. The pisolith sample (160730.04) had formed  
193 in pools at an actively flowing vent. Given the variety of samples from active and recently inactive  
194 sites, the Hveravellir system will be referred to as ‘inactive’ in this work. The geothermally active  
195 Hveravellir region has been geochemically characterized, with the spring vents in the area  
196 generally emerging at *ca.* 90 °C, with a pH of 8.60 and [SiO<sub>2</sub>] of 599 ppm (Camacho, 2017). The  
197 actively flowing spring vents proximal to Hveravellir vent sample “160730.09” ranged from 91.0  
198 to 97.1°C, pH 8.15 to 8.65, [SiO<sub>2</sub>] of 321 to 641 ppm, and [Cl] of 53.1 to 63.6 ppm. Hveravellir  
199 spring waters are considered part of a population of high-temperature geothermal waters, along  
200 with Geysir, Hveragerdi, and Reykjadalir, which contain elevated silica concentrations relative to  
201 the lower-temperature geothermal regions on the island.

202 A second site, Gunnuhver, lies on the Reykjanes peninsula in SW Iceland (Figure 3C). The  
203 sampling site is recently inactive (no spring activity), but with a large active fumarole ~10 m away  
204 and geothermal heat still warming the surface. A downstream (~50 m) channel from the active  
205 vent in the Gunnuhver region was 24.1°C with a pH = 7.34. Samples collected at Gunnuhver

206 included those near-vents, mid- (15 to 23 m) and distal-apron (33 to 49 m) locations from deduced  
207 vent locations (Figure 4).

208 The third site visited was Lýsuhóll, a farmland area with relict sinter deposits and few active vent  
209 systems (Figure 3D) with lower temperatures (*ca.* 41°C) and pH = 6.3. Other low temperature (*ca.*  
210 50 °C) thermal springs are present nearby in the Snæfellsnes region (Camacho, 2017). Silica sinter  
211 (e.g. Opal-A) deposits constitute the mineralogy of older regions in this deposit, overprinted by  
212 younger travertine deposits (Jones and Renaut, 2017, Jones et al., 2005). Vent and near-vent  
213 samples were collected at Lýsuhóll. The determination of vent locations was more difficult at this  
214 relict site as erosion had disrupted many of the deposits; therefore only three well preserved sites  
215 were sampled (Figure 4).

216 All samples were collected with the appropriate permissions. Hveravellir is a protected natural  
217 monument for which we were permitted to sample by the Environment Agency of Iceland. We  
218 were permitted to sample the Gunnuhver site by the National Energy Authority. Lýsuhóll is private  
219 land for which we had owner permission to sample. Lastly, we also had permission from the  
220 Icelandic Institute of Natural History to transport samples out of Iceland.

## 221 2.2. Sampling Procedure

222 To minimize contamination, all sinter samples were collected using solvent-washed and ashed  
223 tools, while wearing gloves, facemasks and hairnets. Samples were collected using solvent cleaned  
224 (MeOH – MeOH/DCM – DCM) and ashed (at 500°C) chisels, spoons, and tweezers. Rock  
225 hammers were covered in a new piece of ashed (at 500°C) foil before use. Samples were placed in  
226 solvent washed and ashed (at 500°C) glass jars and stored on ice for transport back to the  
227 laboratory. Samples were collected both from the surface (denoted with an ‘S’) or from several  
228 centimeters into the subsurface (denoted with an ‘I’). Fresh solvent-washed tools were used for  
229 each surface and subsurface sample to limit cross-contamination between sample depths. Samples  
230 were transported to the laboratory on ice and stored in a -20 °C freezer upon arrival. Samples for  
231 SEM analysis were either collected in the field in sterile whirlpak bags with solvent washed tools,  
232 or subsampled in the lab from the solvent washed and ashed glass jars with solvent washed tools.  
233 In the laboratory, sinter from the glass jars was subsampled by breaking off smaller pieces with  
234 solvent washed and ashed tweezers. These subsamples were powdered with a solvent washed and

235 ashed ceramic mortar and pestle and ground for 3 minutes to a homogenous powder. As samples  
236 were carefully collected using organically-clean tools and techniques in the field, samples were  
237 not treated to remove potential natural external contamination prior to powdering. This follows the  
238 approach of Wilhelm et al. (2017).

### 239 **2.3. Imaging and Mineralogy**

240 The aerial drone maps for Figure 3 was acquired by DJI Mavic Pro with the onboard camera at  
241 12MP resolution. Images were recorded at 2-second intervals to ensure that most surface points  
242 would be captured in >5 images. The images were stitched together using the  
243 commercially available Pix4D Mapping Software to combine the images into the mosaic. Due to  
244 the size of the field sites, map images were acquired at different flight altitudes resulting in the  
245 following pixel resolutions: Gunnuhver (5 m, 0.32 cm/pixel), Hveravellir (33 m, 1.41  
246 cm/pixel), and Lýsuhóll (60 m, 2.76 cm /pixel).

247 Scanning electron microscope (SEM) images of sinter were collected with a Phenom ProX G5  
248 tabletop environmental SEM. Qualitative analyses were conducted using a 3- $\mu\text{m}$  spot size, a beam  
249 current of 2.25–2.33 nA, and an accelerating voltage of 15 kV under low vacuum conditions  
250 consistent with environment SEM operation.

251 Mineralogy was determined via X-ray diffraction with either a Bruker D8 Discover X-Ray  
252 Diffractometer, (Cu K $\alpha$  radiation,  $\lambda = 1.54059 \text{ \AA}$ ); patterns were acquired from 2-70° 2 $\theta$ , or an  
253 Olympus Terra X-Ray Diffractometer, (Co K $\alpha$  radiation,  $\lambda = 1.78897 \text{ \AA}$ ); patterns were acquired  
254 from 5-55° 2 $\theta$ .

### 255 **2.4. Pyrolysis GC-MS Conditions for Analog Samples**

256 Aliquots of each powdered sample underwent pyrolysis (with a commercial Frontier pyrolyzer  
257 3030D) GC-MS analysis to measure the distribution and abundance of FAMEs. Ground rock or  
258 sediment samples were weighed into non-reactive metal cups (sample mass was ca. 5 to 10 mg,  
259 depending on the sample). Just prior to sample analysis, TMAH in the ratio of 1 mg sample to 1  
260  $\mu\text{L}$  TMAH was added to the cup. The *n*-C<sub>19:0</sub> internal standard was injected into each sample  
261 immediately prior to analysis to determine the efficiency of FAME transfer. The sample was then  
262 immediately loaded and dropped into the pyrolyzer oven to begin the run. The GC-MS analysis  
263 was run on either an Agilent 7890A-5975C inertXL GC-MS or a Thermo Trace 1310 GC-MS  
264 equipped with a 30 m Restek capillary column (MXT-5) with a 0.25 mm internal diameter and

1  
2  
3 265 0.25  $\mu\text{m}$  thick Crossbond® 5% diphenyl/95% dimethyl polysiloxane stationary phase, and He as  
4  
5 266 the carrier gas at a 3 mL  $\text{min}^{-1}$  flow rate and 10:1 split. Blank (without TMAH) and “TMAH-  
6  
7 267 blank” (with TMAH) clean up analyses were run between each sample to determine and subtract  
8  
9 268 any background level of residual FAMES in the column.

10  
11 269 Chromatograms and the mass spectra for FAMES were analyzed with ChemStation software  
12  
13 270 (Agilent Technologies). Identifications were based on comparison to known FAME retention  
14  
15 271 times of a Supelco 37 component FAME mixture that contains saturated FAMES, monounsaturated  
16  
17 272 fatty acid methyl esters (MUFAMES), and polyunsaturated fatty acid methyl esters (PUFAMES)  
18  
19 273 (Sigma-Aldrich). In addition, mass spectra were compared to the National Institute of Standards  
20  
21 274 and Technology (NIST) Spectral Library. Quantifications were based on integration under peaks  
22  
23 275 and scaled to the known amount of the n-C<sub>19:0</sub> internal standard.

24  
25 276 The pyrolyzer and GC programs used are detailed below: For the SAM-like pyrolysis ramp and  
26  
27 277 GC column method, the pyrolyzer program started at 50 °C and was ramped at 35 °C/min to 400  
28  
29 278 °C with no hold. The inlet program and Tenax ® TA trap started at 15 °C with a 2 minute hold.  
30  
31 279 The inlet was ramped at 68 °C/min to 150 °C and held for 2 minutes, then ramped at 80 °C/min to  
32  
33 280 305 °C and held for 3 minutes. The GC column program started at 35 °C and was held for 10.9  
34  
35 281 minutes, then ramped at 15 °C/min to 175 °C with no hold, then ramped at 10 °C/min to 305 °C  
36  
37 282 and held for 1 minute. The auxiliary transfer line was held at 135 °C.

38  
39 283 For the 500 °C flash pyrolysis method, the sample was dropped into a 500 °C pyrolyzer oven and  
40  
41 284 held for 1 minute. The inlet program and Tenax ® TA trap started at 15 °C with a 2 minute hold.  
42  
43 285 The inlet was ramped at 900 °C/min to 300 °C and held for 2 minutes. The GC column program  
44  
45 286 started at 35 °C and was held for 2.5 minutes, then ramped at 5 °C/min to 300 °C and held for 5  
46  
47 287 minutes. The auxiliary transfer line was held at 270 °C.

### 48 288 **3. Results**

#### 49 289 **3.1. Mineralogy**

50  
51 290 The mineralogy of the three sample locations was generally dominated by opal-A, opal-CT, and  
52  
53 291 amorphous silica (Table 1). The modern inactive hot spring vent at Hveravellir was dominantly  
54  
55 292 opal-A with lesser percentages of clinopyroxene, plagioclase, gypsum, and magnesite. Several  
56  
57 293 Hveravellir locations were composed entirely of amorphous silica. The modern inactive hot spring

1  
2  
3 294 at Gunnuhver was predominantly opal-CT, with the distal apron being dominantly amorphous  
4  
5 295 silica. This system also contained lesser percentages of halite and quartz. The older relict hot spring  
6  
7 296 at Lýsuhóll is composed almost exclusively of amorphous silica with some minor calcite, and one  
8  
9 297 sample site was composed entirely of calcite.

### 10 11 298 **3.2. FAMES Detected with 500 °C Instantaneous Flash Pyrolysis**

12  
13 299 Samples were initially analyzed using a 500 °C flash pyrolysis technique to enable  
14  
15 300 transesterification of fatty acids with TMAH thermochemolysis. Using this technique there was  
16  
17 301 relatively high diversity and abundance of FAMES in the Hveravellir spring site, a lower diversity  
18  
19 302 and abundance of FAMES at the inactive Gunnuhver spring site, and diminished diversity and  
20  
21 303 abundance of FAMES at the Lýsuhóll relict site (Table 2).

22  
23 304 At the Hveravellir site, samples were collected from the spring vent, mid-apron, and distal-apron,  
24  
25 305 as well as a pisolith-bearing layer near an active vent. FAMES detected ranged from  $n$ -C<sub>6:0</sub> to  $n$ -  
26  
27 306 C<sub>27:0</sub> and included a few monounsaturated fatty acids (MUFAs) and branched *iso*- and *antiso*-C<sub>15:0</sub>  
28  
29 307 (Figure 5, Table 2).  $n$ -C<sub>16:0</sub> and  $n$ -C<sub>18:0</sub> were the most abundant FAMES in all samples, and FAMES  
30  
31 308 longer than  $n$ -C<sub>18:0</sub> were identified at all sites except the vent. In general, the abundance and  
32  
33 309 diversity of FAMES was increased in the surface samples relative to the subsurface samples, which  
34  
35 310 ranged from 3 to 18 cm in depth. The distal apron sample demonstrated an unexpected general  
36  
37 311 trend of greater FAME concentrations in the interior samples than the surface samples.

38  
39 312 At the inactive Gunnuhver site, samples were collected from two spring vents, the mid-apron, and  
40  
41 313 the distal-apron. FAMES detected ranged from  $n$ -C<sub>4:0</sub> to  $n$ -C<sub>27:0</sub> and included few MUFAs (Figure  
42  
43 314 6, Table 2). No methyl-branched fatty acids were detected. Again  $n$ -C<sub>16:0</sub> and  $n$ -C<sub>18:0</sub> were the most  
44  
45 315 abundant FAMES in all samples. FAMES longer than  $n$ -C<sub>18:0</sub> were only identified at the mid- and  
46  
47 316 distal-apron sites. In general, the abundance and diversity of FAMES was increased in the surface  
48  
49 317 samples relative to the subsurface samples, which ranged from 5 to 7 cm in depth.

50  
51 318 At the relict Lýsuhóll site, much of the characteristic features of the proximal slope, mid- and  
52  
53 319 distal-aprons were too degraded to assess, so samples were collected from three (near-)vent sites  
54  
55 320 which could be identified in field observations. FAMES detected ranged from  $n$ -C<sub>6:0</sub> to  $n$ -C<sub>27:0</sub> and  
56  
57 321 included a few MUFAs and one polyunsaturated fatty acid (PUFA) (Figure 7, Table 2). Methyl-  
58  
59 322 branched *iso*- and *antiso*-C<sub>15:0</sub> were also detected in the surface and subsurface. Again  $n$ -C<sub>16:0</sub> and

1  
2  
3 323  $n$ -C<sub>18:0</sub> were the most abundant FAMES in most samples, although this is not the case for Vent #1  
4 (160809.03). FAMES longer than  $n$ -C<sub>18:0</sub> were identified at all sites. The abundance and diversity  
5 324 of FAMES was increased in the surface samples relative to the subsurface samples, which ranged  
6  
7 325 from 2 to 3 cm in depth.  
8  
9 326

### 327 3.3. FAMES Detected with SAM-Instrument-like Pyrolysis Ramp

328 Samples were analyzed using a 35°C/min pyrolysis ramp to approximate the pyrolysis ramp used  
329 by the SAM instrument on the *Curiosity* rover (Mahaffy et al., 2012; Williams et al., 2019). This  
330 technique was utilized to explore how the FAME detection would change if sinter-bearing hot  
331 springs were analyzed on Mars with a SAM comparable TMAH experiment. In general, with this  
332 technique there was lower diversity and abundance of FAMES in all three spring systems (Table  
333 3).

334 At the Hveravellir site, FAMES detected ranged from  $n$ -C<sub>6:0</sub> to  $n$ -C<sub>20:0</sub> and included two MUFAs.  
335  $n$ -C<sub>16:0</sub> and  $n$ -C<sub>18:0</sub> were the most abundant FAMES in all samples, and only one FAME longer  
336 than  $n$ -C<sub>18:0</sub> was identified at the surface mid-apron site. In general, the abundance and diversity  
337 of FAMES was equal to or increased in the surface samples relative to the subsurface samples.

338 At the inactive Gunnuhver site, FAMES detected ranged from  $n$ -C<sub>4:0</sub> to  $n$ -C<sub>27:0</sub> and included  
339 several MUFAs and PUFAs.  $n$ -C<sub>16:0</sub> and  $n$ -C<sub>18:0</sub> were generally the most abundant FAMES in  
340 the mid- and distal-apron samples, and ~~no~~ FAMES longer than  $n$ -C<sub>18:0</sub> were detected in all  
341 Gunnuhver samples. ~~No FAMES larger than  $n$ -C<sub>16:0</sub> were detected in the Vent 2 site.~~ The  
342 abundance and diversity of FAMES was uniformly higher in the surface vent samples relative to  
343 the subsurface samples, but lower than the subsurface samples from the mid- and distal apron  
344 environments, with no FAMES detected in the subsurface Vent 2 sample.

345 At the relict Lýsuhóll site, FAMES detected ranged from  $n$ -C<sub>4:0</sub> to  $n$ -C<sub>18:0</sub> and included three  
346 MUFAs and one PUFA.  $n$ -C<sub>16:0</sub> and  $n$ -C<sub>18:0</sub> were the most abundant FAMES in all samples, and no  
347 FAMES longer than  $n$ -C<sub>18:0</sub> were detected. The only site with both a surface and subsurface sample  
348 was the 160807.04 Vent #2 location, and the abundance and diversity of FAMES was higher in the  
349 surface sample relative to the subsurface sample.

#### 350 4. Discussion

351 The samples studied here comprise silica sinter and calcite hot spring deposits ranging from  
352 modern and partially inactive to older, relict spring systems. Thus, they represent a range in  
353 geochemistry, mineralogy, age, and diagenetic history, as well as spanning a diversity of  
354 subenvironments present in hot spring deposits: vents, mid-aprons, and distal aprons. Despite these  
355 differences, several trends in FAME abundance and diversity are shared between the hot spring  
356 locations.

357 A high abundance of  $n\text{-C}_{16:0}$  and  $n\text{-C}_{18:0}$  is observed in all of the surface and subsurface samples  
358 relative to other FAMES in all three locations, which is consistent with these two fatty acids being  
359 the most common in both bacteria and eukarya (Kaur et al., 2015; Johnson et al., 2019; Wilhelm  
360 et al., 2017; Williams et al., 2019). Elevated  $n\text{-C}_{16:0}$  and  $n\text{-C}_{18:0}$  FAME abundances relative to other  
361 FAMES are expected in younger sinters versus much older, inactive sinters (Kaur et al., 2011),  
362 reflecting more extant diverse microbial communities.

363 ~~Most-All~~ samples contain FAMES longer than  $n\text{-C}_{18:0}$ , which most often represent a terrigenous  
364 plant or algal origin (Volkman et al., 1989; Kaur et al., 2011; Wilhelm et al., 2017; Johnson et al.,  
365 2020), although these fatty acids may be derived from bacterial free, long chain fatty acids  
366 (Summons et al., 2013) or intact polar lipids. Based on the abundance of surface vegetation near  
367 the spring sites, the FAMES here are assumed to represent waxy longer chain FAMES that are often  
368 more stable than short chain FAMES and may be more resistant to diagenetic processes and  
369 preserved longer after entombment (Wilhelm et al., 2017). Long chain FAMES were identified in  
370 ~~nearly~~ all surface and subsurface samples. At Hveravellir, they are present in greater abundance at  
371 the mid- and distal apron sites than the vent site. At Gunnuhver, long chain FAMES are ~~also~~ present  
372 in the mid- and distal-apron sites, ~~and present at a lower abundance with a noticeable absence~~ in  
373 the vent sites. This increase in abundance is consistent with models of lower temperatures ( $<40^\circ\text{C}$ )  
374 downstream from vents, which support plant growth and the potential for their entombment in the  
375 sinter (Walter, 1976). At Lýsuhóll these FAMES are present in the relict vent sites both on the  
376 surface and in the subsurface. Much higher concentrations of long chain FAMES in the surface  
377 samples suggest a more modern source. Observation of the long chain FAMES at depth suggests  
378 that plant matter grew and was entrained in the sinter as flow decreased and the spring system  
379 cooled and became dormant. The lower concentrations in the subsurface samples indicate some

380 higher plants were also incorporated into the rock record, with some degradation of the FAMES  
381 due to burial processes.

382 A strong even-over-odd carbon number preference is consistent with the production of the fatty  
383 acids due to enzymatic effects (e.g. the monomer addition of a 2-C sub-unit, (Georgiou and  
384 Deamer, 2014; Summons et al., 2008) and of microbial cellular metabolic processes (Elias et al.,  
385 1997; Grimalt and Albaiges, 1987; Nishimura and Baker, 1986; Volkman, 2006). All surface and  
386 subsurface samples from the Hveravellir site exhibit this trend, as do all surface samples from  
387 Gunnuhver.

#### 388 4.1. Organics Load in the Hveravellir Hot Spring System

389 The surface FAME load in the Hveravellir spring system varied greatly between the vent, mid-  
390 apron, and distal apron facies. The abundance of FAMES detected with flash pyrolysis at the opal-  
391 A vent facies is an order of magnitude (>35000 ng total FAME/ mg TOC) greater than the FAME  
392 abundance at the amorphous silica distal apron (ca. 1875 ng total FAME/ mg TOC). The  
393 amorphous silica mid-apron FAME abundance was lower still, at ca. 600ng total FAME/ mg TOC.  
394 The elevated FAME abundance at the Hveravellir vent sample indicates that the change in water  
395 flow and/or temperature at this recently ceased vent has likely led to colonization by bacteria with  
396 a lower temperature tolerance. Water flow and temperature are major factors in determining  
397 microbial diversity and biofilm texture in hydrothermal systems. When flow shifts, as it has at this  
398 vent and outflow system, change to abundance and diversity of the fatty acids should be expected.  
399 At the vent, the presence of the C<sub>20</sub> and C<sub>22</sub> fatty acids indicate that thermophilic bacterial growth  
400 has occurred, however, the relatively high FA abundance, together with an increase in the  
401 cyanobacterial C<sub>18:09</sub> indicates that a reduction in water temperature probably resulted in increased  
402 growth of a thermotolerant cyanobacterium such as the high-temperature form of *Mastigocladus*  
403 isolated from an Iceland hot spring (Castenholz, 1969) or perhaps diatoms (e.g. Figure 4), or both.  
404 This is consistent with the relatively low abundance and lack of C<sub>18:09</sub> FA in the subsurface sample.  
405 As we report on and discuss our results for the sinter samples below it is important to keep in mind  
406 that these hydrothermal siliceous springs are dynamic systems suited for the production and  
407 preservation of fatty acid, and that under conditions ranging from recent to ancient, fatty acids  
408 survive for extended time entombed in siliceous sinter.



~~409 indicates that lower thermal tolerance bacteria have recolonized this recently ceased vent (i.e. no  
410 water flow). The expected trend for an active high temperature spring system would include a  
411 higher fatty acid abundance further from the high temperature vent (Kaur et al., 2011), as a greater  
412 diversity of organisms with higher growth rates tend to colonize these regions.~~

413 The opal-A pisolith layer was collected further away from the other three Hveravellir samples but  
414 exhibits a similar trend in even-over-odd carbon number predominance, with  $n\text{-C}_{16:0}$  and  $n\text{-C}_{18:0}$   
415 dominating the FAME profile. Pisoliths are composed of >2 mm diameter spherical concretionary  
416 grains called pisolites which are often formed from calcite or amorphous silica. Due to the  
417 concretionary nature of the grains, pisolites and related ooids often contain bound organic matter  
418 between cortices or submillimeter concretion layers (Davies et al., 1978; Folk and Lynch, 2001;  
419 O'Reilly et al., 2017). Therefore, the trends in presence and preservation of FAMES in the pisolith  
420 layer is expected.

421 The FAME load at Hveravellir changed with depth between the vent, mid-apron, and distal apron  
422 facies as well. Most subsurface FAMES decreased in abundance or were not detected relative to  
423 their complementary surface samples (e.g. Figure 8, where the total organic carbon load is  
424 generally lower in subsurface samples relative to surface samples). A few exceptions exist in this  
425 data set, in particular in the mid-apron sample, in which all of the subsurface FAMES were elevated  
426 in concentration relative to the surface (Figure 5). In the distal-apron samples, several individual  
427 subsurface FAMES were elevated relative to their complementary surface samples. This trend was  
428 distinct at the distal apron site, where FAME concentration often increased stepwise from the  
429 surface to progressively deeper sampling depths (e.g. Figure 5). We interpret this trend to represent  
430 either a temporary increase in organic preservation potential at spring sites that experience lower  
431 temperatures and diagenetic alteration, or a modern and robust subsurface microbial community  
432 that is generating the elevated FAME abundance. This trend could also be an indication of  
433 preservational bias related to differences in mineralization rates, where early mineralization  
434 minimizes the degradation of organics and body fossils (Alleon et al., 2016). If a diversion of water  
435 flow precipitated a localized temperature shift, the supply of dissolved silica would also decrease  
436 leaving the microbial community more vulnerable to biotic and abiotic degradation processes since  
437 it was not mineralizing. These observations reinforce the complexity of these spring systems.  
438 While general trends can be preserved, changes in the local physical conditions, (like diverted

1  
2  
3 439 water flow) can have cascading impacts on the microbial composition and their ecology, that  
4  
5 440 together end up biasing their fossil record.  
6

7 441 Although our sampling techniques were optimized to decrease any chance of anthropogenic  
8  
9 442 contamination, we cannot rule out the possibility that some of these samples reflect an  
10  
11 443 anthropogenic influence, as this hot spring is a tourist destination and the mid-apron sample was  
12  
13 444 collected below a visitor boardwalk. Elevated FAMES from eukaryotic (e.g. human) sources may  
14  
15 445 have been entrained in the sinter. Contamination of this type should yield a much higher FAME  
16  
17 446 concentration, especially in the C<sub>16:0</sub> and C<sub>18:0</sub> FAMES, in the surface relative to the subsurface  
18  
19 447 samples. Surface samples did contain higher concentrations of FAMES than the subsurface, but  
20  
21 448 C<sub>16:0</sub> and C<sub>18:0</sub> FAMES were not uniformly elevated. Therefore, we interpret the Hveravellir FAME  
22  
23 449 concentration trends as reflecting a diverse and active microbial community.

24 450 In general, branched and unsaturated fatty acids were either not detected in subsurface samples or  
25  
26 451 were decreased in abundance relative to the surface samples. This degradation was likely as the  
27  
28 452 unsaturation in the fatty acid is a more reactive location for the molecule to be cleaved by microbial  
29  
30 453 heterotrophy or diagenesis (Killops and Killops, 2005; Sun et al., 1997). A decrease in unsaturated  
31  
32 454 fatty acids and branching is documented in older sinter samples (Kaur et al., 2011) and is consistent  
33  
34 455 with the expectation for preferential degradation of unsaturated fatty acids. However, this trend  
35  
36 456 was reverse in the mid-apron sample, in which all of the subsurface FAMES, including branched  
37  
38 457 and MUFAs were elevated in concentration relative to the surface.

#### 38 458 **4.2. Organics Load in the Gunnuhver Hot Spring System**

39  
40 459 The surface FAME load in the inactive Gunnuhver spring system also varied between the vent,  
41  
42 460 mid-apron, and distal apron facies, although to a lesser degree than in the Hveravellir spring  
43  
44 461 system. The abundance of FAMES detected with flash pyrolysis is comparable between the opal-  
45  
46 462 CT vent and opal-CT mid-apron facies (ca. 18,000 to 20,000 ng total FAME/ mg TOC) and is  
47  
48 463 elevated in the amorphous silica distal apron facies (ca. 37000 ng total FAME/ mg TOC). This  
49  
50 464 trend in abundance is expected, as a greater diversity of organisms with higher growth rates tend  
51  
52 465 to colonize lower temperature geothermal regions such as the distal apron.

53 466 The abundance and diversity of lower molecular weight FAMES increases to the mid-apron and  
54  
55 467 distal apron facies relative to the vent facies, suggesting that higher molecular weight (longer)

1  
2  
3 468 FAMES are not less well preserved further from the vent and are likely degraded to shorter chains,  
4  
5 469 potentially due to microbial heterotrophy.

6  
7 470 The FAME load in this system changed with depth more dramatically than in the Hveravellir  
8  
9 471 system. Subsurface FAME abundance decreases substantially or to no detection relative to their  
10  
11 472 complementary surface samples, indicating continued degradation of the FAMES in the subsurface  
12  
13 473 due to diagenesis and microbial recycling. Only four FAMES (*n*-C<sub>8:0</sub>, *n*-C<sub>10:0</sub>, *n*-C<sub>15:0</sub> and *trans*-  
14  
15 474 C<sub>18:1</sub>) are detected in the subsurface distal apron site with flash pyrolysis(~~*n*-C<sub>8:0</sub>, *n*-C<sub>10:0</sub>, *n*-C<sub>15:0</sub>~~  
16  
17 475 ~~and *trans*-C<sub>18:1</sub>~~). The subsurface FAME profiles exhibit strong to weak even-over-odd trends in  
18  
19 476 carbon number. The subsurface distal apron sample does not exhibit any trends in carbon number  
20  
21 477 or molecular weight, which likely reflect whole-scale degradation of all FAMES in the subsurface.  
22  
23 478 Mono- and polyunsaturated fatty acids are already limited in detection at the surface, a dearth that  
24  
25 479 continues into the subsurface and indicates preferential degradation of the unsaturated FAMES.  
26  
27 480 Branched fatty acids are not detected in ~~neither the surface nor~~ subsurface samples with flash  
28  
29 481 pyrolysis.

### 4.3. Organics Load in the Lýsuhóll Hot Spring System

30  
31 483 In contrast with the opal and amorphous silica-dominated Hveravellir and Gunnhver sites, the  
32  
33 484 Lýsuhóll relict spring system is composed of both amorphous silica and calcite, with Vent #3  
34  
35 485 composed entirely of calcite. The spring waters that precipitate these deposits, and the subsequent  
36  
37 486 microbial communities that colonize the springs, may be different from those present in the  
38  
39 487 siliceous sinter deposits (Zhang et al., 2004). Additionally, several samples from the relict Lýsuhóll  
40  
41 488 system were destroyed and/or contaminated during transport back to the United States, reducing  
42  
43 489 the completeness of the surface versus subsurface samples originally collected at these sites.  
44  
45 490 Regardless, several conclusions can be drawn from the remaining samples about the organics load  
46  
47 491 in this relict system. All sample sites at Lýsuhóll represented near-vent environments as these were  
48  
49 492 the most readily identified in older, degraded hot spring deposits. The surface FAME load in the  
50  
51 493 relict Lýsuhóll system also varied between the vent facies, ranging from *ca.* 3,700 ng total FAME/  
52  
53 494 mg TOC in the amorphous silica and calcite dominated Vent #2, to *ca.* 350 ng total FAME/ mg  
54  
55 495 TOC in the amorphous silica Vent #1, to *ca.* 83 ng total FAME/ mg TOC in the calcite Vent #3  
56  
57 496 (Table 2). As in the other spring systems, the surface FAME concentrations are higher than in  
58  
59 497 subsurface samples. The total FAME abundance ranges from *ca.* 3700 ng total FAME/ mg TOC

1  
2  
3 498 in the Vent 160807.04 (Vent #2) surface sample to ca. 1500 ng total FAME / mg TOC in the  
4  
5 499 complementary subsurface sample. This indicates continued degradation of the FAMES in the  
6  
7 500 subsurface due to diagenesis and microbial recycling. The Lýsuhóll samples exhibit a moderate to  
8  
9 501 weak even-over-odd preference, which may indicate degradation of fatty acids from the original  
10  
11 502 microbial community, and limited input of organics from a modern community. The relict  
12  
13 503 Lýsuhóll system is the only site where the PUFA *trans*-C<sub>18:2</sub> algal biomarker is detected (Campbell  
14  
15 504 et al., 2015), and then only in two subsurface samples. Branched and unsaturated fatty acids are  
16  
17 505 already limited in detection at the surface, a dearth that continues into the subsurface and indicates  
18  
19 506 preferential degradation of these fatty acids.

#### 20 507 4.4. Comparison between Flash Pyrolysis and Flight-like Ramp Pyrolysis

21  
22 508 The 500 °C flash pyrolysis method, combined with TMAH thermochemolysis, represents a  
23  
24 509 laboratory-optimum method to release organics bound within macromolecules and sorbed to  
25  
26 510 minerals. The SAM instrument on the *Curiosity* rover, while able to perform similar experiments,  
27  
28 511 is limited to a 35°C/ min pyrolysis ramp. To assess how this method variation might affect the  
29  
30 512 detection of FAMES in these Mars-analog siliceous sinters, samples were analyzed with the SAM-  
31  
32 513 like 35°C/ min pyrolysis ramp.

33 514 The flash pyrolysis method improved FAME detection uniformly across the sample suite (Figure  
34  
35 515 8), and with ~~few one~~ exceptions, FAMES >*n*-C<sub>18:0</sub> were only detected with the flash method. The  
36  
37 516 ramp method released between 30 and 86% of the FAMES detected in the flash method. ~~Two~~  
38  
39 517 ~~outliers to this data set include sample 160727.03.I (200% improvement in the ramp versus flash~~  
40  
41 518 ~~method) and 160727.02.I (no fatty acids liberated using the ramp method).~~ Although the flash  
42  
43 519 method is more efficient at FAME release, it is promising for the search for organics on Mars that  
44  
45 520 FAMES are also detectable with the SAM instrument pyrolysis ramp rate.

46 521 Correlations vary between the number of FAMES detected, the concentration of FAMES in the  
47  
48 522 sample suite. In the Hveravellir system, the greatest diversity of FAMES detected was generally  
49  
50 523 correlated with low % TOC. This diversity is invaluable in assessing microbial community  
51  
52 524 composition, even at low concentrations. In the inactive spring system, the greatest diversity of  
53  
54 525 FAMES correlates with higher % TOC, which represents an anticipated trend with regards to  
55  
56 526 FAME preservation (and degradation), FAME diversity and concentration, and the total organic  
57  
58 527 carbon load in the sample (Table 1). The greatest diversity of FAMES was detected in the relict

1  
2  
3 528 spring system, as were some of the highest % TOC values. In particular, *n*-C<sub>16:0</sub> concentrations  
4  
5 529 were predictably low in the subsurface samples and high in the surface samples. These lines of  
6  
7 530 evidence indicate the relict system is experiencing higher delivery of organic load now than it  
8  
9 531 perhaps did in the past. The organics that constitute these higher % TOC values may represent a  
10  
11 532 far greater diversity of organic molecules than just fatty acids, which are preserved but in low  
12  
13 533 concentrations in the relict system.

#### 14 534 **4.5. Mineralogic Evolution and the Preservation of Organic Matter**

15  
16 535 It is important to note that the primary controls on lipid preservation in siliceous sinter deposits  
17  
18 536 are thermal alteration, high temperatures, and the pH of each spring (Kaur et al., 2015; Pancost et  
19  
20 537 al., 2006), which can vary greatly between spring systems. A great variety of lipid biomarkers have  
21  
22 538 been identified in modern to *ca.* 900 year old geothermal environments (Jahnke et al., 2001; Shiea  
23  
24 539 et al., 1991; Pancost et al., 2005, 2006; Zhang et al., 2007, Kaur et al., 2011). Often branched or  
25  
26 540 unsaturated fatty acids are prevalent in younger deposits (Kaur et al., 2011), as was also  
27  
28 541 demonstrated in the work presented here. In an actively flowing spring system, temperature in  
29  
30 542 particular tends to ~~drive~~ strongly influence the length and therefore abundance of ~~fatty~~  
31  
32 543 ~~acid~~ hydrocarbons, such as bacterial diethers in active geothermal systems (Kaur et al., 2015).  
33  
34 544 Thermophilic organisms tend to have ~~longer~~ higher molecular weight diethers with higher  
35  
36 545 thermostabilities, longer fatty acid chains (Oshima and Miyagawa, 1974; Russell, 1984;  
37  
38 546 Weerkamp and Heinen, 1972), increased ~~the ratios~~ of *iso*- to *anteiso*-branched fatty acids (Oshima  
39  
40 547 and Miyagawa, 1974; Shen et al., 1970), and fewer unsaturations and branches (Daron, 1970; Ray  
41  
42 548 et al., 1971). All of these adaptations serve to maintain optimal membrane fluidity at higher  
43  
44 549 temperatures.

45  
46 550 Although the Hveravellir vent sampled here was no longer actively flowing, several of these trends  
47  
48 551 were observed (Figure 9), consistent with the recent cessation of flowing spring water. The average  
49  
50 552 chain length ranges from 15.1 to 16.4 for all surface and subsurface Hveravellir samples, with the  
51  
52 553 exception of the surface vent sample with an ACL of 13.6. These values suggest that the surface  
53  
54 554 vent microbial community and/or FAME preservation is different from the other surface and  
55  
56 555 subsurface samples at Hveravellir. The low ACL value of the surface vent suggests that this spring  
57  
58 556 vent never experienced very high fluid temperatures and even if thermophilic organisms were  
59  
60 557 present, their organic biosignatures are not well preserved. An increase in heterotrophy may have

1  
2  
3 558 occurred concurrent with the cessation of water activity that hosted the thermophiles. The ratio of  
4 559 *i-* to *a-C*<sub>15</sub> is elevated near the vent where elevated temperatures affect community composition  
5  
6 560 and the *iso* to *anteiso* ratio trend was consistent at this site between the surface and subsurface  
7  
8 561 samples. Additionally, the ratio of branched to total fatty acids increases with distance from the  
9  
10 562 higher temperature vent, consistent with a decrease in the dominance of any thermophilic  
11  
12 563 organisms.

13  
14 564 In contrast, the inactive Gunnuhver site had a relatively consistent, elevated total FAME  
15  
16 565 concentration at the vent and mid-apron facies, and a much higher FAME concentration at the  
17  
18 566 distal apron facies, consistent with secondary recolonization of the sinter by more diverse, lower  
19  
20 567 temperature mesophiles after spring flow ceased. Average chain length also decreases from the  
21  
22 568 vent facies (ACL = 10.7 to 13.4 at vents) to the distal apron facies (ACL = 0 to 8.9), suggesting  
23  
24 569 the degradation of fatty acids to shorter chains due to heterotrophy or diagenesis.

25  
26 570 Only relict inactive vents were sampled at the Lýsuhóll spring site, which allows a comparison  
27  
28 571 with the modern inactive vent systems. The total FAME concentrations at the relict site was within  
29  
30 572 the range of the inactive spring sites, with higher abundances in the surface samples and lower  
31  
32 573 abundances in the subsurface samples. Average chain length at the relict site (ACL = 14.9 to 17.4)  
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34 574 is within the range of the Hveravellir spring site, suggesting that these Lýsuhóll vents were higher  
35  
36 575 temperature when flowing and the thermophile FAMES are well preserved. The *iso* to *anteiso* ratio  
37  
38 576 was lower than at the modern inactive spring sites, suggesting degradation of the branched fatty  
39  
40 577 acids from an original, higher temperature population. The ratio of branched to total fatty acids,  
41  
42 578 however, was comparable to or higher than the Hveravellir vent site, indicating that more branched  
43  
44 579 fatty acids relative to the total fatty acid load are preserved. This trend may be due to a greater  
45  
46 580 initial branched fatty acid load that, along with the total fatty acids, has experienced degradation.  
47  
48 581 The branched to total fatty acid ratio at Vent #3 (160807.12) is higher than all other samples and  
49  
50 582 contains the longest ACL of all samples from this study. This vent is composed solely of travertine,  
51  
52 583 which may introduce a preservation bias between this younger calcite deposit and the other vents  
53  
54 584 that are older and silica-dominated. The high branched to total fatty acid ratio and ACL value of  
55  
56 585 Vent #3 may also correspond to an active mesophilic microbial community at this site.

57  
58 586 Low temperature degradative processes such as heterotrophy and diagenesis control the fatty acid  
59  
60 587 distribution observed in this study. The primary fatty acid distribution from thermophilic versus  
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589 mesophilic communities is discernable in some environments (e.g. Hveravellir vent) but not

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3 589 diagnostic of more specific temperature and pH conditions of each subenvironment at each spring  
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5 590 system. The vents from the three systems studied here are compared to assess the preservation of  
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7 591 organics as a concentration of FAMES in the surface and subsurface (Figure 10). Additional work  
8  
9 592 can be completed to assess the preserved lipids separate from modern lipids in all three systems,  
10  
11 593 including solvent washing sinter samples to remove surface lipids prior to analysis. At the recently  
12  
13 594 inactive Hveravellir vent (sample 09), the C<sub>16</sub> and C<sub>18</sub> FAMES are likely elevated due to these being  
14  
15 595 the dominant FAMES in both bacteria and eukarya, although anthropogenic influences, especially  
16  
17 596 in the surface sample, cannot be completely dismissed. The lower concentrations of the other  
18  
19 597 FAMES in the surface and subsurface reflect a nominal FAME profile from an active microbial  
20  
21 598 community, with the expected even-over-odd carbon number preference. The inactive Gunnuhver  
22  
23 599 vent generally had the lowest FAME concentrations relative to the Hveravellir and Lýsuhóll  
24  
25 600 systems, but still exhibited the even-over-odd carbon number preference characteristic of an active  
26  
27 601 microbial community. The low FAME concentrations reflects FAME degradation leading to low  
28  
29 602 preservation potential, and a lack of extensive recolonization of the sinter by secondary microbial  
30  
31 603 communities. The higher FAME concentrations at the relict Lýsuhóll Vent #2 (especially relative  
32  
33 604 to the inactive Gunnuhver site) likely reflect microbial (re)colonization by a community distinct  
34  
35 605 from that found in active/recent silica-precipitating hot springs. The mixed amorphous silica and  
36  
37 606 calcite mineralogy of this vent likely also contributes to this marked difference in FAME  
38  
39 607 abundance.

40  
41 608 The mineralogic evolution and diagenesis of siliceous hot spring sinters is well studied (Herdianita  
42  
43 609 et al., 2000). Initial silica precipitates are composed of amorphous silica or opal-A (Tobler et al.,  
44  
45 610 2017), and the very fine size of these precipitates are documented to preserve microbial cells in  
46  
47 611 extraordinary physical detail (Cady and Farmer, 1996; Jones et al., 2000; Renaut et al., 1998),  
48  
49 612 although progressive silicification and preservation biases over long time scales can also destroy  
50  
51 613 these physical biosignatures (Jones et al., 1997; Jones and Renaut, 1996). As siliceous sinter  
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53 614 deposits experience diagenesis, opal-A incrementally recrystallizes to Opal-C (containing  
54  
55 615 disordered  $\alpha$ -cristobalite) or opal-CT (containing cristobalite and tridymite). This transition may  
56  
57 616 occur on the order of *ca.* 50 years up to tens of thousands of years. Over longer time scales (*ca.*  
58  
59 617 10,000 to 50,000 years), opal-C/ -CT will reorder to microcrystalline quartz. Although sinter age  
60  
618 can sometimes be used as an estimate for crystallinity, variables such as fluid mobilization and  
619  
620 composition can also greatly affect crystallinity (Herdianita et al., 2000). This recrystallization

620 process, and specifically the transition to microquartz, reduces porosity and decreases the  
621 preservation of organic matter in siliceous sinter (Teece et al., 2020).

622 These trends in mineralogy and organics preservation were upheld in the current study. In the  
623 Hveravellir system, which is composed of amorphous silica and opal-A, the diversity and  
624 abundance of FAMEs was elevated in both the surface and subsurface. In contrast, the inactive  
625 Gunnuhver system, which is composed primarily of opal-CT, had a much lower diversity and  
626 abundance of preserved FAMEs. Lynne et al. (2006) suggest that silica diagenesis is accelerated  
627 in hot spring systems that have transformed from flowing water to steam dominated. The  
628 transformation to opal-CT in this relatively young system was likely controlled by fluid  
629 composition and fluid or steam mobilization. This mobilization led to accelerated diagenesis and  
630 fumarolic degradation of organics in the Gunnuhver system. The surprisingly elevated diversity  
631 and abundance of FAMEs in the relict Lýsuhóll system was controlled in part by the mineralogy,  
632 which is dominated by calcite and amorphous silica, as well as a modern microbial community  
633 that had colonized the surface of this deposit. These results further reinforce the importance of  
634 ascertaining sample mineralogy prior to sampling for organics in hydrothermal systems, especially  
635 on life detection missions beyond Earth.

#### 636 **4.6. Biomarker Preservation and Implications for Organics Detection on Mars**

637 The size and distribution of hot spring facies varies across systems, yet trends in preservation  
638 styles, from the outcrop- to the submicrometer scale, are present and can be used to provide  
639 depositional context as we continue to explore life in other planets. On Earth, preservation at the  
640 vent is often less favorable with microfossils usually in the form of smaller diameter molds that  
641 can be readily infilled and obscured (Cady and Farmer, 1996). Vents may be identified by finely  
642 laminated deposits in a circular form that are often associated with geyserite. The directed water  
643 flow in channel facies produces bacterial streamer fabrics, long filaments parallel to water  
644 direction, sometimes composed of cyanobacterial filaments (Parenteau et al., 2014). This streamer  
645 morphology is readily identified in outcrop (e.g., Parenteau et al, 2014, Walter et al., 1998). The  
646 lower temperatures of the distal apron are usually accompanied by thick microbial mats creating  
647 palisade fabrics with abundant microfossils and organics that can be preserved (Cady and Farmer,  
648 1996; Campbell et al., 2015). Apron facies are one of the better-preserved lithofacies in degraded  
649 spring systems, with the mid-to-distal apron composing the largest aerial extent, followed by the



650 channel facies and then vent facies. Based solely on aerial extent, Martian hot spring apron facies  
651 may be most easily detected from orbital imaging and targeted for *in situ* exploration.

652 Biosignature preservation within all facies is complicated **by** preservation bias and diagenesis over  
653 geologic time scales. Even the proximal slopes and aerially extensive mid- and distal-aprons may  
654 become challenging to identify as the fine scale micro-textures that define these spring sub-  
655 environments can be lost to diagenesis and preservation biases (Cady and Farmer, 1996). Although  
656 diagenesis can and does obscure some features, the apron facies contain structures such as pools  
657 and terracettes that are identifiable in surface deposits and layered deposits with palisade fabric  
658 identifiable in outcrop cross-section. Vents are just as prone to erosion and infilling (this was  
659 observed at Lýsuhóll and Gunnuhver), and with a much smaller footprint, can be challenging to  
660 identify from orbital images. However, the primary macro-scale textural characteristics of spring  
661 vents can enable their identification after some geologic reworking (Cady and Farmer, 1996, and  
662 references therein).

663 Given these complexities, it is therefore important to develop a facies-oriented understanding of  
664 biosignature preservation potential in terrestrial hot springs to better predict the ideal locations to  
665 search for organics in Martian relict hydrothermal spring systems. The temperature of fluids at  
666 subaerial hydrothermal vent systems can vary greatly, up to or above surface boiling temperatures.  
667 Non-photosynthetic microbes can live in these temperatures, with hyperthermophiles preferring  
668  $>80$  °C up to a maximum of 110 °C where elevated hydrostatic pressure allows waters to reach  
669 higher temperature without boiling (Seeger et al., 1993). For reference, the upper temperature  
670 limit for archaeal life is 121 °C in submarine black smoker environments (Kashefi and Lovley,  
671 2003); in contrast, photosynthetic microbes are present in waters below 73 °C (Ward et al., 2012).  
672 Although these organisms do not directly contribute to silica precipitation in these spring systems,  
673 they can be entombed and preserved in the rapidly precipitating silica, contributing to the organic  
674 load and microstructures found in near-vent environments (Cady and Farmer, 1996). In our study,  
675 silicified traces of microorganisms are readily identified in SEM images from the Hveravellir and  
676 Lýsuhóll sample sites, but no distinct biogenic features were identified from the Gunnuhver sinter  
677 samples (Figure 4). These observations appear to correlate with the relative abundances of organics  
678 detected at these three locations.

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3 679 The search for (biogenic) organic molecules on Mars' near-surface requires new and novel  
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5 680 techniques to detect these molecules with miniaturized space-flight payloads. In the terrestrial  
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7 681 realm, many biomarkers are preserved and detectable in siliceous sinter environments with  
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9 682 laboratory instrumentation such as off-line lipid fraction extraction, derivatization and/or  
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11 683 saponification, followed by multiple GC-MS and/or liquid chromatography-mass spectrometry  
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13 684 (LC-MS) analyses. However, suites of these laboratory techniques are often too complex to be  
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15 685 successfully miniaturized for space-flight and remote operation. Therefore, "one-step"  
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17 686 thermochemolysis or derivatization pyrolysis GC-MS is an optimized technique that is  
18  
19 687 miniaturized to work on space-flight missions such as with the SAM and MOMA instruments. To  
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21 688 be more "SAM-like" in this study, samples that may contain low-temperature surface biofilms  
22  
23 689 were collected and ground as whole rock samples in a manner similar to *Curiosity's* sample  
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25 690 acquisition system. Therefore, the results from this study likely reflect a greater abundance of  
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27 691 organics than may be expected to be preserved in a Martian hot spring deposit with a hypothetical  
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29 692 microbial community. Although more diagnostic biomolecules are present in Mars-analog  
30  
31 693 hydrothermal spring environments, this work demonstrates that FAMES are preserved and  
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33 694 detectable with space-flight-like GC-MS techniques in modern inactive and relict siliceous sinter  
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35 695 and calcite deposits from Iceland. The MOMA instrument on the *Rosalind Franklin* rover will be  
36  
37 696 able to ramp the pyrolysis oven at >200 °C/min (from thermal vacuum), achieving a much faster  
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39 697 ramp than the SAM instrument and perhaps increasing the likelihood of detecting Martian FAMES,  
40  
41 698 if present.

37  
38 699 The rate of siliceous sinter formation, hydrothermal conditions, and those effects on biomass  
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40 700 growth and potential for silicification in each individual Icelandic sinter deposit reflect differences  
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42 701 in the diversity and abundance of biomolecules detected in this work. The geologic process of a  
43  
44 702 hot spring becoming quiescent should be similar on Earth and Mars. The evolution of a putative  
45  
46 703 microbial community within that hot spring can be approximated in specific ways by assessing  
47  
48 704 microbial community evolution in hot springs on Earth. The approximations ~~by definition~~  
49  
50 705 ~~require~~ require assumptions that do not account for Mars-specific conditions such as atmospheric  
51  
52 706 composition, radiation flux, energy sources, etc. By assessing the microbial communities and  
53  
54 707 preserved organics from lower temperature inactive and relict spring systems, this work  
55  
56 708 approximates the organics load in Martian hot springs with a hypothetical microbial community  
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58 709 that would have evolved as the spring system cooled to dormancy. Understanding the distribution

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3 710 of organics in terrestrial systems is vital for choosing an ideal Martian hydrothermal target for  
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5 711 organic biosignature exploration.

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7 712 A final consideration in the search for organics on Mars is the deleterious effects of oxidants and  
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9 713 radiation on organics. Although organic matter has been identified in Mars near-surface substrates  
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11 714 with the SAM instrument (Eigenbrode et al., 2018; Freissinet et al., 2015), fewer molecules  
12  
13 715 detected by SAM represent reduced carbon, with many of these organics being oxidized. For  
14  
15 716 example, polycyclic aromatic hydrocarbons (PAHs) and kerogen from exogenous delivery via  
16  
17 717 meteoritic infall may undergo several oxidation reactions to form stable simple organics such as  
18  
19 718 phthalic, mellitic, or oxalic acids (Benner et al., 2000; Freissinet et al., 2015) on the Martian  
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21 719 surface, which are then detectable by the SAM instrument (e.g. possible source organics for  
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23 720 benzene in the Sheepbed mudstone, Freissinet et al. (2015)). These oxidation processes are  
24  
25 721 expected to occur both on the surface and in the subsurface of Mars.

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27 722 Ionizing radiation is another serious concern for the long-term preservation of Martian organics.  
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29 723 The ionizing radiation dose on the surface of Mars is estimated at 0.54–0.85 Gray/year (Dartnell  
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31 724 et al., 2007). Ionizing radiation destroys biomolecules (e.g. proteins or DNA), and most bacteria  
32  
33 725 cannot survive doses greater than 200 Gray (Daly, 2009). Recent studies indicate that within 300  
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35 726 million years, preservation of functionalized organic molecules greater than 500 amu in the top ~5  
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37 727 cm of rock drops 1000x due to ionizing radiation (Pavlov et al., 2012). There is some indication  
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39 728 that rapid burial of organics, short exposure ages due to relatively slow Martian weathering rates,  
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41 729 and the characteristics of some entombing mineralogies may assist in organics preservation in the  
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43 730 Martian near-surface (Hays et al., 2017 and references therein), but the need remains to sample the  
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45 731 Martian subsurface to reduce the impact of degradation via radiation effects.

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47 732 The *Curiosity* rover can drill to a maximum of six centimeters and the *Rosalind Franklin* rover  
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49 733 will be able to drill to two meters depth. The work presented here indicates that even with  
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51 734 diagenetic effects and heterotrophic cycling of organics, organic molecules are preserved and  
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53 735 detectable over short timescales in the near-subsurface (2 to 18 cm in depth) of modern and relict  
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55 736 terrestrial hot springs. Organics therefore may be detectable on Mars in relatively recent hot spring  
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57 737 deposits, but preservation likely deteriorates over geologic time scales. These depths are within  
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59 738 the drill sampling range of the *Curiosity* and *Rosalind Franklin* rovers (partial to complete depths,  
60

739 respectively), and support the search for near-surface organics preserved in Martian hydrothermal  
740 systems.

## 741 **5. Conclusions**

742 Hydrothermal spring deposits host unique microbial ecosystems and have the capacity to  
743 physically preserve both microbial communities and their associated organic matter as  
744 biosignatures within siliceous sinter. These terrestrial geothermal springs serve as Mars analog  
745 environments to prepare future Mars missions to target their search for biosignatures in Martian  
746 hot spring deposits. This study quantified the preservation and detectability of fatty acids in three  
747 Icelandic hot spring deposits ranging from modern and inactive to relict, and samples were  
748 collected both at the surface and at depth (ranging from 2 to 18 cm in depth). Analyses performed  
749 with TMAH thermochemolysis pyrolysis-GC-MS indicate that fatty acids were preserved and  
750 detectable in all three spring deposits, with the subsurface samples degraded relative to the surface  
751 samples. TMAH thermochemolysis pyrolysis-GC-MS analyses using a SAM-instrument-like  
752 procedure were also successful in liberating preserved fatty acids from these sinter samples, albeit  
753 generally in lower abundances and with less diversity. These results build confidence in our ability  
754 to detect fatty acids and distinguish biogenicity using the GC-MS instrumentation available on  
755 current and future Mars missions.

## 756 **6. Acknowledgements**

757 Financial support for sample collection was provided by the NASA Planetary Science and  
758 Technology through Analog Research (PSTAR) program grant #NNX15AJ38G to J.R. Skok and  
759 the American Philosophical Society's Lewis and Clark Fund for Exploration and Field Research  
760 in Astrobiology to K. Craft. S.S. Johnson would like to acknowledge support from NASA grant  
761 #NNX17AJ68G ("Using Organic Molecule Detections in Mars Analog Environments to Interpret  
762 the Results of the SAM Investigation on the Mars Science Laboratory Mission") as well as a  
763 Georgetown University Pilot Research Grant.

## 764 **7. Author Disclosure Statement**

765 No competing financial interests exist.

**8. References**

- 766 Allen, C.C., Oehler, D.Z., 2008. A Case for Ancient Springs in Arabia Terra, Mars. *Astrobiology*  
767 8, 1093–1112. <https://doi.org/10.1089/ast.2008.0239>
- 769 Benner, S.A., Devine, K.G., Matveeva, L.N., Powell, D.H., 2000. The missing organic molecules  
770 on Mars. *Proc. Natl. Acad. Sci.* 97, 2425 LP – 2430.  
771 <https://doi.org/10.1073/pnas.040539497>
- 772 Biemann, K., Oro, J., Toulmin, P., Orgel, L.E., Nier, A.O., Anderson, D.M., Simmonds, P.G.,  
773 Flory, D., Diaz, A. V., Rushneck, D.R., Biller, J.E., Lafleur, A.L., 1977. The search for  
774 organic substances and inorganic volatile compounds in the surface of Mars. *J. Geophys.*  
775 *Res.* 82, 4641–4658. <https://doi.org/10.1029/JS082i028p04641>
- 776 Boynton, W. V., Bailey, S.H., Hamara, D.K., Williams, M.S., Bode, R.C., Fitzgibbon, M.R., Ko,  
777 W., Ward, M.G., Sridhar, K.R., Blanchard, J.A., Lorenz, R.D., May, R.D., Paige, D.A.,  
778 Pathare, A. V., Kring, D.A., Leshin, L.A., Ming, D.W., Zent, A.P., Golden, D.C., Kerry,  
779 K.E., Lauer, H.V., Quinn, R.C., 2001. Thermal and Evolved Gas Analyzer: Part of the Mars  
780 Volatile and Climate Surveyor integrated payload. *J. Geophys. Res. Planets* 106, 17683–  
781 17698. <https://doi.org/10.1029/1999JE001153>
- 782 Brasier, M.D., Antcliffe, J.B., Callow, R.H., 2011. Evolutionary trends in remarkable fossil  
783 preservation across the Ediacaran–Cambrian transition and the impact of metazoan mixing,  
784 in: *Taphonomy*. Springer, Dordrecht, pp. 519–567.
- 785 Bray, E., Evans, E., 1961. Distribution of n-paraffins as a clue to recognition of source beds.  
786 *Geochim. Cosmochim. Acta* 22, 2–15.
- 787 Breitbart, M., Wegley, L., Leeds, S., Schoenfeld, T., Rohwer, F., 2004. Phage Community  
788 Dynamics in Hot Springs. *Appl. Environ. Microbiol.* 70, 1633 LP – 1640.  
789 <https://doi.org/10.1128/AEM.70.3.1633-1640.2004>
- 790 Buch, A., Glavin, D.P., Sternberg, R., Szopa, C., Rodier, C., Navarro-González, R., Raulin, F.,  
791 Cabane, M., Mahaffy, P.R., 2006. A new extraction technique for in situ analyses of amino  
792 and carboxylic acids on Mars by gas chromatography mass spectrometry. *Planet. Space Sci.*  
793 54, 1592–1599. <https://doi.org/https://doi.org/10.1016/j.pss.2006.05.041>

- 1  
2  
3 794 Cady, S.L., Farmer, J.D., 1996. Fossilization processes in siliceous thermal springs: trends in  
4  
5 795 preservation along thermal gradients. *Evol. hydrothermal Ecosyst. Earth (and Mars?)* 150–  
6  
7 796 173. <https://doi.org/10.1002/9780470514986.ch9>  
8  
9 797 Campbell, K.A., Lynne, B.Y., Handley, K.M., Jordan, S., Farmer, J.D., Guido, D.M., 2015.  
10  
11 798 Tracing Biosignature Preservation of Geothermally Silicified Microbial Textures into the  
12  
13 799 Geological Record. *Astrobiology* 15, 858–882. <https://doi.org/10.1089/ast.2015.1307>  
14  
15 800 Castenholz, R.W., 1969. THE THERMOPHILIC CYANOPHYTES OF ICELAND AND THE  
16  
17 801 UPPER TEMPERATURE LIMIT. *J. Phycol.* 5, 360–368. <https://doi.org/10.1111/j.1529->  
18  
19 802 8817.1969.tb02626.x  
20  
21 803 Chaves Torres, L., Kaur, G., Melbourne, L.A., Pancost, R.D., 2019. Selective chemical  
22  
23 804 degradation of silica sinters of the Taupo Volcanic Zone (New Zealand). Implications for  
24  
25 805 early Earth and Astrobiology. *Geobiology* 0, 1–16. <https://doi.org/10.1111/gbi.12340>  
26  
27 806 Chefetz, B., Chen, Y., Clapp, C.E., Hatcher, P.G., 2000. Characterization of organic matter in  
28  
29 807 soils by thermochemolysis using tetramethylammonium hydroxide (TMAH). *Soil Sci. Soc.*  
30  
31 808 *Am. J.* 64, 583–589.  
32  
33 809 Cronin, J., Pizzarello, S., Epstein, S., Krishnamurthy, R., 1993. Molecular and isotopic analyses  
34  
35 810 of the hydroxy acids, dicarboxylic acids, and hydroxydicarboxylic acids of the Murchison  
36  
37 811 meteorite. *Geochim. Cosmochim. Acta* 57, 4745–4752. <https://doi.org/10.1016/0016->  
38  
39 812 7037(93)90197-5  
40  
41 813 Daly, M.J., 2009. A new perspective on radiation resistance based on *Deinococcus radiodurans*.  
42  
43 814 *Nat. Rev. Microbiol.* 7, 237–245. <https://doi.org/10.1038/nrmicro2073>  
44  
45 815 Daron, H.H., 1970. Fatty acid composition of lipid extracts of a thermophilic *Bacillus* species. *J.*  
46  
47 816 *Bacteriol.* 101, 145–151.  
48  
49 817 Dartnell, L.R., Desorgher, L., Ward, J.M., Coates, A.J., 2007. Modelling the surface and  
50  
51 818 subsurface Martian radiation environment: Implications for astrobiology. *Geophys. Res.*  
52  
53 819 *Lett.* 34, 1–6. <https://doi.org/10.1029/2006GL027494>  
54  
55 820 Davies, P.J., Bubela, B., Ferguson, J., 1978. The formation of ooids. *Sedimentology* 25, 703–  
56  
57 821 730.

- 1  
2  
3 822 del Rio, J.C., Martin, F., Gonzalez-Vila, F.J., 1996. Thermally assisted hydrolysis and alkylation  
4 823 as a novel pyrolytic approach for the structural characterization of natural biopolymers and  
5 824 geomacromolecules. *Trends Anal. Chem.* 15, 70–79.
- 6  
7  
8  
9 825 Deport, C., Lemee, L., Ambles, A., 2006. Comparison between humic substances from soil and  
10 826 peats using TMAH and TEAAc thermochemolysis. *Org. Geochem.* 37, 649–664.
- 11  
12  
13 827 Eigenbrode, J.L., Summons, R.E., Steele, A., Freissinet, C., Millan, M., Navarro-González, R.,  
14 828 Sutter, B., McAdam, A.C., Franz, H.B., Glavin, D.P., Archer, P.D., Mahaffy, P.R., Conrad,  
15 829 P.G., Hurowitz, J.A., Grotzinger, J.P., Gupta, S., Ming, D.W., Sumner, D.Y., Szopa, C.,  
16 830 Malespin, C., Buch, A., Coll, P., 2018. Organic matter preserved in 3-billion-year-old  
17 831 mudstones at Gale crater, Mars. *Science* (80-. ). 360, 1096–1101.  
18 832 <https://doi.org/10.1126/science.aas9185>
- 19  
20  
21  
22  
23  
24 833 Elias, V., Simoneit, B., Cardoso, J., 1997. Even N-Alkane Predominances on the Amazon Shelf  
25 834 and A Northeast Pacific Hydrothermal System. *Naturwissenschaften* 84, 415–420.
- 26  
27  
28 835 Folk, R.L., Lynch, F., 2001. Organic matter, putative nannobacteria and the formation of ooids  
29 836 and hardgrounds. *Sedimentology* 48, 215–229.
- 30  
31  
32 837 Freissinet, C., Glavin, D.P., Buch, A., Szopa, C., Teinturier, S., Archer, P.D., Williams, R.,  
33 838 Millan, M., Williams, A.J., Malespin, C., Mahaffy, P., 2019. Detection of long-chain  
34 839 hydrocarbons on mars with the Sample Analysis at Mars (SAM) instrument, in: *Mars 9*  
35 840 *Conference*. p. 6123.
- 36  
37  
38  
39  
40 841 Freissinet, C., Glavin, D.P., Mahaffy, P.R., Miller, K.E., Eigenbrode, J.L., Summons, R.E.,  
41 842 Brunner, A.E., Buch, A., Szopa, C., Archer, P.D., Franz, H.B., Atreya, S.K., Brinckerhoff,  
42 843 W.B., Cabane, M., Coll, P., Conrad, P.G., Marais, D.J. Des, Dworkin, J.P., Fairén, A.G.,  
43 844 François, P., Grotzinger, J.P., Kashyap, S., Kate, I.L., Leshin, L.A., Malespin, C.A., Martin,  
44 845 M.G., Mcadam, A.C., Ming, D.W., Pavlov, A.A., Prats, B.D., Squyres, S.W., Steele, A.,  
45 846 Stern, J.C., Sumner, D.Y., Sutter, B., Zorzano, M., 2015. Organic molecules in the  
46 847 Sheepbed Mudstone, Gale Crater, Mars. *J. Geophys. Res.* 120, 495–514.  
47 848 <https://doi.org/10.1002/2014JE004737>.Received
- 48  
49  
50  
51  
52  
53  
54 849 Frydenvang, J., Gasda, P.J., Hurowitz, J.A., Grotzinger, J.P., Wiens, R.C., Newsom, H.E.,  
55  
56  
57  
58  
59  
60

- 1  
2  
3 850 Edgett, K.S., Watkins, J., Bridges, J.C., Maurice, S., Fisk, M.R., Johnson, J.R., Rapin, W.,  
4 851 Stein, N.T., Clegg, S.M., Schwenzer, S.P., Bedford, C.C., Edwards, P., Mangold, N.,  
5 852 Cousin, A., Anderson, R.B., Payré, V., Vaniman, D., Blake, D.F., Lanza, N.L., Gupta, S.,  
6 853 Van Beek, J., Sautter, V., Meslin, P.-Y., Rice, M., Milliken, R., Gellert, R., Thompson, L.,  
7 854 Clark, B.C., Sumner, D.Y., Fraeman, A.A., Kinch, K.M., Madsen, M.B., Mitrofanov, I.G.,  
8 855 Jun, I., Calef, F., Vasavada, A.R., 2017. Diagenetic silica enrichment and late-stage  
9 856 groundwater activity in Gale crater, Mars. *Geophys. Res. Lett.* 44, 4716–4724.  
10 857 <https://doi.org/10.1002/2017GL073323>  
11  
12 858 Geffroy-Rodier, C., Grasset, L., Sternberg, R., Buch, A., Amblès, A., 2009. Thermochemolysis  
13 859 in search for organics in extraterrestrial environments. *J. Anal. Appl. Pyrolysis* 85, 454–459.  
14 860 <https://doi.org/https://doi.org/10.1016/j.jaap.2008.10.005>  
15  
16 861 Georgiou, C.D., Deamer, D.W., 2014. Lipids as Universal Biomarkers of Extraterrestrial Life.  
17 862 *Astrobiology* 14, 541–549.  
18  
19 863 Gibson, R.A., Sherry, A., Kaur, G., Pancost, R.D., Talbot, H.M., 2014. Bacteriohopanepolyols  
20 864 preserved in silica sinters from Champagne Pool (New Zealand) indicate a declining  
21 865 temperature gradient over the lifetime of the vent. *Org. Geochem.* 69, 61–69.  
22 866 <https://doi.org/https://doi.org/10.1016/j.orggeochem.2014.02.004>  
23  
24 867 Goesmann, F., Brinckerhoff, W.B., Raulin, F., Goetz, W., Danell, R., Getty, S., Siljeström, S.,  
25 868 Steininger, H., Arevalo Jr., R., Buch, A., Freissinet, C., Grubisi, A., Meierhenrich, U.,  
26 869 Mißbach, H., Pinnick, V., Stalport, F., Szopa, C., Brucato, J.R., Glavin, D.P., Grand, N., Li,  
27 870 X., van Amerom, F.H.W., the MOMA Science Team, 2017. The Mars Organic Molecule  
28 871 Analyzer (MOMA) Instrument: Characterization of Organic Material in Martian Sediments.  
29 872 *Astrobiology* 17, 655–685. <https://doi.org/10.1089/ast.2016.1551>  
30  
31 873 Goetz, W., Brinckerhoff, W.B., Arevalo, R., Freissinet, C., Getty, S., Glavin, D.P., Siljeström, S.,  
32 874 Buch, A., Stalport, F., Grubisi, A., Li, X., Pinnick, V., Danell, R., van Amerom, F.H.W.,  
33 875 Goesmann, F., Steininger, H., Grand, N., Raulin, F., Szopa, C., Meierhenrich, U., Brucato,  
34 876 J.R., 2016. MOMA: the challenge to search for organics and biosignatures on Mars. *Int. J.*  
35 877 *Astrobiol.* 15, 239–250. <https://doi.org/DOI: 10.1017/S1473550416000227>  
36  
37 878 Grasset, L., Guignard, C., Amblès, A., 2002. Free and esterified aliphatic carboxylic acids in



- 879 humin and humic acids from a peat sample as revealed by pyrolysis with  
880 tetramethylammonium hydroxide or tetraethylammonium acetate. *Org. Geochem.* 33, 181–  
881 188. [https://doi.org/http://dx.doi.org/10.1016/S0146-6380\(01\)00150-4](https://doi.org/http://dx.doi.org/10.1016/S0146-6380(01)00150-4)
- 882 Grimalt, J., Albaiges, J., 1987. Sources and occurrence of C<sub>12</sub> to C<sub>22n</sub>-alkane distributions with  
883 even carbon-number preference in sedimentary environments. *Geochim. Cosmochim. Acta*  
884 51, 1379–1384.
- 885 Guido, D.M., Channing, A., Campbell, K.A., Zamuner, A., 2010. Jurassic geothermal landscapes  
886 and fossil ecosystems at San Agustín, Patagonia, Argentina. *J. Geol. Soc. London.* 167, 11–  
887 20.
- 888 Guignard, C., Lemee, L., Ambles, A., 2005. Lipid constituents of peat humic acids and humin.  
889 Distinction from directly extractable bitumen components using TMAH and TEAAc  
890 thermochemolysis. *Org. Geochem.* 36, 287–297.
- 891 Hays, L.E., Achenbach, L.A., Bailey, J., Barnes, R., Baross, J., Bertka, C., Boston, P.J., Boyd,  
892 E.S., Cable, M., Chen, I., Ciesla, F., Des Marais, D.J., Domagal-Goldman, S.D., Cook, J.E.,  
893 Goldman, A., Hud, N., Laine, P., Lloyd, K., Lyons, T.W., Meadows, V.S., Mix, L.J.,  
894 Mojzsis, S.J., Muller, S., Pasek, M., Powell, M., Robinson, T., Rosenzweig, F., Schmidt, B.,  
895 Seelig, B., Springsteen, G., Vance, S., Welander, P., Williams, L., Wordsworth, R., 2015.  
896 NASA Astrobiology Strategy 2015.
- 897 Hays, L.E.L.E., Graham, H.V.H.V., Des Marais, D.J.D.J., Hausrath, E.M.E.M., Horgan, B.,  
898 McCollom, T.M.T.M., Parenteau, M.N.N., Potter-McIntyre, S.L., Williams, A.J.A.J.,  
899 Lynch, K.L.K.L., 2017. Biosignature Preservation and Detection in Mars Analog  
900 Environments. *Astrobiology* 17, 363–400. <https://doi.org/10.1089/ast.2016.1627>
- 901 He, Y., Buch, A., Morisson, M., Szopa, C., Freissinet, C., Williams, A.J., Millan, M., Guzman,  
902 M., Gonzalez, R.N., Bonnet, J.Y., Coscia, D., Mahaffy, P., Johnson, S.S., 2019. Application  
903 of TMAH thermochemolysis to the detection of nucleobases: application to the MOMA and  
904 SAM space experiment. *Talanta* submitted.
- 905 Herdianita, N.R., Browne, P.R.L., Rodgers, K.A., Campbell, K.A., 2000. Mineralogical and  
906 textural changes accompanying ageing of silica sinter. *Miner. Depos.* 35, 48–62.

- 1  
2  
3 907 <https://doi.org/10.1007/s001260050005>  
4  
5  
6 908 Jahnke, L.L., Eder, W., Huber, R., Hope, J.M., Hinrichs, K.U., Hayes, J.M., Des Marais, D.J.,  
7  
8 909 Cady, S.L., Summons, R.E., 2001. Signature lipids and stable carbon isotope analyses of  
9  
10 910 Octopus Spring hyperthermophilic communities compared with those of Aquificales  
11  
12 911 representatives. *Appl. Environ. Microbiol.* 67, 5179–5189.  
13  
14 912 <https://doi.org/10.1128/AEM.67.11.5179-5189.2001>  
15  
16 913 Johnson, S.S., Millan, M., Graham, H., Benison, K.C., Williams, A.J., Mcadam, A., Knudson,  
17  
18 914 C.A., Andrejkovic, S., Achilles, C.N., 2020. Lipid Biomarkers in Ephemeral Acid Salt Lake  
19  
20 915 Mudflat/Sandflat Sediments: Implications for Mars. *Astrobiology* 20, 167–178.  
21  
22 916 <https://doi.org/10.1089/ast.2017.1812>  
23  
24 917 Jones, B., Renaut, R.W., 1996. Influence of thermophilic bacteria on calcite and silica  
25  
26 918 precipitation in hot springs with water temperatures above 90 °C: evidence from Kenya and  
27  
28 919 New Zealand. *Can. J. Earth Sci.* 33, 72–83. <https://doi.org/10.1139/e96-008>  
29  
30 920 Jones, B., Renaut, R.W., Rosen, M.R., 2000. Stromatolites Forming in Acidic Hot-Spring  
31  
32 921 Waters, North Island, New Zealand. *Palaios* 15, 450–475. [https://doi.org/10.1669/0883-](https://doi.org/10.1669/0883-1351(2000)015<0450:SFIAHS>2.0.CO;2)  
33  
34 922 [1351\(2000\)015<0450:SFIAHS>2.0.CO;2](https://doi.org/10.1669/0883-1351(2000)015<0450:SFIAHS>2.0.CO;2)  
35  
36 923 Jones, B., Renaut, R.W., Rosen, M.R., 1997. Biogenicity of silica precipitation around geysers  
37  
38 924 and hot-spring vents, North Island, New Zealand. *J. Sediment. Res.* 67, 88–104.  
39  
40 925 Kashefi, K., Lovley, D.R., 2003. Extending the Upper Temperature Limit for Life. *Science* (80-  
41  
42 926 ). 301, 934 LP – 934. <https://doi.org/10.1126/science.1086823>  
43  
44 927 Kaur, G., Mountain, B.W., Hopmans, E.C., Pancost, R.D., 2011. Preservation of Microbial  
45  
46 928 Lipids in Geothermal Sinters. *Astrobiology* 11, 259–274.  
47  
48 929 <https://doi.org/10.1089/ast.2010.0540>  
49  
50 930 Kaur, G., Mountain, B.W., Pancost, R.D., 2008. Microbial membrane lipids in active and  
51  
52 931 inactive sinters from Champagne Pool, New Zealand: Elucidating past geothermal  
53  
54 932 chemistry and microbiology. *Org. Geochem.* 39, 1024–1028.  
55  
56 933 <https://doi.org/https://doi.org/10.1016/j.orggeochem.2008.04.016>  
57  
58 934 Kaur, G., Mountain, B.W., Stott, M.B., Hopmans, E.C., Pancost, R.D., 2015. Temperature and

- 1  
2  
3 935 pH control on lipid composition of silica sinters from diverse hot springs in the Taupo  
4 936 Volcanic Zone, New Zealand. *Extremophiles* 19, 327–344. <https://doi.org/10.1007/s00792->  
5  
6 937 014-0719-9  
7  
8  
9 938 Killops, S., Killops, V., 2005. *Introduction to Organic Geochemistry*, 2nd ed. Blackwell  
10 939 Publishing.  
11  
12  
13 940 Konhauser, K.O., Phoenix, V.R., Bottrell, S.H., Adams, D.G., Head, I.M., 2001. Microbial–silica  
14 941 interactions in Icelandic hot spring sinter: possible analogues for some Precambrian  
15 942 siliceous stromatolites. *Sedimentology* 48, 415–433. <https://doi.org/10.1046/j.1365->  
16 943 3091.2001.00372.x  
17  
18  
19  
20 944 Larter, S.R., Horsfield, B., 1993. Determination of structural components of kerogens by the use  
21 945 of analytical pyrolysis methods, in: Engel, M.H., Macko, S.A. (Eds.), *Organic*  
22 946 *Geochemistry*. Plenum Press, pp. 271–288.  
23  
24  
25  
26 947 Leira, M., Meijide-Failde, R., Torres, E., 2017. Diatom communities in thermo-mineral springs  
27 948 of Galicia (NW Spain). *Diatom Res.* 32, 29–42.  
28 949 <https://doi.org/10.1080/0269249X.2017.1286266>  
29  
30  
31  
32 950 Mahaffy, P.R., Webster, C.R., Cabane, M., Conrad, P.G., Coll, P., Atreya, S.K., Arvey, R.,  
33 951 Barciniak, M., Benna, M., Bleacher, L., Brinckerhoff, W.B., Eigenbrode, J.L., Carignan, D.,  
34 952 Cascia, M., Chalmers, R.A., Dworkin, J.P., Errigo, T., Everson, P., Franz, H., Farley, R.,  
35 953 Feng, S., Frazier, G., Freissinet, C., Glavin, D.P., Harpold, D.N., Hawk, D., Holmes, V.,  
36 954 Johnson, C.S., Jones, A., Jordan, P., Kellogg, J., Lewis, J., Lyness, E., Malespin, C.A.,  
37 955 Martin, D.K., Maurer, J., McAdam, A.C., McLennan, D., Nolan, T.J., Noriega, M., Pavlov,  
38 956 A.A., Prats, B., Raaen, E., Sheinman, O., Sheppard, D., Smith, J., Stern, J.C., Tan, F.,  
39 957 Trainer, M., Ming, D.W., Morris, R. V, Jones, J., Gundersen, C., Steele, A., Wray, J., Botta,  
40 958 O., Leshin, L.A., Owen, T., Battel, S., Jakosky, B.M., Manning, H., Squyres, S.,  
41 959 Christopher, R.N., Francois, P.M., Sternberg, R., Buch, A., Sorensen, P., Coscia, D., Szopa,  
42 960 C., Teinturier, S., Baffes, C., Feldman, J., Flesch, G., Forouhar, S., Garcia, R., Keymeulen,  
43 961 D., Woodward, S., Block, B.P., Arnett, K., Miller, R., Edmonson, C., Gorevan, S., 2012.  
44 962 *The Sample Analysis at Mars Investigation and Instrument Suite*. *Space Sci. Rev.* 1–78.  
45 963 <https://doi.org/10.1007/s11214-012-9879-z>  
46  
47  
48  
49  
50  
51  
52  
53  
54  
55  
56  
57  
58  
59  
60

- 1  
2  
3 964 McCollom, T.M., Hynek, B.M., Rogers, K.L., 2016. Potential for preservation of biosignatures  
4 965 from endolithic microbial communities in a Mars analog fumarole environment, in:  
5 966 Biosignature Preservation and Detection in Mars Analog Environments. p. 2006.  
6  
7  
8  
9 967 McCollom, T.M., Ritter, G., Simoneit, B., 1999. Lipid Synthesis under Hydrothermal Conditions  
10 968 by Fischer-Tropsch-Type Reactions. *Orig. Life Evol. Biosph.* 29, 153–166.  
11  
12  
13 969 Metcalffe, L.D., Wang, C.N., 1981. Rapid Preparation of Fatty Acid Methyl Esters Using  
14 970 Organic Base-Catalyzed Transesterification. *J. Chromatogr. Sci.* 19, 530–535.  
15  
16  
17 971 Millan, M., Malespin, C.A., Freissinet, C., Glavin, D.P., Mahaffy, P.R., Buch, A., Szopa, C.,  
18 972 Srivastava, A., Teinturier, S., Williams, R., Williams, A.J., McAdam, A., Coscia, D.,  
19 973 Eigenbrode, J., Raaen, E., Dworkin, J., Navarro-Gonzalez, R., Johnson, S.S., 2019. Lessons  
20 974 learned from the full cup wet chemistry experiment performed on mars with the sample  
21 975 analysis at mars instrument, in: Mars 9 Conference. p. 6210.  
22  
23  
24  
25  
26 976 Munoz-Saez, C., Salthiel, S., Manga, M., Nguyen, C., Gonnermann, H., 2016. Physical and  
27 977 hydraulic properties of modern sinter deposits: El Tatio, Atacama. *J. Volcanol. Geotherm.*  
28 978 *Res.* 325, 156–168.  
29  
30  
31  
32 979 Nishimura, M., Baker, E., 1986. Possible origin of n-alkanes with a remarkable even-to-odd  
33 980 predominance in recent marine sediments. *Geochim. Cosmochim. Acta* 50, 299–305.  
34  
35  
36 981 O'Reilly, S.S., Mariotti, G., Winter, A.R., Newman, S.A., Matys, E.D., McDermott, F., Pruss,  
37 982 S.B., Bosak, T., Summons, R.E., Klepac-Ceraj, V., 2017. Molecular biosignatures reveal  
38 983 common benthic microbial sources of organic matter in ooids and grapestones from Pigeon  
39 984 Cay, The Bahamas. *Geobiology* 112–130. <https://doi.org/10.1111/gbi.12196>  
40  
41  
42  
43  
44 985 Oshima, M., Miyagawa, A., 1974. Comparative studies on the fatty acid composition of  
45 986 moderately and extremely thermophilic bacteria. *Lipids* 9, 476–480.  
46 987 <https://doi.org/10.1007/BF02534274>  
47  
48  
49  
50 988 Pancost, R.D., Pressley, S., Coleman, J.M., Benning, L.G., Mountain, B.W., 2005. Lipid  
51 989 biomolecules in silica sinters: indicators of microbial biodiversity. *Environ. Microbiol.* 7,  
52 990 66–77. <https://doi.org/10.1111/j.1462-2920.2004.00686.x>  
53  
54  
55  
56 991 Pancost, R.D., Pressley, S., Coleman, J.M., Talbot, H.M., Kelly, S.P., Farrimond, P., Schouten,  
57  
58  
59  
60

- 1  
2  
3 992 S., Benning, L., Mountain, B.W., 2006. Composition and implications of diverse lipids in  
4 993 New Zealand Geothermal sinters. *Geobiology* 4, 71–92. <https://doi.org/10.1111/j.1472->  
5 994 [4669.2006.00069.x](https://doi.org/10.1111/j.1472-4669.2006.00069.x)  
6  
7  
8  
9 995 Pavlov, A.A., Vasilyev, G., Ostryakov, V.M., Pavlov, A.K., Mahaffy, P., 2012. Degradation of  
10 996 the organic molecules in the shallow subsurface of Mars due to irradiation by cosmic rays.  
11 997 *Geophys. Res. Lett.* 39. <https://doi.org/10.1029/2012GL052166>  
12  
13  
14  
15 998 Pulchan, J., Abrajano, T.A., Helleur, R., 1997. Characterization of tetramethylammonium  
16 999 hydroxide thermochemolysis products of near-shore marine sediments using gas  
17 1000 chromatography/mass spectrometry and gas chromatography/combustion/isotope ratio mass  
18 1001 spectrometry. *J. Anal. Appl. Pyrolysis* 42, 135–150.  
19  
20  
21  
22 1002 Pumas, C., Pruetiworanan, S., Peerapornpisal, Y., 2018. Diatom diversity in some hot springs of  
23 1003 northern Thailand. *Botanica* 24, 69–86. [https://doi.org/https://doi.org/10.2478/botlit-2018-](https://doi.org/10.2478/botlit-2018-)  
24 1004 [0007](https://doi.org/10.2478/botlit-2018-0007)  
25  
26  
27  
28 1005 Ray, P.H., White, D.C., Brock, T.D., 1971. Effect of temperature on the fatty acid composition  
29 1006 of *Thermus aquaticus*. *J. Bacteriol.* 106, 25–30.  
30  
31  
32 1007 Renaut, R.W., Jones, B., 2011. Hydrothermal environments, terrestrial, in: Reitner, J., Thiel, V.  
33 1008 (Eds.), *Encyclopedia of Geobiology*. Springer, pp. 467–479.  
34  
35  
36 1009 Renaut, R.W., Jones, B., Tiercelin, J.J., 1998. Rapid in situ silicification of microbes at Loburu  
37 1010 hot springs, Lake Bogoria, Kenya rift valley. *Sedimentology* 45, 1083–1103.  
38 1011 <https://doi.org/10.1046/j.1365-3091.1998.00194.x>  
39  
40  
41  
42 1012 Rossi, A.P., Neukum, G., Pondrelli, M., van Gasselt, S., Zegers, T., Hauber, E., Chicarro, A.,  
43 1013 Foing, B., 2008. Large-scale spring deposits on Mars? *J. Geophys. Res. E Planets* 113, 1–  
44 1014 17. <https://doi.org/10.1029/2007JE003062>  
45  
46  
47  
48 1015 Ruff, S.W., Farmer, J.D., 2016. Silica deposits on Mars with features resembling hot spring  
49 1016 biosignatures at El Tatio in Chile. *Nat. Commun.* 7, 13554.  
50  
51  
52 1017 Ruff, S.W., Farmer, J.D., Calvin, W.M., Herkenhoff, K.E., Johnson, J.R., Morris, R. V, Rice,  
53 1018 M.S., Arvidson, R.E., Bell III, J.F., Christensen, P.R., Squyres, S.W., 2011. Characteristics,  
54 1019 distribution, origin, and significance of opaline silica observed by the Spirit rover in Gusev  
55  
56  
57

- 1  
2  
3 1020 crater, Mars. *J. Geophys. Res. E Planets* 116. <https://doi.org/10.1029/2010JE003767>
- 4  
5  
6 1021 Russell, N.J., 1984. Mechanisms of thermal adaptation in bacteria: blueprints for survival.  
7  
8 1022 *Trends Biochem. Sci.* 9, 108–112. <https://doi.org/https://doi.org/10.1016/0968->  
9 1023 0004(84)90106-3
- 10  
11 1024 Saini, J., Kaur, R., Sharma, P.K., 2017. Metagenomic Insights into Microbial Diversity and  
12 1025 Metabolic Potential of Hot Spring Ecosystems BT - Mining of Microbial Wealth and  
13 1026 MetaGenomics, in: Kalia, V.C., Shouche, Y., Purohit, H.J., Rahi, P. (Eds.), . Springer  
14 1027 Singapore, Singapore, pp. 143–159. [https://doi.org/10.1007/978-981-10-5708-3\\_9](https://doi.org/10.1007/978-981-10-5708-3_9)
- 15  
16  
17  
18  
19 1028 Schulten, H.R., 1996. Direct pyrolysis-mass spectrometry of soils: a novel tool in agriculture,  
20 1029 ecology, forestry and soil science, in: Yamasaki, S., Boutton, T.W. (Eds.), *Mass*  
21 1030 *Spectrometry of Soils*. Marcel Dekker, New York, pp. 373–436.
- 22  
23  
24  
25 1031 Segerer, A.H., Burggraf, S., Fiala, G., Huber, G., Huber, R., Pley, U., Stetter, K.O., 1993. Life in  
26 1032 hot springs and hydrothermal vents. *Orig. Life Evol. Biosph.* 23, 77–90.
- 27  
28  
29 1033 Shen, P.Y., Coles, E., Foote, J.L., Stenesh, J., 1970. Fatty Acid Distribution in Mesophilic and  
30 1034 Thermophilic Strains of the Genus <em>Bacillus</em>; *J. Bacteriol.* 103, 479  
31 1035 LP – 481.
- 32  
33  
34  
35 1036 Shiea, J., Brassel, S.C., Ward, D.M., 1991. Comparative analysis of extractable lipids in hot  
36 1037 spring microbial mats and their component photosynthetic bacteria. *Org. Geochem.* 17,  
37 1038 309–319. [https://doi.org/https://doi.org/10.1016/0146-6380\(91\)90094-Z](https://doi.org/https://doi.org/10.1016/0146-6380(91)90094-Z)
- 38  
39  
40  
41 1039 Skirmisdottir, S., Hreggvidsson, G.O., Hjörleifsdottir, S., Marteinson, V.T., Petursdottir, S.K.,  
42 1040 Holst, O., Kristjansson, J.K., 2000. Influence of Sulfide and Temperature on Species  
43 1041 Composition and Community Structure of Hot Spring Microbial Mats. *Appl. Environ.*  
44 1042 *Microbiol.* 66, 2835 LP – 2841. <https://doi.org/10.1128/AEM.66.7.2835-2841.2000>
- 45  
46  
47  
48 1043 Skok, J.R., Mustard, J.F., Ehlmann, B.L., Milliken, R.E., Murchie, S.L., 2010. Silica deposits in  
49 1044 the Nili Patera caldera on the Syrtis Major volcanic complex on Mars. *Nat. Geosci.* 3, 838–  
50 1045 841.
- 51  
52  
53  
54 1046 Squyres, S.W., Arvidson, R.E., Ruff, S., Gellert, R., Morris, V., Ming, D.W., Crumpler, L.,  
55 1047 Farmer, J.D., Marais, D.J. Des, Yen, A., McLennan, S.M., Calvin, W., Bell, J.F., B. C.

- 1  
2  
3 1048 Clark, H. A. Wang, T. J. McCoy, M. E. Schmidt, P. A. de Souza Jr. 2008.  
4  
5 1049 Detection of Silica-Rich Deposits on Mars. *Science* (80-. ). 320, 1063–1067.  
6  
7 1050 <https://doi.org/10.1126/science.1155429>  
8  
9 1051 Summons, R.E., Albrecht, P., McDonald, G., Moldowan, J.M., 2008. Molecular Biosignatures.  
10  
11 1052 *Space Sci. Rev.* 135, 133–159. <https://doi.org/10.1007/s11214-007-9256-5>  
12  
13 1053 Summons, R.E., Bird, L.R., Gillespie, A.L., Pruss, S.B., Roberts, M., Sessions, A.L., 2013. Lipid  
14  
15 1054 biomarkers in ooids from different locations and ages: evidence for a common bacterial  
16  
17 1055 flora. *Geobiology* 11, 420–436. <https://doi.org/10.1111/gbi.12047>  
18  
19 1056 Sun, M.-Y., Wakeham, S., Lee, C., 1997. Rates and mechanisms of fatty acid degradation in oxic  
20  
21 1057 and anoxic coastal marine sediments of Long Island Sound, New York, USA. *Geochim.*  
22  
23 1058 *Cosmochim. Acta* 61, 341–355.  
24  
25 1059 Tarhan, L.G., Hood, A. v. S., Droser, M.L., Gehling, J.G., Briggs, D.E.G., 2016. Exceptional  
26  
27 1060 preservation of soft-bodied Ediacara Biota promoted by silica-rich oceans. *Geology* 44,  
28  
29 1061 951–954. <https://doi.org/10.1130/G38542.1>  
30  
31 1062 Teece, B.L., George, S.C., Djokic, T., Campbell, K.A., Ruff, S.W., Kranendonk, M.J. Van, 2020.  
32  
33 1063 Biomolecules from fossilized hot-spring sinters: implications for the search for life on Mars.  
34  
35 1064 *Astrobiology* 20, 537–551.  
36  
37 1065 Vestal, J.R., White, D.C., 1989. Lipid Analysis in Microbial Ecology. *Bioscience* 39, 535–541.  
38  
39 1066 Volkman, J., 2006. Lipid biomarkers for marine organic matter, in: Volkman, J. (Ed.), *Marine*  
40  
41 1067 *Organic Matter: Biomarkers, Isotopes and DNA*. Heidelberg: Springer-Verlag, Berlin, pp.  
42  
43 1068 27–70.  
44  
45 1069 Volkman, J., Jeffrey, S., Nichols, P., Rogers, G., Farland, C., 1989. Fatty acid and lipid  
46  
47 1070 composition of 10 species of microalgae used in mariculture. *J. Exp. Mar. Bio. Ecol.* 128.  
48  
49 1071 Walter, M.R., 1996. Ancient hydrothermal ecosystems on earth: a new palaeobiological frontier,  
50  
51 1072 in: *Ciba Foundation Symposium*. pp. 112–127 (discussion 127–130).  
52  
53 1073 Walter, M.R., 1976. Hot-spring sediments in Yellowstone National Park, in: Walter, M.R. (Ed.),  
54  
55 1074 *Stromatolites*. Elsevier, Amsterdam, pp. 489–198.  
56  
57  
58  
59  
60

- 1  
2  
3 1075 Walter, M.R., Des Marais, D.J., 1993. Preservation of Biological Information in Thermal Spring  
4 1076 Deposits: Developing a Strategy for the Search for Fossil Life on Mars. *Icarus* 101, 129–  
5 1077 143.
- 6  
7  
8  
9 1078 Ward, D.M., Castenholz, R.W., Miller, S.R., 2012. Cyanobacteria in Geothermal Habitats, in:  
10 1079 Whitton, B.A. (Ed.), *Ecology of Cyanobacteria II: Their Diversity in Space and Time*.  
11 1080 Springer, pp. 39–63. [https://doi.org/10.1007/978-94-007-3855-3\\_3](https://doi.org/10.1007/978-94-007-3855-3_3)
- 12  
13  
14  
15 1081 Weerkamp, A., Heinen, W., 1972. Effect of temperature on the fatty acid composition of the  
16 1082 extreme thermophiles, *Bacillus caldolyticus* and *Bacillus caldotenax*. *J. Bacteriol.* 109, 443–  
17 1083 446.
- 18  
19  
20  
21 1084 Westall, F., Campbell, K.A., Bréhéret, J.G., Foucher, F., Gautret, P., Hubert, A., Sorieul, S.,  
22 1085 Grassineau, N., Guido, D.M., 2015. Archean (3.33 Ga) microbe-sediment systems were  
23 1086 diverse and flourished in a hydrothermal context. *Geology* 43, 615–618.  
24 1087 <https://doi.org/10.1130/G36646.1>
- 25  
26  
27  
28 1088 Wilhelm, M.B., Davila, A.F., Eigenbrode, J.L., Parenteau, M.N., Jahnke, L.L., Liu, X.,  
29 1089 Summons, R.E., Wray, J.J., Stamos, B.N., O'Reilly, S.S., Williams, A.J., 2017.  
30 1090 Xeropreservation of functionalized lipid biomarkers in hyperarid soils in the Atacama  
31 1091 Desert. *Org. Geochem.* 103, 97–104. <https://doi.org/10.1016/j.orggeochem.2016.10.015>
- 32  
33  
34  
35  
36 1092 Williams, A.J., Eigenbrode, J., Floyd, M., Wilhelm, M.B., O'Reilly, S., Johnson, S.S., Craft,  
37 1093 K.L., Knudson, C.A., Andrejkovičová, S., Lewis, J.M.T., Buch, A., Glavin, D.P., Freissinet,  
38 1094 C., Williams, R.H., Szopa, C., Millan, M., Summons, R.E., McAdam, A., Benison, K.,  
39 1095 Navarro-González, R., Malespin, C., Mahaffy, P.R., 2019. Recovery of Fatty Acids from  
40 1096 Mineralogic Mars Analogs by TMAH Thermochemolysis for the Sample Analysis at Mars  
41 1097 Wet Chemistry Experiment on the Curiosity Rover. *Astrobiology* 19, 522–546.  
42 1098 <https://doi.org/10.1089/ast.2018.1819>
- 43  
44  
45  
46  
47  
48  
49 1099 Zhang, C.L., Fouke, B.W., Bonheyo, G.T., Peacock, A.D., White, D.C., Huang, Y., Romanek,  
50 1100 C.S., 2004. Lipid biomarkers and carbon-isotopes of modern travertine deposits  
51 1101 (Yellowstone National Park, USA): Implications for biogeochemical dynamics in hot-spring  
52 1102 systems. *Geochim. Cosmochim. Acta* 68, 3157–3169.  
53 1103 <https://doi.org/https://doi.org/10.1016/j.gca.2004.03.005>
- 54  
55  
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2  
3 1104 Zhang, C.L., Huang, Z., Li, Y.-L., Romanek, C.S., Mills, G.L., Gibson, R.A., Talbot, H.M.,  
4  
5 1105 Wiegel, J., Noakes, J., Culp, R., White, D.C., 2007. Lipid Biomarkers, Carbon Isotopes, and  
6  
7 1106 Phylogenetic Characterization of Bacteria in California and Nevada Hot Springs.  
8  
9 1107 Geomicrobiol. J. 24, 519–534. <https://doi.org/10.1080/01490450701572515>

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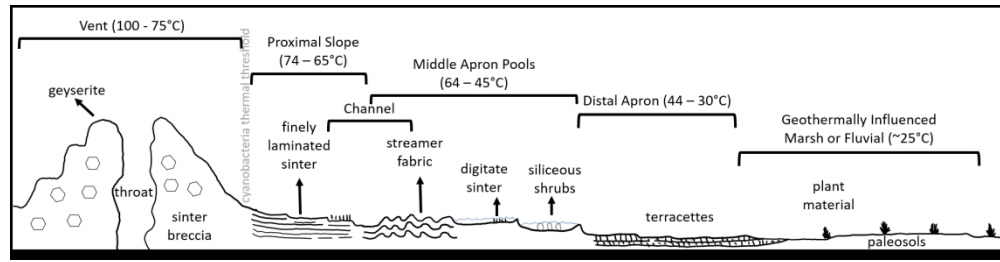


Figure 1. Environmental settings and facies of a generalized geothermal spring system. Modified from Campbell et al. (2015).

321x81mm (150 x 150 DPI)

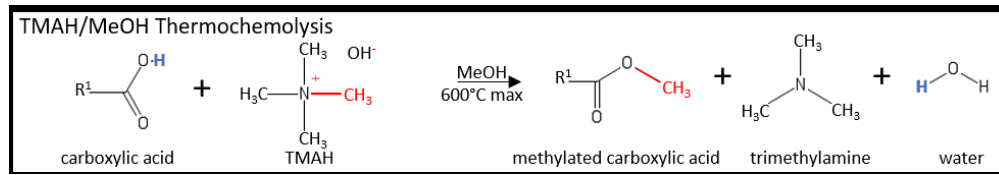


Figure 2. Example TMAH thermochemolysis reaction with a carboxylic acid to form a methylated carboxylic acid. After the reaction, the molecule is more volatile and detectable to GC-MS. TMAH in methanol is available on the SAM and MOMA instruments.

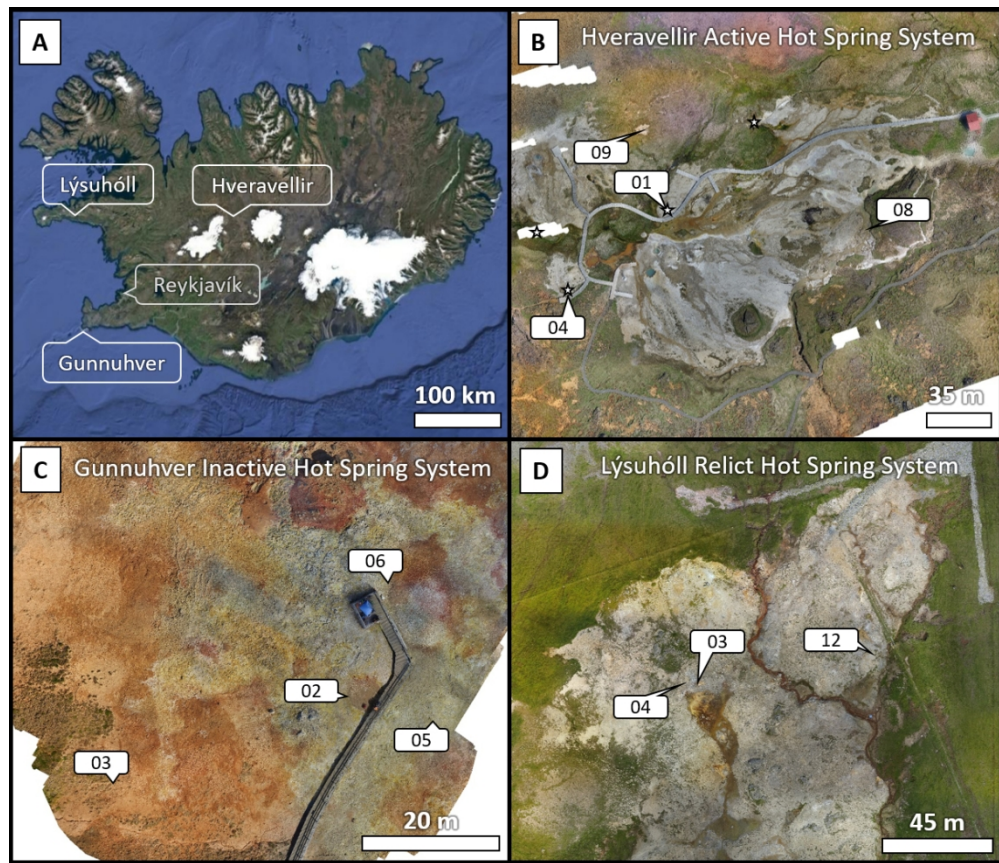


Figure 3. A) Map of field locations in Iceland, with the capitol of Reykjavík for reference. B-D) Field site numbers at the Hveravellir, Gunnuhver, and Lysuhöll spring systems. Stars at Hveravellir indicate active flowing spring vents where geochemical measurements were taken.

209x180mm (150 x 150 DPI)

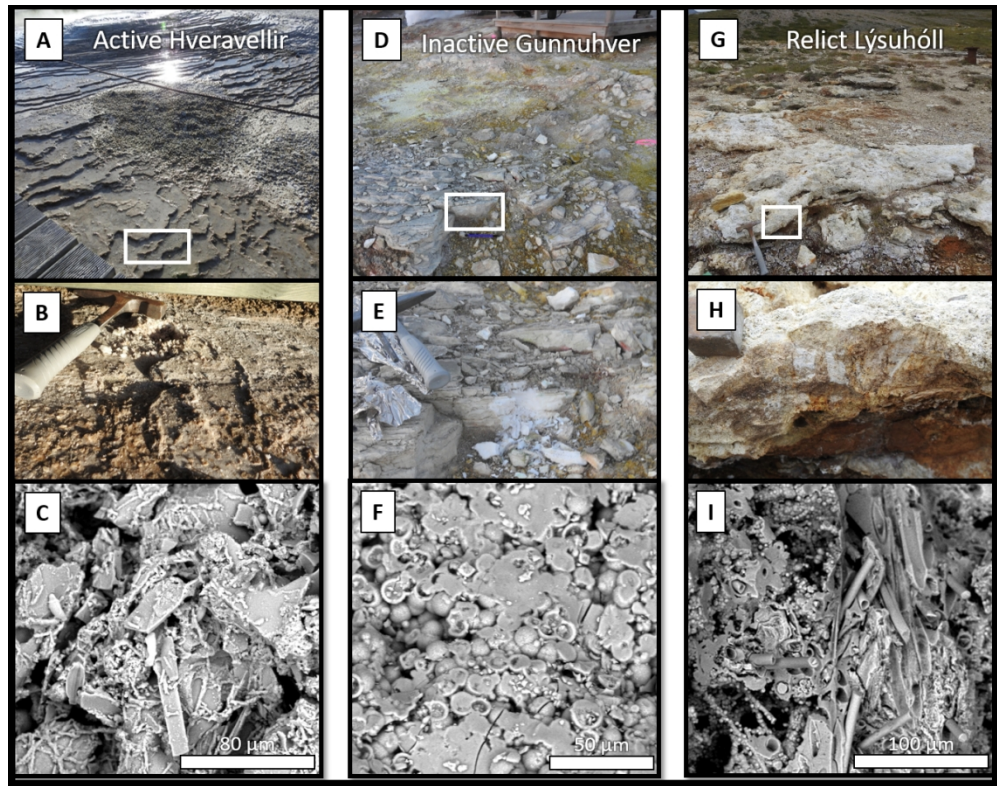


Figure 4. Field sample site images and corresponding SEM images of sinter from A-C) Hveravellir (active vent in top left of panel A), D-F) Gunnuhver, and G-I) Lýsuhóll. White box indicates zoom image B, E, or H. Rock hammer for scale in surface images. C) Silicified diatoms, microbial filaments, and putative plant matter present in the Hveravellir site. F) Opal-CT accretionary spheres with no confirmed biogenic features present in the Gunnuhver site. I) Abundant silicified casts of microbial filaments present in the Lýsuhóll site.

229x179mm (150 x 150 DPI)

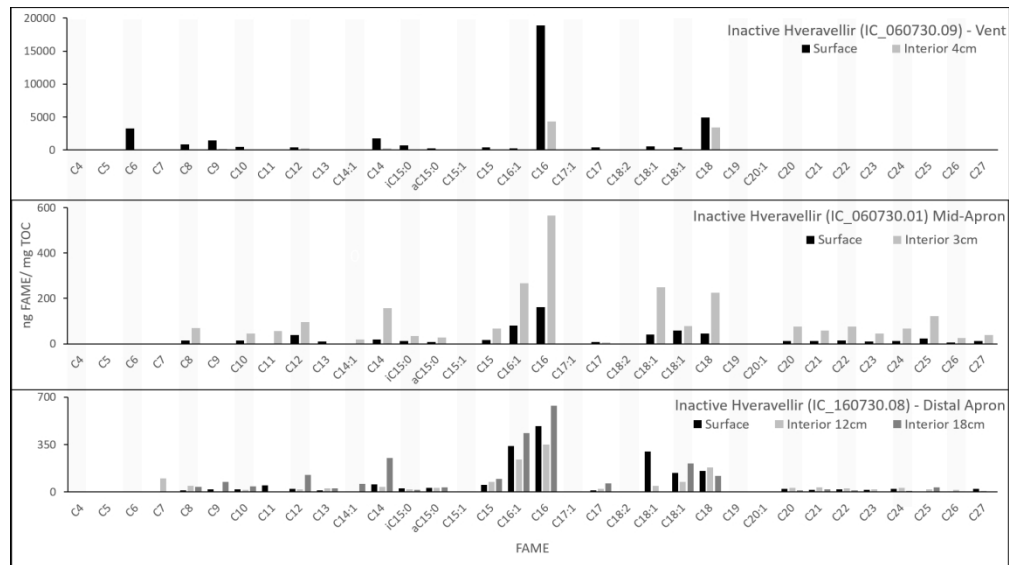


Figure 5. FAME profile for the active Hveravellir Hot Spring system analyzed with a 500 °C flash pyrolysis step.

339x190mm (150 x 150 DPI)

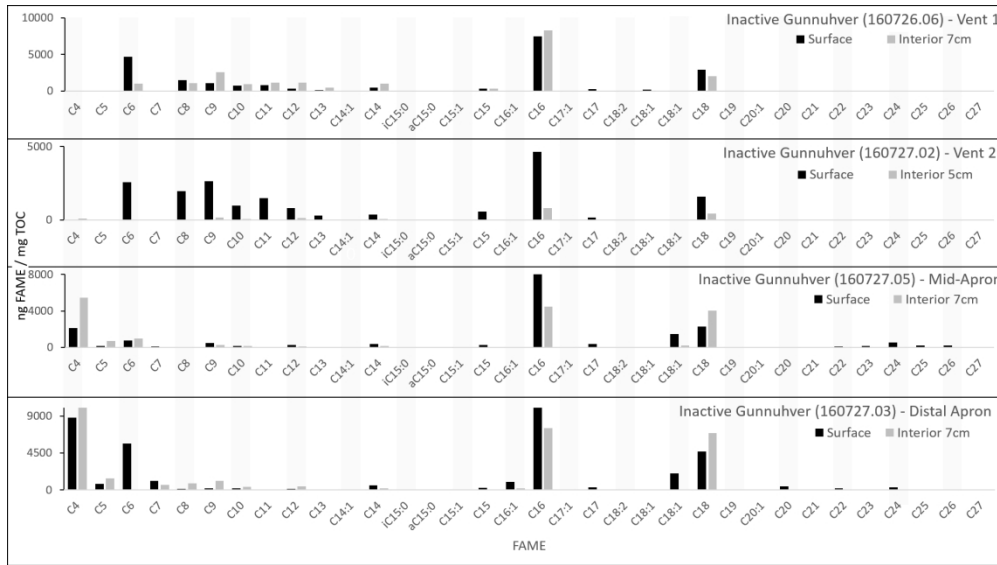


Figure 6. FAME profile for the inactive Gunnuhver Hot Spring system analyzed with a 500 °C flash pyrolysis step.

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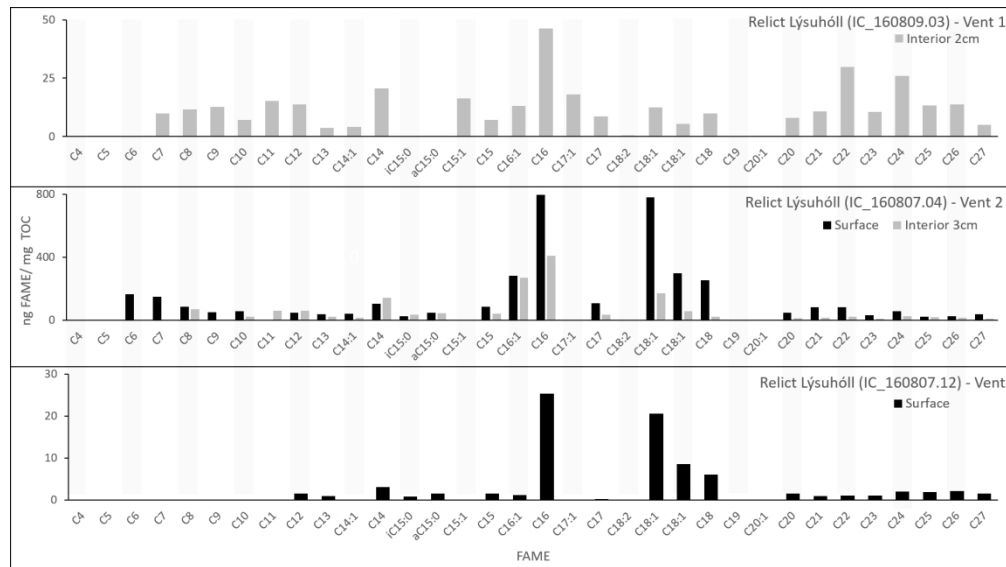


Figure 7. FAME profile for the relict Lýsuhóll Hot Spring system analyzed with a 500 °C flash pyrolysis step.

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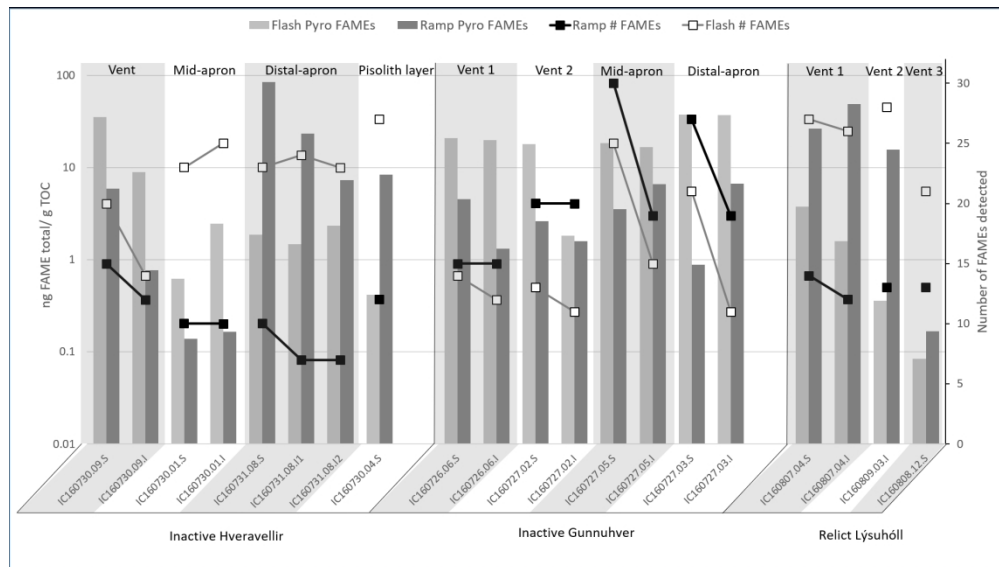


Figure 8. Comparison of total FAME abundance (in ng FAME/ g TOC) detected with the 500°C flash pyrolysis and the SAM-like 35°C/min ramp pyrolysis methods at each sinter sample location. The number of unique FAMES detected with each pyrolysis method is included on the secondary (right) y-axis. Flash pyrolysis nearly always yielded comparable numbers of FAMES relative to ramp pyrolysis.

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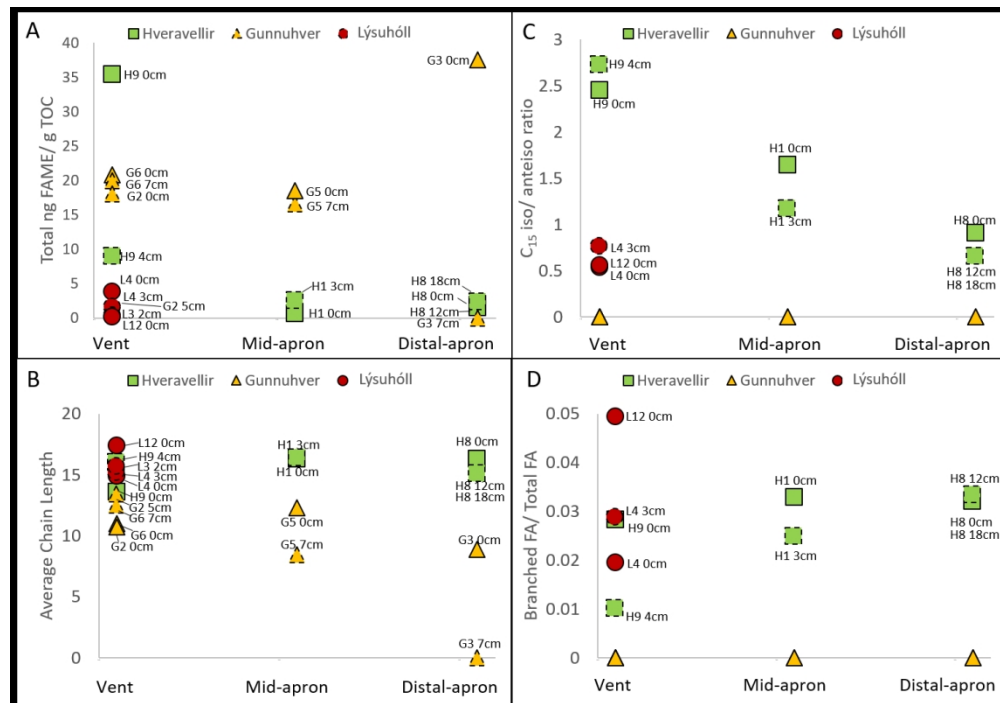


Figure 9. Fatty acid distribution across hot spring facies, with locations “vent”, “mid-apron” and “distal-apron” representing environmental facies as determined by textural characterization. A) Total ng of FAMES / g TOC. The total FAME concentration at the Hveravellir site is highest at the vent, whereas the concentration at the Gunnuhver site is highest at the distal apron. B) The average carbon chain length at the Hveravellir site slightly increases from vent to distal apron, whereas at the Gunnuhver site the average chain length dramatically decreases. C) The ratio of iso- to anteiso-C<sub>15</sub> is highest in the Hveravellir vent site and lowest in the distal apron facies. No br-FA were detected in the Gunnuhver site. The br-FA ratio was low for the Lýsuhóll vent sites. D) The ratio of branched fatty acids to total fatty acids was lowest at the Hveravellir vent site and highest at the distal apron site. No br-FA were detected in the Gunnuhver site. The ratio of br-FA to total FA was higher for the Lýsuhóll vent sites.

256x179mm (150 x 150 DPI)

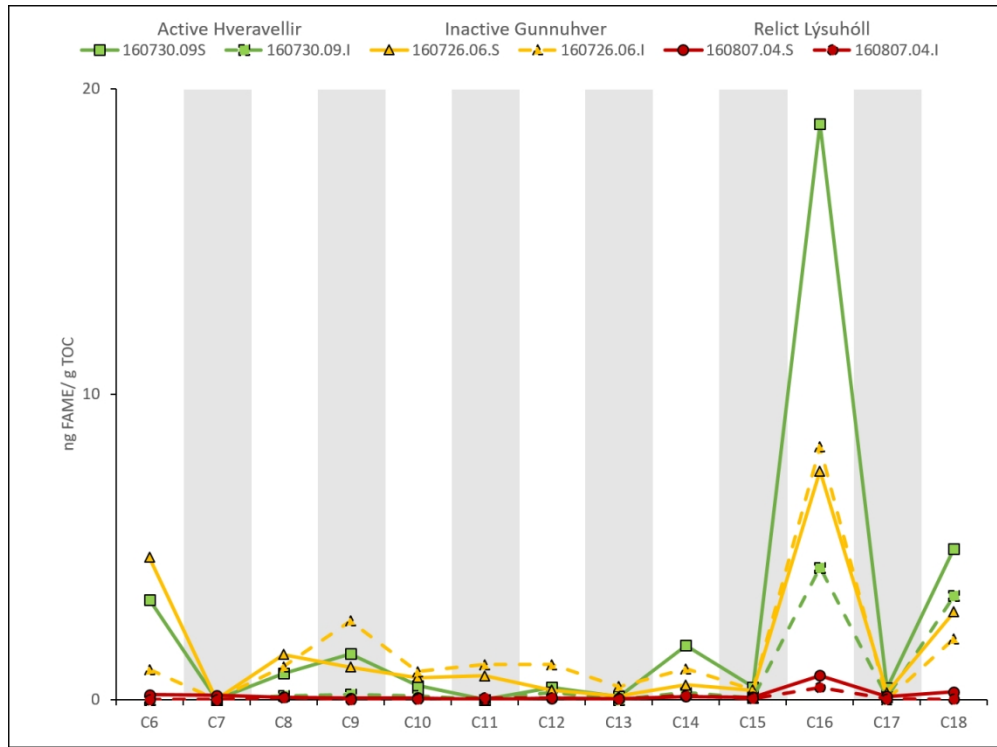


Figure 10. Comparison of FAME concentrations determined with flash pyrolysis experiments from the surface and subsurface vent sites from the Hveravellir, Gunnuhver, and Lýsuhóll spring systems. The even-over-odd carbon number preference is most apparent in the Hveravellir and Gunnuhver locations.

242x180mm (150 x 150 DPI)

Table 1. Sample location, ID, depth, mineralogy, latitude, longitude, percent total carbon and percent total organic carbon.								
Group	Sample ID	Sample	Sample Depth	Mineralogy	Latitude	Longitude	% TC	% TOC ±0.06
Hveravellir Modern Active Spring System	IC160730.09.S	Vent deposit	0 cm	Opal-A, clinopyroxenes (e.g. augite, diopside), plagioclase, gypsum	N64° 51.940'	W19° 33.513'	0.24	0.24
	IC160730.09.I	Vent deposit	4 cm	Opal-A, plagioclase, diopside, magnesite			0.08	0.08
	IC160730.01.S	Mid-apron deposit	0 cm	Amorphous	N64° 51.917'	W19° 33.437'	0.43	0.43
	IC160730.01.I	Mid-apron deposit	3 cm	Amorphous			0.10	0.10
	IC160731.08.S	Distal apron deposit	0 cm	Amorphous	N64° 51.911'	W19° 33.363'	0.08	0.08
	IC160731.08.I1	Distal apron deposit	12 cm	Amorphous			0.13	0.13
	IC160731.08.I2	Distal apron deposit	18 cm	Amorphous			0.15	0.15
	IC160730.04.S	Pisolith layer	0 cm	Opal-A, plagioclase, magnesite	N64° 51.893'	W19° 33.564'	0.30	0.30
Gunnahver Modern Inactive Spring System	IC160726.06.S	Vent deposit	0 cm	Opal-CT	N63° 49.148'	W022° 41.159'	0.02	0.02
	IC160726.06.I	Vent deposit	7 cm	Opal-CT			0.01	0.01
	IC160727.02.S	Vent deposit	0 cm	Opal-CT, halite, quartz	N63° 49.138'	W022° 41.168'	0.13	0.13
	IC160727.02.I	Vent deposit	5 cm	Opal-CT, halite, quartz			0.02	0.02
	IC160727.05.S	Mid-apron deposit	0 cm	Opal-CT	N63° 49.136'	W022° 41.149'	0.17	0.17
	IC160727.05.I	Mid-apron deposit	7 cm	Opal-CT			0.02	0.02
	IC160727.03.S	Distal apron deposit	0 cm	Amorphous	N63° 49.131'	W022° 41.205'	0.15	0.15
	IC160727.03.I	Distal apron deposit	7 cm	Amorphous			0.02	0.02
Lýsuhóll Relict Spring System	IC160807.04.S	Vent deposit	0 cm	Amorphous, calcite	N64° 50.651'	W023° 13.281'	1.83	0.61
	IC160807.04.I	Vent deposit	3 cm	Amorphous, calcite			0.85	0.18
	IC160809.03.I	Vent deposit	2 cm	Amorphous, minor calcite	N64° 50.650'	W023° 13.274'	0.61	0.43
	IC160808.12.S	Vent deposit	0 cm	Calcite	N64° 50.657'	W023° 13.177'	10.38	6.31

Table 1. Sample location, ID, depth, mineralogy, latitude, longitude, percent total carbon and percent total organic carbon.

Table 2. Flash pyrolysis FAME data in ng FAME/ mg TOC.

Group	Sample ID	C4	C5	C6	C7	C8	C9	C10	C11	C12	C13	C14:1	C14	iC15:0	aC15:0	C15:1	C15
Hveravellir Modern Active Spring System	IC160730.09.S	0.0	0.0	3281.4	0.0	862.1	1501.0	467.5	0.0	400.4	110.5	0.0	1780.5	711.3	290.4	0.0	419.3
	IC160730.09.I	0.0	0.0	0.0	0.0	136.2	171.8	127.7	0.0	247.3	0.0	0.0	231.9	66.4	24.4	0.0	68.2
	IC160730.01.S	0.0	0.0	0.0	0.0	13.8	0.0	13.9	0.0	38.9	9.2	0.0	19.5	12.8	7.8	0.0	17.3
	IC160730.01.I	0.0	0.0	0.0	0.0	68.5	0.0	45.0	57.0	95.0	1.6	18.5	157.3	33.1	28.2	0.0	67.4
	IC160731.08.S	0.0	0.0	0.0	0.0	11.0	18.5	18.4	49.8	22.3	14.1	0.0	57.2	28.5	31.4	0.0	53.2
	IC160731.08.I1	0.0	0.0	0.0	100.0	47.0	0.0	16.9	0.0	19.6	26.7	0.0	37.6	19.4	29.7	0.0	75.0
	IC160731.08.I2	0.0	0.0	0.0	0.0	38.7	74.3	41.1	0.0	126.4	25.5	61.8	252.4	14.7	36.0	0.0	98.5
Gunnhver Modern Inactive Spring System	IC160726.06.S	0.0	0.0	4663.7	0.0	1478.2	1070.1	731.4	788.8	319.7	118.2	0.0	485.3	0.0	0.0	0.0	313.2
	IC160726.06.I	0.0	0.0	990.0	0.0	1085.4	2578.5	925.8	1147.4	1148.3	432.7	0.0	1017.4	0.0	0.0	0.0	329.4
	IC160727.02.S	0.0	0.0	2558.7	0.0	1960.6	2630.9	976.6	1494.4	807.0	310.5	0.0	377.5	0.0	0.0	0.0	554.9
	IC160727.02.I	94.7	0.0	0.0	0.0	0.0	163.1	72.9	0.0	120.3	14.8	0.0	59.1	0.0	0.0	0.0	18.9
	IC160727.05.S	2146.2	165.5	770.5	129.1	14.1	467.1	149.3	0.0	247.1	11.8	0.0	387.7	0.0	0.0	0.0	287.5
	IC160727.05.I	5466.7	683.5	1007.4	0.0	0.0	246.3	139.1	52.6	104.9	0.0	0.0	149.8	0.0	0.0	0.0	27.0
	IC160727.03.S	8777.9	715.6	5654.0	1077.1	141.7	181.6	188.0	0.0	123.7	0.0	0.0	554.0	0.0	0.0	0.0	282.1
IC160727.03.I	17486.5	1401.0	0.0	604.7	794.8	1075.5	366.6	0.0	440.7	0.0	0.0	219.2	0.0	0.0	0.0	0.0	
Lýsuhóll Relict Spring System	IC160807.04.S	0.0	0.0	163.4	147.0	84.3	49.0	57.3	0.0	46.9	36.8	39.7	104.8	25.8	48.0	0.0	83.2
	IC160807.04.I	0.0	0.0	0.0	0.0	69.1	0.0	22.3	60.2	58.6	21.8	13.7	142.7	33.8	44.4	0.0	39.5
	IC160809.03.I	0.0	0.0	0.0	9.9	11.5	12.6	7.2	15.3	13.7	3.8	4.0	20.5	0.0	0.0	16.3	7.1
	IC160808.12.S	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0	1.6	1.0	0.0	3.1	0.9	1.5	0.0	1.5

Table 2 con't. Flash pyrolysis FAME data in ng FAME/ mg TOC.

Group	Sample ID	C16:1	C16	C17:1	C17	trans-C18:2	cis-C18:1	trans-C18:1	C18	C20	C21	C22	C23	C24	C25	C26	C27
Hveravellir Modern Active Spring System	IC160730.09.S	244.0	18860.8	5.5	383.7	0.0	580.2	406.4	4953.4	133.6	0.0	10.5	0.0	0.0	0.0	0.0	0.0
	IC160730.09.I	0.0	4334.5	0.0	85.0	0.0	0.0	0.0	3405.9	49.5	0.0	6.2	0.0	0.0	0.0	0.0	0.0
	IC160730.01.S	79.4	160.6	0.0	7.5	0.0	40.1	59.2	44.5	11.4	11.3	13.5	10.8	11.9	22.9	6.0	11.8
	IC160730.01.I	266.7	565.5	0.0	4.5	0.0	249.3	77.2	224.9	75.1	57.8	75.8	46.0	66.5	122.8	24.9	39.5
	IC160731.08.S	340.9	488.0	0.0	13.7	0.0	300.6	141.3	156.8	25.3	15.8	20.5	17.8	24.6	0.0	0.0	25.3
	IC160731.08.I1	239.8	351.6	0.0	22.5	0.0	47.1	73.6	180.1	29.1	35.5	26.2	21.3	29.9	18.1	17.2	9.0
	IC160731.08.I2	433.4	638.5	0.0	64.6	0.0	0.0	211.3	119.3	11.3	20.5	11.1	5.7	9.9	36.1	4.4	0.0
Gunnuhver Modern Inactive Spring System	IC160726.06.S	0.0	7484.0	0.0	237.7	0.0	165.8	0.0	2881.8	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
	IC160726.06.I	0.0	8282.2	0.0	0.0	0.0	0.0	0.0	1993.1	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
	IC160727.02.S	0.0	4655.8	0.0	152.7	0.0	0.0	0.0	1571.3	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
	IC160727.02.I	0.0	820.1	0.0	17.6	0.0	0.0	0.0	439.4	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
	IC160727.05.S	62.2	8073.3	0.0	394.7	0.0	0.0	1469.2	2295.9	76.1	16.0	110.5	143.0	556.9	216.9	212.2	18.4
	IC160727.05.I	0.0	4444.4	0.0	32.1	0.0	0.0	195.7	4021.3	0.0	0.0	0.0	0.0	0.0	0.0	0.0	0.0
	IC160727.03.S	1010.7	10647.0	0.0	296.2	0.0	0.0	1984.8	4710.6	407.2	0.0	223.2	81.4	340.7	0.0	80.4	0.0
Lýsuhóll Relict Spring System	IC160807.04.S	282.3	796.2	0.0	108.5	0.0	780.4	296.8	254.0	46.0	80.2	82.6	29.0	57.2	20.5	22.7	35.4
	IC160807.04.I	268.0	409.4	0.0	33.5	0.5	171.4	54.5	21.4	10.6	14.9	20.7	8.8	22.8	16.7	13.4	9.8
	IC160809.03.I	13.2	46.2	18.1	8.7	0.5	12.5	5.4	9.8	8.1	10.8	29.9	10.5	25.9	13.2	13.8	5.0
	IC160808.12.S	1.2	25.3	0.0	0.3	0.0	20.6	8.6	6.0	1.5	1.0	1.0	1.1	2.0	1.9	2.1	1.5

Table 2. Flash pyrolysis FAME data in ng FAME/ mg TOC. Black cells = concentrations >100 ng/ mg TOC. Dark gray cells = Concentrations 10 to 100 ng/ mg TOC. Light gray cells = 0.01 to 10 ng/ mg TOC. White cells = No detection (0.0 ng / mg TOC).

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Group	Sample ID	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R	
		C <sub>4:0</sub>	C <sub>5:0</sub>	C <sub>6:0</sub>	C <sub>7:0</sub>	C <sub>8:0</sub>	C <sub>9:0</sub>	C <sub>10:0</sub>	C <sub>11:0</sub>	C <sub>12:0</sub>	C <sub>13:0</sub>	C <sub>14:1</sub>	C <sub>14:0</sub>	iC <sub>15:0</sub>	aC <sub>15:0</sub>	C <sub>15:1</sub>	C <sub>15:0</sub>					
Modern Active Spring System at Hveravellir	IC160730.09.S			x	X	x	X	x	X	X	X		x	X	x	X				x	X	
	IC160730.09.I					x	X	x	X	x	X		X	x	X		X				x	X
	IC160730.01.S				X	x	X		x		x		x	X	x	X					x	X
	IC160730.01.I					x		X	x	X	x		x	X	x	X					x	X
	IC160731.08.S					x	X	x	X	x	X	x		x	X	x	X				x	X
	IC160731.08.11				x	x			x		x	X	x	X	x	X					x	X
	IC160731.08.12					x	X	x		x	X		x	X	x	X					x	X
	IC160730.04.S			x	X	x	X	x		x	X			x	X	x	X				x	X
Modern Inactive Spring System at Gunnhver	IC160726.06.S			x		x	X	x	X	x	X	x		x	X						x	X
	IC160726.06.I			x	X	x	X	x	X	x	X	x		x	X						x	X
	IC160727.02.S		X	X	x	X	X	x	X	x	X	x		x	X						x	X
	IC160727.02.I	x	X	X	X	X	X	X	X	X	X	X		x	X						x	X
	IC160727.05.S	x	X	x	X	x	X	x	X	x	X	x		x	X						x	X
	IC160727.05.I	x	X	x	X	x	X	X	X	x	X	x		x	X						x	X
	IC160727.03.S	x	X	x	X	x	X	x	X	x	X		X	x	X		X	X			x	X
	IC160727.03.I	x	X	x		x	X	x	X	x	X			x	X							X
Relict Spring System at Lýsuhóll	IC160807.04.S		X		x	X	x		x	X	x		x	X	x	X				x	X	
	IC160807.04.I				X		x	X	X	X	x	X	x	X	x	X					x	X
	IC160809.03.I				x		x	X	x	X	x	X	x	X					x		x	X
	IC160808.12.S					X			X		x	X	x	X	X	X					x	X

Table 3 con't. Comparison of flash and SAM-like ramp pyrolysis.

Group	Sample ID	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R	F	R				
		C <sub>16:1</sub>	C <sub>16:0</sub>	C <sub>17:1</sub>	C <sub>17:0</sub>	trans-C <sub>18:2</sub>	cis-C <sub>18:1</sub>	trans-C <sub>18:1</sub>	C <sub>18:0</sub>	C <sub>19:0</sub> std	C <sub>20:0</sub>	C <sub>21:0</sub>	C <sub>22:0</sub>	C <sub>23:0</sub>	C <sub>24:0</sub>	C <sub>25:0</sub>	C <sub>26:0</sub>	C <sub>27:0</sub>							
Modern Active Spring System at Hveravellir	IC160730.09.S	x	X	x	X	x		x	X	x	X	x		x											
	IC160730.09.I			x	X					x	X	x	X	x											
	IC160730.01.S	x		x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X		
	IC160730.01.I	x	X	x	X			x		x	X	x	X	x		x	X	x	X	x	X	x	X		
	IC160731.08.S	x		x	X			x		x	X	x	X	x		x	X	x	X	x			x	X	
	IC160731.08.11	x		x	X			x		x	X	x	X	x		x	X	x	X	x			x	X	
	IC160731.08.12	x		x	X			x		x	X	x	X	x		x	X	x	X	x			x	X	
	IC160730.04.S	x	X	x	X	x		x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
Modern Inactive Spring System at Gunnuhver	IC160726.06.S			x	X			x		x	X	x	X	x	X	x	X	x	X	x	X				
	IC160726.06.I			x	X			x	X	x	X	x	X	x	X										
	IC160727.02.S			x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X		
	IC160727.02.I			x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X				
	IC160727.05.S	x		x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
	IC160727.05.I			x	X			x		x	X	x	X	x	X	x	X	x	X	x	X				
	IC160727.03.S	x		x	X	x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
	IC160727.03.I	x	X	x	X			x		x	X	x	X	x	X	x	X	x	X	x	X				
Relict Spring System at Lysuhóll	IC160807.04.S	x	X	x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
	IC160807.04.I	x		x	X			x		x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
	IC160809.03.I	x	X	x	X	x		x		x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X
	IC160808.12.S	x	X	x	X			x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X	x	X



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*Table 3. Comparison of flash pyrolysis and SAM-like ramp pyrolysis of sinter samples. Green = FAME detection in both experiments. Yellow = FAME detection in flash pyrolysis experiment only. Orange = FAME detection in ramp pyrolysis experiment only. Blank cells = no FAME detection. Detection of the C<sub>19</sub> standard is included for completeness.*

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