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Modelling the interplate domain in thermo-mechanical simulations of subduction: Critical effects of resolution and rheology, and consequences on wet mantle melting

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Abstract

The present study aims at better deciphering the different mechanisms involved in the functioning of the subduction interplate. A 2D thermo-mechanical model is used to simulate a subduction channel, made of oceanic crust, free to evolve. Convergence at constant rate is imposed under a 100 km thick upper plate. Pseudo-brittle and non-Newtonian behaviours are modelled. The influence of the strength contrast modelling the subduction channel, represented by the difference in activation energy between crust and mantle (ΔE_a) is investigated to examine in detail the variations in depth of the subduction plane down-dip extent, z_{coup} . First, simulations show that numerical resolution may be responsible for an artificial and significant shallowing of z_{coup} if the weak crustal layer is not correctly discretized. Second, if the age of the subducting plate is 100 Myr, subduction occurs for any strength reduction ΔE_a . The stiffer the crust is, the shallower z_{coup} is (60 km depth if $\Delta E_a = 20$ kJ/mol) and the hotter the fore-arc base is. Conversely, imposing a very weak subduction channel ($\Delta E_a > 135$ kJ/mol) leads to an extreme mantle wedge cooling and inhibits mantle melting in wet conditions. Partial kinematic coupling at the fore-arc base occurs if $\Delta E_a = 145$ kJ/mol. If the incoming plate is 20 Myr old, subduction can occur if the crust is either stiff and denser than the mantle, or weak and buoyant. In the latter case cold crust plumes rise from the subduction channel and ascend through the upper lithosphere, triggering (1) partial kinematic coupling at the fore-arc base, (2) fore-arc lithosphere cooling, and (3) partial or complete hindrance of wet mantle melting. z_{coup} then ranges from 50 to more than 250 km depth and is time-dependent if crust plumes form. Finally, subduction plane dynamics is intimately linked to the regime of subduction-induced corner flow. Two different intervals of ΔE_a are underlined: 80-120 kJ/mol to reproduce the range of slab surface temperature inferred from geothermometry, and 10-40 kJ/mol to reproduce the shallow hot mantle wedge core inferred

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from conditions of last equilibration of near-primary arc magmas and seismic tomographies. Therefore, an extra process controlling mantle wedge dynamics is needed to satisfy simultaneously the aforementioned observations. A mantle viscosity reduction, by a factor 4 to 20, caused by metasomatism in the mantle wedge is proposed. From these results, I conclude that the subduction channel down-dip extent should depend on the subduction setting, in agreement with the observed variability of sub-arc depths of the subducting plate surface. r

Keywords: subduction interface, thermo-mechanical modelling, rheology, numerical resolution, oceanic crust buoyancy, wet mantle partial melting

1. Introduction

The mantle wedge tip and subduction plane are key structures of the subduction factory, affecting not only the subduction viability but also seismogenic properties of the subduction plane, fluid transfers, and melt migration towards the volcanic arc (Fig. 1). Using a thermo-kinematic model applied in 17 subduction zones constrained by heat flux profiles, Wada and Wang (2009) suggested that the Maximum Depth of kinematic Decoupling (MDD) along the subduction plane was constant for all considered subduction zones and equal to 80 km depth, the process controlling it remaining to be determined. Other observations advocate on the contrary for a certain degree of variability, as indicated for instance by the depth to slab under volcanic front ("sub-arc slab depth", Fig. 1) : 110 ± 20 km in Jarrard (1986), varying more broadly in recent studies: 60-120 km depth in England and Wilkins (2004) and 70-170 km in Syracuse and Abers (2006). Even if the average sub-arc slab depth is always centered at 105 km, corresponding to the most frequent overlapping plate thickness (Heuret and Lallemand, 2005), the volcanic arc location may vary as a function of convergence parameters, even if the exact governing process is debated (Grove et al., 2009; England and Katz, 2010). To assess the exact MDD, heat flow data are useful: from a minimum at trench, heat flux increases landwards to a maximum plateau assumed to start right above the MDD (Honda, 1985; Furukawa, 1993). Nonetheless, the assertion by Wada and Wang (2009) of a MDD uniformly equal to 80 km depth is based on heat-flow data presenting either an extremely low spatial density or a significant scattering. Even if the down-dip extent of subduction planes might actually be located at 80 km depth for all subduction zones, the observations used to state it are not completely convincing yet, whereas the variability of well-constrained sub-arc slab depths is meaningful. In any case, the processes regulating the subduction plane geometry have to be better understood.

Notable progress have been made in understanding the subduction plane dynamics. On one hand, van Keken et al. (2002) using thermo-kinematic models showed that modelling a temperature- and stress-dependent viscosity drastically increased thermal gradients at the slab surface. Wada et al. (2008) and Hirauchi and Katayama (2013) showed that the formation of low-temperature hydrated

minerals such as serpentine in the thermal boundary layer covering the subducting slab reduces viscous couplings between slab and asthenosphere and cools the mantle wedge. Thermo-kinematic models have been pointed out by geochemists studying mantle melting at the slab surface (e.g., Kelemen et al., 2003; Plank et al., 2009; Grove et al., 2012; Cooper et al., 2012, Fig. 1), this because simulated slab surfaces are hotter than obtained in thermo-mechanical models, in better agreement with geochemical analyses. The cause likely comes from the prescribed rigid lid within the upper plate that partly includes the underneath low viscosity wedge, which should otherwise flow onto the slab surface (Arcay et al., 2007b; Wada and King, 2015), and cool it. Conder (2005) proposed a parameterization of the subduction fault taking into account temperature- and strain-rate dependence to limit the arbitrariness of the imposed rigid lid thickness. Moreover, partial kinematic coupling has been assumed to occur at the fore-arc base (Kneller et al., 2005, 2007; van Keken et al., 2011), i.e., velocities in the thin layer covering the down-dip extent of the subduction channel would be much lower than corner flow velocity, typically about a few percents of subduction speed (Fig. 1). The origin of partial kinematic coupling being unclear, a possible explanation is proposed in this paper.

On the other hand, thermo-mechanical models showed that if the subduction plane was free to evolve its structure (depth of brittle-ductile transition and MDD) depended on the modelled brittle rheology and on ductile strength, as well as on subduction speed and age at trench of the subducting plate (Arcay et al., 2006; Arcay, 2012). However, the influence of interplate rheology, addressed by Wada et al. (2008) by imposing the MDD, has not been completely established in a self-consistent model, although interactions between interplate shearing and corner flow dynamics may be crucial to regulate the subduction interplate. This motivates the present study. In thermo-mechanical models, the tangential velocity discontinuity along the subduction interface (Fig. 1) is classically modelled by a thin layer weaker than the surrounding mantle by several orders of magnitude (e.g., Kopitzke, 1979; Gurnis and Hager, 1988; King and Hager, 1990; Gerya et al., 2008; Quinquis et al., 2011; Agrusta et al., 2014; Garel et al., 2014; Crameri and Tackley, 2015; Marques and Kaus, 2016). A similar set-up is used in this paper to study carefully how the subduction channel geometry is regulated and how its strength affects the global subduction zone structure. As the subduction plane is likely to equilibrate at the time-scale of tens of Myr, long-term experiments are performed (~ 20 to 40 Myr of convergence). If the upper plate is thick (~ 100 km), forming a viable subduction channel is tough when the incoming plate is young and buoyant (Arcay, 2012). That is why the oceanic crust density will also have to be tested when modelling a young plate subduction. Results show that modifying both the subduction channel strength and density generates a broad variety of subduction regimes. At last, as numerical resolution can be critical (Schmeling et al., 2008), it has to be investigated. These different reasons finally yield a long study but, I hope, consistent. The paper starts with a short summary of key processes implied in the functioning of a subduction interplate. Then, the model set-up is described. Next, the strong influence of numerical resolution is detailed first. Afterwards,

the paper presents what kind of thermal structure is obtained in the vicinity of the subduction plane down-dip extent as a function of rheology for two end-member subducting plate ages at trench (either 100 Myr, section 5.1, or 20 Myr, section 5.2). The case of an middle age incoming plate (60 Myr) is briefly described to derive a comprehensive regime diagram at the end of the result section. Three types of outputs are described: (1) depth of maximum decoupling along the interplate and velocity field in the mantle wedge, (2) subduction viability, and (3) prediction of wet mantle melting. This study shows that the subduction plane structure and extent are intimately linked to the regime of the subduction-induced corner flow in the asthenosphere. In the last part of the paper, by combining modelling outputs with two main observations typically used to discuss the mantle wedge thermal state (pressure-temperature conditions at the slab surface and spatial extent of the cold fore-arc nose), the range of realistic ductile strength to impose along the subduction plane will be discussed. To conclude, the study argues for an interplate decoupling depth dependent on the subduction setting and likely to vary from a subduction zone to another.

2. Subduction interplate functioning inferred from previous studies

To sustain subduction, the interplate channel has to allow simultaneously for two opposite processes (Arcay, 2012). On the one hand, the velocity discontinuity at shallow depth is maintained by a highly localized shearing (Fig. 1). Strain localization depends on the strength contrast between the soft crustal layer and the surrounding mantle. On the other hand, next to the upper plate base, localized strain has to jump from the subduction channel to the asthenosphere to thermally decouple converging plates (Doin and Henry, 2001). This occurs if the subduction channel is stiffer than hot asthenosphere. Reversely, if the subduction channel is not significantly weak, the asthenospheric flow focussing at mantle wedge tip is needed to mechanically separate converging plates. This phenomenon was well exemplified by Eberle et al. (2002). If the subduction fault is modeled solely by a kinematic boundary condition, and not mechanically (such as by low stress or viscosity condition), when a no-slip condition is imposed on top surface, then the asthenosphere rises along the subduction plane up to $z \sim 0$. Indeed, the only means for the flow to satisfy the two kinematic conditions is to propel the weak asthenosphere along the kinematic fault, up to the surface. In this case the asthenosphere acts as the missing decoupling agent. On the contrary, a very weak subduction channel, and/or the absence of corner flow formation prevent localized strain from jumping towards the asthenospheric wedge tip. As a result, the continuous localization of convergence within the weak interplate channel yields horizontal subduction, as modelled for a young and rather buoyant subducting lithosphere (Arcay, 2012).

3. Model set-up

Simulations are performed in 2D by solving the equations of mass, momentum, and energy conservation, using the thermo-chemico-mechanical code of

Christensen (1992), based on spline-finite elements on an Eulerian grid. Fluid incompressibility is assumed except in the momentum equation. A uniform volumetric heating, as well as shear heating and adiabatic heating are included in the energy equation (extended Boussinesq approximation). Compositional tracers, with specific density and rheology, are advected as a function of the velocity field (fourth order Runge-Kutta scheme, van Keken et al., 1997).

3.1. Boundary conditions and initial state

The 2D simulation box, 2220 km wide and 555 km deep, is filled with two different lithologies: a ~ 7 km thick layer overlies the incoming lithosphere, while the rest of the box is made of mantle (Fig. 2). The initial thermal structure of the overriding lithosphere results from a preliminary experiment in which thermal equilibrium between heating from below and cooling from above was performed without plate kinematics, but keeping identical other thermo-mechanical parameters. For simplicity, the upper lithosphere structure is the same for all simulations. A constant convergence rate of 6.5 cm/yr is applied on the top surface far away from the subduction zone (Fig. 2). The crust layer covering the incoming plate surface also fills the initial subduction plane whose dip angle is 30° until 55 km depth. As crustal rocks are assumed to be significantly weaker than mantle (see below), strain will focus within the initial crustal interplate at convergence start (Fig. 2), triggering plate bending in the vicinity of the trench ($x = 1110$ km). The thermal, mechanical and compositional structure of the subduction channel is then free to evolve.

To sustain a constant age at trench of the subducting plate, labelled A_t , a specific thermal condition is imposed in the top left hand corner of the simulation box. A constant plate age is imposed 710 km away ahead of the trench, by setting on a 400 km wide domain a plate cooling from asthenosphere outcropping at $x = 0$ to thermal thickening corresponding to an oceanic plate age, A_{lith} , reached at $x = 400$ km (Fig. 2). As it takes ~ 11 Myr for a column of oceanic plate to drift from $x = 400$ km to the trench, A_{lith} is set to: $A_{lith} = A_t - 11$ (in Myr). Crustal tracers are continuously supplied down to the Moho to renew the layer covering the subducting plate in the top left corner. The Moho depth is either ~ 7 or ~ 8 km depending on resolution mesh (Table 1). Other kinematical, mechanical, and thermal conditions are summarised in Figure 2.

To estimate the upper plate stress state, the regional horizontal tectonic force, F_{res} , is computed at $x = 1665$ km (Fig. 2) from σ_{xx} deviation from lithostatic stresses in an oceanic column of density ρ_{ref} (Christensen, 1992): $F_{res} = -\int_0^{z_i} \sigma_{xx} dz + \int_0^{z_i} \rho_{ref}(z)gz dz$, where z_i is the compensation depth (278 km) and σ_{xx} is the horizontal normal component of stress tensor. F_{res} corresponds to the force applied to keep constant convergence velocity.

3.2. Rheology

A non-Newtonian viscous rheology and a pseudo-brittle rheology are combined. Strength, for both oceanic crust and mantle rocks. The non-Newtonian

ductile viscosity, ν_d , depending on depth (z), temperature (T), and second invariant of the deviatoric strain rate tensor ($\dot{\epsilon}$), writes as:

$$\nu_d = A_0 \exp \left(\frac{E_a(C) + V_a \rho g z}{nRT} \right) \dot{\epsilon}^{1/n-1} \quad (1)$$

where A_0 is a pre-exponential factor, E_a is the activation energy depending on the lithology (C), V_a is the activation volume, P is the lithostatic pressure, n is a exponent higher than 1, ρ is the mantle density, g is the gravity acceleration, and R is the ideal gas constant. In the brittle domain, the yield stress, τ_y , writes as :

$$\tau_y = \tau_0 + \gamma(C) \rho g z \quad (2)$$

where τ_0 is the frictional cohesive strength at the surface, γ is a parameter related to the friction coefficient (Turcotte and Schubert, 1982) depending on lithology. The brittle strain rate, $\dot{\epsilon}_b$, is defined as a function of stress, τ : $\dot{\epsilon}_b = (\tau/\tau_y)^{n_p} \dot{\epsilon}_{ref}$, where n_p is a stress exponent higher than 10 and $\dot{\epsilon}_{ref}$ is reference strain rate. Using the yield stress and the brittle deformation rate, an equivalent viscosity in the brittle domain, ν_b , is defined as:

$$\nu_b = \tau_y \left(\frac{\dot{\epsilon}_b^{\frac{1}{n_p}-1}}{\dot{\epsilon}_{ref}^{\frac{1}{n_p}-1}} \right) \quad (3)$$

The effective viscosity, ν_{eff} , is computed assuming that the total strain rate is the sum of brittle and ductile strain rates: $\nu_{eff} = 1/(\nu_d^{-1} + \nu_b^{-1})$. The crust is supposed to be much weaker than mantle both in the brittle and ductile realms, to promote strain localization within the interplate crustal layer (Fig. 2). The low crustal friction coefficient ($\gamma_c = 0.05$) belongs to the classical range used to simulate realistic subduction dynamics (e.g., Tackley, 2000; Bird, 2003; Gerya et al., 2008; van Heck and Tackley, 2008). Other constants and parameters are listed in Table 2.

3.3. Numerical resolution and investigated parameters

The influence of mesh refinement is investigated by testing three numerical meshes, with an increasing refinement of the smallest mesh. The refined domain is limited to the rectangular area that encompasses the subduction plane and the mantle wedge tip (Table 1) where gradients of temperature, strain and stress are maximum. For the lowest resolution ("LR"), minimum node spacings are 2.85 and 2.27 km in the horizontal (x)- and vertical (z)-directions. Both are decreased to 1.26 km in the most refined mesh ("HR"). A third mesh of intermediate resolution (labelled "MHR") is also tested. In all cases, the tracer density, uniform all over the simulation box, respects a minimum number of 7 markers in the smallest cell, which offers a sufficient accuracy in resolving the lithological difference (Arcay et al., 2005). Therefore, the total number of tracers increases from 1.2544×10^6 for grid LR to 5.76×10^6 for grid HR.

The numerical mesh refinement will be first investigated. The influence of the subducting crust properties is tested next. As mantle characteristics are fixed, I focus on the offset in activation energy between crust and mantle, ΔE_a : $\Delta E_a = E_a^m - E_a^c$, mantle being always stiffer than crust. The impact of ΔE_a on the down-dip extent of the subduction channel will be explored in details for two end-member ages at the trench of the subducting lithosphere (20 Myr and 100 Myr). All other parameters in equation 1 are kept identical. The density difference between crust (ρ_c) and mantle (ρ_m), $\Delta\rho_c (= \rho_c - \rho_m)$, will be tested only for a young (20 Myr old) subducting plate, only moderately negatively buoyant. I will **show how** the pair $(\Delta E_a, \Delta\rho_c)$ modify the regime of young plate subduction. As a thick (100 km) upper plate is an usual feature of subduction zones but also an unfavorable condition for subduction initiation, it is imposed in all simulations.

3.4. Estimate of the interplate down-dip extent

The down-dip extent of the subduction channel may be considered either as a maximum depth of *kinematic* decoupling (Furukawa, 1993; Arcay et al., 2007b; Wada et al., 2008; Wada and Wang, 2009; Arcay, 2012), or as a minimum depth of viscous and kinematic *coupling* between subducting slab and asthenosphere. Since partial kinematic coupling might occur at depth shallower than the asthenosphere (e.g., Wada et al., 2008; Syracuse et al., 2010; van Keken et al., 2011, Fig. 1), it is more convenient to tag the subduction channel down-dip extent as a depth of significant slab-asthenosphere coupling, z_{coup} . To estimate z_{coup} , let us zoom in the interplate structure. At convergence start, as the incipient subducting slab enters the asthenosphere, the corner flow is directed towards the mantle wedge tip before turning downwards, dragged by the subducting slab (Fig. 3a, 2.35 Myr). The following cooling of the incoming plate rises thermal gradients at the mantle wedge tip, which enhances the corner flow and slightly shallows z_{coup} (Fig. 3a, 4.4 Myr). The thermal structure and its associated z_{coup} are stable (Fig. 3a, 25.3 Myr). Along the subduction channel, stress (τ) and strain ($\dot{\epsilon}$), and thence rate of energy dissipation ($\tau \times \dot{\epsilon}$), are very high up to the location where high strain rate jumps into the asthenosphere, where viscosity is low (Fig. 3d). The interplate geometry can hence be tracked by a $x - z$ profile joining maximum dissipation rates modelled at depth shallower than ~ 120 km (Arcay, 2012). Along this profile, the average strain rate, $\dot{\epsilon}$, is close to $2.6 \times 10^{-13} \text{ s}^{-1}$ (Fig. 3d1), i.e., $\sim v_c/h_c$, where v_c is convergence velocity (6.5 cm/yr) and h_c is the weak layer thickness. The interplate geotherm increases roughly linearly with depth until a high temperature increase that marks the proximity to asthenosphere (Fig. 3d2). Therefore, z_{coup} is defined by the depth along the interplate profile where a maximum thermal gradient and a strong strain rate decrease occur. Note that the distribution of high deformation rates, here obtained with a free slip-non deforming top surface and a weak crustal layer, reproduces well the strain rate pattern obtained in convection models including a free surface and a weak crust that simulate self-consistent singled-sided subduction (Schmeling et al., 2008)(Quinquis et al.,

2011; Crameri and Tackley, 2015).

4. Influence of the numerical mesh

A preliminary set of experiments with a ~ 100 Myr old, ~ 100 km thick subducting lithosphere is performed to study how the numerical resolution affects the maximum interplate decoupling depth, z_{coup} (simulations S1-10a to d, Table 3). The physical properties of the crust are kept constant. Using a coarse mesh (labelled LR, x - and z -spacings of 2.8 and 2.3 km, Table 1) yields a shallow z_{coup} , that stabilizes at the location where the crust content at computational nodes is lower than 99.9%. To understand this result, note that the weak layer slightly thins, probably as a result of slab bending and/or internal shearing. Consequently, the weak layer is discretized by less than 4 nodes by the Eulerian mesh at ~ 80 km depth (Figure 4a, top). Close to the crust surface the average crust content at nodes drops from 100% to less than 99.1%. A "contamination" by mantle tracers lower than 1% is sufficient to notably increase the strength averaged at nodes, which triggers the strain rate jump towards asthenosphere at low depth (80 km, Figure 4a, top). With a more refined mesh (HRlim, x - and z -spacings of 1.26 km), the weak crust is discretized by more than 4 Eulerian nodes, among which 3 nodes at least display crust contents close to 100%: this enables the computation of an effective low strength at "crustal" nodes and deepens z_{coup} to the level where the Eulerian mesh is not refined anymore (102 km depth, Figure 4a, bottom). By increasing the downwards extent of the refined area, z_{coup} becomes independent of the numerical mesh (114 km, Figure 4b). A minimum number of 3 nodes to discretise the weak layer core, where crust content equals 100%, appears to be the condition to simulate a mesh-independent coupling depth, as also observed by Crameri et al. (2012). Refining the mesh increases node number but also the number of compositional tracers, to respect a minimum of 7 particles in the smallest mesh (tracer number multiplied by ~ 5 , Table 1). The increase in node number and tracer density between grids LR and HR considerably rises computation time. To find a balance between discretization quality and computation time, a third numerical mesh is tested, with intermediate node spacings in the subduction plane area (MHR mesh, node spacing of ~ 1.7 km). To improve crust discretization, the crustal layer is also slightly thickened (from 6.9 to 8.3 km, Table 1, Figure 4c). The z_{coup} depth obtained with mesh MHR (109 km) is not exactly the one simulated with mesh HR (114 km). However, $\sim 85\%$ of the difference between high and low resolution meshes is accounted for with the intermediate mesh, while mesh MHR allows for saving noticeably computation time. Mesh MHR is used in all simulations presented in the following. Numerical accuracy will be further discussed in section 6.1.

5. Modelling results

5.1. Case 1: Subduction of a cold lithosphere

In simulations S1-2 to S6-15 (Table 3), only the offset in activation energy between crust and mantle, ΔE_a , is varied, from 20 to 155 kJ/mol (E_a^c varying from 445 to 310 kJ/mol). This large range of crust viscous strength enables to test different hypotheses on the lithology forming the subduction channel (sediments, basalt, gabbro, mantle, altered/hydrated phases, composite material). The lowest activation energy (310 kJ/mol, $\Delta E_a = 155$ kJ/mol) mimics a ductile strength encompassed between the ones of a dry quartzite (Ranalli and Murphy, 1987) and of a wet diorite (Ranalli, 1995, Fig. 5a). The strongest rheology ($E_a^c = 445$ kJ/mol, $\Delta E_a = 20$ kJ/mol) simulates the strength of a dry mafic granulite (Wilks and Carter, 1990).

5.1.1. Influence of interplate strength on the corner flow structure

Softening the interplate channel (increasing ΔE_a) increases z_{coup} (Figure 5b). The largest z_{coup} , 163 km depth, is modelled for $\Delta E_a = 135$ kJ/mol. Conversely, the stiffest subducting crust ($\Delta E_a = 20$ kJ/mol) shallows z_{coup} to 60 km depth. This $\Delta E_a - z_{coup}$ trend was already observed by Wada et al. (2008) and described in Arcay (2012), but the values of z_{coup} obtained here are significantly different thanks to the improved resolution. Surprisingly, for the weakest crust ($\Delta E_a = 155$ kJ/mol), the jump of localized strain from the interplate to the asthenosphere does not occur in the numerically refined domain (Fig. 6a2): The crust is so weak that it remains softer than the asthenosphere, which maintains strain localization within the subducting crust. z_{coup} is shifted at the boundary between fine and coarse meshes, at 250 km depth. As the warm corner flow is offset downwards, the fore-arc lithosphere significantly thickens by cooling. The major part of the cold mantle wedge is motionless. For $135 < \Delta E_a < 155$ kJ/mol, an in-between behaviour is modelled. Near the base of the overriding plate (~ 100 km depth), convergence is partly transmitted to the fore-arc base (velocities < 0.5 cm/yr, Fig. 6a1). I label $z_{partial-coup}$ the shallowest depth along the subduction channel where partial kinematic coupling is modelled (Fig. 6a1). It equals 110 ± 10 km in simulation S1-14 ($\Delta E_a = 145$ kJ/mol). Partial kinematic coupling between slab and asthenosphere progressively increases downwards. z_{coup} is defined instead by the coupling depth where rocks overlying the downgoing plate are asthenospheric (hot temperature) and are dragged with velocities close to subduction speed (> 3 cm/yr). According to this definition, z_{coup} for $\Delta E_a = 145$ kJ/mol varies through time but globally deepens, up to 200 km depth (Figure 5b). The z_{coup} dependence on time is likely to be a numerical effect, related to transient thinnings of the weak layer that would possibly disappear by increasing resolution. Nonetheless, the layer of partial kinematic decoupling forms on the whole at the upper plate-asthenosphere boundary, independently of numerical resolution. Note that the ΔE_a range resulting in partial kinematic coupling is quite narrow ($135 < \Delta E_a < 155$ kJ/mol). For a bit younger subducting plate ($A_t = 60$ Myr, simulations S6-2 to S6-15), the dependence of the interplate structure on crust strength is similar to the

one simulated for a 100 Myr old lithosphere (Figure 5b). The main differences are subduction failure if the oceanic crust is too resistant ($\Delta E_a = 20$ kJ/mol) and a general slight z_{coup} deepening compared to $A_t = 100$ Myr. The striking cooling of the fore-arc lithosphere is modelled when $\Delta E_a \geq 145$ kJ/mol, and partial kinematic coupling occurs at the base of the fore-arc since $\Delta E_a = 135$ kJ/mol (Figure 5b). The numerically-induced variations in z_{coup} are however much wider than simulated when $A_t = 100$ Myr.

To summarise, the fore-arc base heating or cooling depends on the corner flow structure, which depends on ΔE_a . The downwelling stream of the corner flow is well developed whatever crust strength, since the asthenosphere sticking to the slab ("viscous blanket", Kincaid and Sacks, 1997, Fig. 3b) does not depend on ΔE_a . On the contrary, the horizontal part of the corner flow is fast and focussed at the upper plate base only if ΔE_a is rather low (< 135 kJ/mol, Fig. 3a), and provides in this case a significant advected heat at the fore-arc base that thins it. The horizontal mantle flow gets more and more diffuse as ΔE_a is risen, which restrains fore-arc basal heating (compare Fig. 6a to Fig. 3a). When $\Delta E_a = 155$ kJ/mol, the heat advected towards mantle wedge tip is so reduced that the fore-arc conductive cooling is unbalanced.

5.1.2. *Effect of a very weak subduction channel on partial melting in the mantle wedge*

To explore the consequences of an extreme cooling of the fore-arc lithosphere when $\Delta E_a \geq 145$ kJ/mol, water transfers associated with slab dehydration and overlying rock hydration are computed in simulations S1-14 and S1-15 (Fig. 6a3). The subducting oceanic crust and the underlying layer of serpentinised mantle (7 km thick) are both assumed to be water-saturated before subduction (containing respectively 2.93 and 6.5% in weight of water). Phase transitions controlling water transfers are predicted by using pressure-temperature paths of lithological tracers, based on accurate phase diagrams determined for a water-saturated peridotite (Schmidt and Poli, 1998) and a gabbroic composition (Bousquet et al., 1997). The model of water expelling, upwards migration, and absorption by the nearest under-saturated marker above the dehydrating element, extensively described and tested in previous papers (Arcay et al., 2005, 2006, 2007b,a, 2008) will not be detailed here. Water transfers are passive and do not affect mantle and crust properties. The key result of the experiment S1-15 is that the mantle wedge hydrated by slab dehydration is too cold to allow for the triggering of mantle melting in wet conditions. Let us assume that the lowest temperature of wet mantle solidus is 980°C (Green, 1973; Millhollen et al., 1974; Schmidt and Poli, 1998). The progressive cooling of the fore-arc lithosphere shifts away from the trench this critical isotherm (Fig. 6a3): at times longer than 16.4 Myr, all the water released by slab dehydration saturates a mantle wedge whose temperature is well below 980°C , as above one slab dehydration front a column of chlorite forms ($T \lesssim 850^\circ\text{C}$, water content of 2.1 %wt, Fig. 6a3, inset). Consequently, the model predicts that the volcanic arc activity of the subduction zone should stop despite a significant fluid flow ex-

pelled from the slab, which is the eventual outcome of a very weak subduction channel. Regarding anhydrous mantle melting, it is all the more impeded as it required temperature hotter than wet partial melting. When $\Delta E_a = 145$ kJ/mol (simulation S1-14), the fore-arc cooling is not as excessive as in the preceeding case: Partial melting within the hydrated mantle wedge is allowed in a limited hot region. For $A_t = 60$ Myr (simulation S6-13), partial kinematic coupling enhances fore-arc cooling through time. The mantle located ontop of the slab dehydration front the farthest from the trench (chlorite out) is predicted to be too cold to enable mantle melting ($T \leq 980^\circ\text{C}$) when convergence duration exceeds 20 Myr. The use of the wet solidus determined by Green (1973) and Millholen et al. (1974) to assess mantle melting occurence in water-saturated condition will be discussed in section 6.

5.2. Case 2: Subduction of a young and warm lithosphere

In this section, experiments simulate the subduction of a young incoming lithosphere, ~ 40 km thick and aged of 20 Myr at trench. For a hot downgoing plate, the efficiency of mechanical decoupling across the subduction interface may be a function not only of the subduction channel strength (ΔE_a) but also of the weak layer density ($\Delta \rho_c$). Indeed, the total slab weight favouring subduction is more sensitive to crust density than in the case of a cold incoming plate, whose mantle density is high and lower the relative effect of crust density. The range of investigated ΔE_a is enlarged to 220 kJ/mol ($E_a^c = 245$ kJ/mol), that simulates the weak strength of a granite (Mackwell et al., 1998, Fig. 5a). The density of the oceanic crust varies between 2920 kg.m^{-3} ($\Delta \rho_c = -380 \text{ kg.m}^{-3}$), which is close to the density of gabbro (2940 kg.m^{-3}) or andesite (2900 kg.m^{-3}) in the greenschist facies (Bousquet et al., 1997), and 3550 kg.m^{-3} ($\Delta \rho_c = +250 \text{ kg.m}^{-3}$), corresponding to the density of an eclogitized gabbro (Bousquet et al., 1997). The mantle density is kept constant (3300 kg.m^{-3} at the surface).

5.2.1. Subduction regime as a funtion of ΔE_a and $\Delta \rho_c$

Figure 7 sums up the broad variety of subduction patterns modelled for $A_t = 20$ Myr depending on the pair ($\Delta E_a, \Delta \rho_c$) (simulations S2-2a to S2-22g, Table 3). Different behaviours may be modelled simultaneously, yielding overlappings domains. Domain 1 may be defined where subduction aborts rapidly (< 10 Myr): the subduction plane is blocked and convergence localises ahead of the trench (white domain in Fig. 7, inset 1). For a stiff crust ($\Delta E_a \leq 40$ kJ/mol), subduction is possible only if crust is denser than the mantle ($\Delta \rho_c > 0$). The softer the crust is, the wider the range of crust density allowing subduction is (yellow domain). However, near the outline of the subduction domain, stresses within the upper plate are highly compressive (light brown domain in Fig. 7, inset 4): subduction would probably stop if convergence was driven by a constant force less than 10^{13} N/m (computed as explained in section 3.1) instead of constant velocity. When the crust is denser than the mantle, early slab break-off occurs if crust is too stiff, followed by subduction obstruction ($\Delta E_a \leq 40$ kJ/mol, left green domain, inset 6). At slightly higher ΔE_a and positive $\Delta \rho_c$,

the incipient subduction zone is unstable, and the slab finally detaches from the surface (purple domain, inset 7). On the contrary, for a weak and dense crust, the early slab break-off does not repeat and subduction is stable ($\Delta E_a \geq 105$ kJ/mol, right green domain, inset 5). The initial slab drip could be related to the weakness of the incipient plate (Andrews and Billen, 2009), and/or to the low activation volume (Cramer et al., 2012).

When the crust is weak and buoyant, small crust plumes initiate near the upper lithosphere-asthenosphere boundary ($\sim 110 \pm 15$ km depth, if the crustal geotherm at this depth is hot enough ($T \sim 1000^\circ\text{C}$, Fig. 6b2; red domain in Fig. 7 inset 2). Two levels of crust plume formation are modelled in a few simulations, the deeper one occurring at ~ 130 km depth (section 5.2.2). Crust plumes have been modeled by Vogt et al. (2013) and Maunder et al. (2016) in the same depth range for similar densities and rheological parameters, but for a much thinner upper plate (~ 50 km) and for slower convergence. In the present study, crust diapirism may be promoted by shear heating in the subduction channel. In the Archean, the underplating of crust plumes along the overlapping plate Moho ("relamination") has been suggested to explain the lower continental crust differentiation, as crust plumes may include partially molten crust (Hacker et al., 2011). Crust plumes do not form for an extremely weak and buoyant crust ($\Delta E_a = 220$ kJ/mol and $\Delta \rho_c = -380 \text{ kg.m}^{-3}$, simulation S2-22a). The subducting crust thickens the subduction channel by accumulating at depth shallower than ~ 100 km and locally convect (velocity of ~ 5 cm/yr, blue domain in Fig. 7, inset 3). The subduction plane thickening, modeled by Maunder et al. (2016) for a similar crustal rheology, was deemed to be a plate tectonics regime in the Archean (Hoffman and Ranalli, 1988).

5.2.2. Partial kinematic coupling and thermal state of the fore-arc lithosphere

As the formation of crustal plumes triggers partial kinematic coupling at the extremity of the subduction channel, the experiments showing crustal diapirism are detailed in this section, illustrated by simulation S2-17b. When instabilities of buoyant crust initiate at ~ 100 km depth, z_{coup} is shifted to 137 km depth while partial kinematic coupling occurs between the plume root and z_{coup} (Fig. 8a, 9 Myr). Soon a secondary branch of asthenospheric corner flow forms reaching the root of the main crust plume (~ 90 km depth, 14.2 Myr). This rises z_{coup} to $z_{partial-coup}$, leading to a classical velocity field with a fast corner flow (> 3 cm/yr, Fig. 8a, 15.2 Myr), before z_{coup} deepens again to ~ 190 km at 18.9 Myr. The crust flow within the main plume (clock-wise current) cools the fore-arc. The crust spinning sense is reversed when z_{coup} rises up to the crust plume head, resulting in a fore-arc warming (Fig. 8, 14.2 Myr). The sequence of fore-arc cooling induced by crust upwelling, followed by a fast warming related to asthenospheric rising, occurs repeatedly. Crust instability thus basically entails (1) a high time-dependence of the fore-arc structure, and (2) an intermittent but intense cooling of the fore-arc domain.

Outside the domain of crust instability, partial kinematic coupling appears when crust buoyancy exceeds -140 kg.m^{-3} at ~ 114 km depth. Figure 8b depicts the

typical evolution simulated in simulation S2-13g: z_{coup} deepens with time (from 175 km to ~ 200 km depth between 8.1 and 10.2 Myr), before being suddenly shallowed when the asthenosphere rises up to $z_{partial-coup}$ (113 km, 12.8 Myr). The z_{coup} deepening followed by quick shallowing is once again a repeating process (Fig. 8).

5.2.3. z_{coup} dependence on ΔE_a , $\Delta \rho_c$, and time

Figure 9 displays z_{coup} variations with crust strength and buoyancy after ~ 10 Myr of convergence (panel a). z_{coup} is mainly constant through time in the absence of crust plume, but becomes strongly time-dependent otherwise. The transient state resulting from fast z_{coup} shallowings up to the depth of $z_{partial-coup}$ is captured in panel b, where the shallowest z_{coup} value simulated between 10 and 13.5 Myr of convergence is mapped. First, the large range of z_{coup} modelled for $A_t = 20$ Myr varies from 50 km to 250 km depth. This maximum depth modelled at 13.5 Myr in simulation S2-17c ($\Delta E_a = 175$ kJ/mol; $\Delta \rho_c = -240$ kg.m $^{-3}$) is controlled by the mesh coarsening. Second, in the domain where z_{coup} is stable (for $\Delta \rho_c \gtrsim -120$ kg.m $^{-3}$), the coupling depth deepens when ΔE_a increases, as already described for $A_t = 100$ Myr (section 5.1.1). At constant crust strength, z_{coup} slightly deepens with increasing crust buoyancy, possibly because of a decrease in slab dip angle (compare modelled slab dips in insets 2, 4 and 5 in Figure 7), though the relationship between slab dip angle and crust buoyancy may not be so straightforward (compare insets 5 and 3). Third, when z_{coup} is time-dependent, crust plumes and partial kinematic coupling strongly increase z_{coup} (> 150 km), this deepening being partly counterbalanced by punctual upwards z_{coup} jumps that temporarily reverse the $z_{coup} - \Delta E_a$ trend observed for $\Delta \rho_c > -120$ kg.m $^{-3}$ (solid black squares in Fig. 9b).

5.2.4. Wet mantle melting as a function of z_{coup} evolution

Water transfers are computed in 10 experiments of young plate subduction to evaluate the likelihood of wet mantle melting (purple symbols in Fig. 9a). As mantle melting is not directly computed, I focus on the water flux that should trigger melting. To trigger melting at the slab surface, the water flux expelled by slab dehydration must be not only significant but also absorbed at temperature hotter than 980°C (lowest wet solidus for peridotite, section 5.1.2). In the present modelling this water is first dissolved in the asthenosphere and fore-arc base (water content at saturation < 0.05 wt%, Arcay et al., 2005). If there is some left, the remaining water hydrates the cold overlapping lithosphere where nominally hydrated phases form (amphibole, talc, serpentine), increasing water contents above 0.5 wt%. Water contents in the upper plate, that here acts as a "sponge", are thus tracked to estimate the water amount available at slab surface at $T > 980^\circ\text{C}$ that would normally yield mantle melting (Fig. 10). Three melting regimes are modelled. If fore-arc cooling is limited ($z_{coup} \lesssim 140$ km), one part of dehydration water should always be absorbed at $T > 980^\circ\text{C}$ in amounts greater than 0.5 wt%, which might be sufficient for sustaining the volcanic front

activity even if the corresponding metasomatized domain is narrow (Fig. 10a, solid triangles in Fig. 9). If fore-arc cooling is intense (z_{coup} always ≥ 160 km), almost all the released water is absorbed at temperature below wet solidus (Fig. 10c, crosses in Fig. 9). Therefore, the volcanic front activity would probably vanish. An intermediate regime occurs if mantle wedge cooling is disturbed by sudden z_{coup} shallownings. Compared to the previous situation, transient fore-arc warmings extend mantle melting for a while (Fig. 10b, empty triangles in Fig. 9). Magma supply and volcanic front activity might be intermittent in this case. In conclusion, the mantle wedge cooling resulting from the subduction channel weakness significantly restricts wet mantle melting for $A_t = 20$ Myr, as already described for older plate subduction.

5.3. Synthesis

5.3.1. Decoupling along the subduction interface as a function of slab age, crust buoyancy, and interplate rheology

Figure 11 schematizes how crust strength and slab weight regulate kinematic decoupling between converging plates in this modelling. As already mentioned, the weight of the downgoing slab depends here on crust density and on lithosphere age at trench (Turcotte and Schubert, 1982). If $A_t \geq 60$ Myr, the average slab density exceeds the asthenospheric one whatever crust buoyancy. The resulting slab weight excess promotes high dip angle ($> 30^\circ$) when the incipient slab enters the asthenosphere. This activates the corner flow that eases kinematic decoupling at the interplate down-dip extent and, finally, subduction. This process is only slightly sensitive to the subduction channel strength (purple domain in Fig. 11). Nonetheless, the latter controls the structure of the horizontal branch of the corner flow: fast, focussed and warming the fore-arc base when ΔE_a is low (red area in Fig. 12), and inversely slow, diffuse and advecting a low heat flux at the mantle wedge tip when ΔE_a is high (blue area). On the contrary, if $A_t = 20$ Myr, as mantle slab density is minimum, crust buoyancy has a significant influence on the total slab weight excess. If $\Delta \rho_c = -380 \text{ kg.m}^{-3}$, the low dip angle of the incipient slab hampers corner flow activation. Subduction initiates only if the interplate decoupling is very efficient up to asthenospheric levels, i.e., if crust is quite weak (yellow domain in Fig. 11). If the subducted crust is not so buoyant ($\Delta \rho_c \geq 0$), the increase in slab dip angle helps the corner flow triggering and lowers the minimum ΔE_a enabling subduction.

5.3.2. Fore-arc structure as a function of subducting plate age and crust weakness

Figure 12 sums up the major characteristics of the fore-arc structure modelled as a function of incoming plate age, A_t , and crust weakness, ΔE_a , for a crust buoyancy set to -380 kg.m^{-3} . To begin with, along the boundary delimiting subduction, the depth of slab-asthenosphere coupling is the shallowest that can be modelled for a given A_t , since ΔE_a is the lowest (Fig. 5b). Therefore, the fore-arc lithosphere base is predicted to be there the warmest (red area).

This is nonetheless not true for a 20 Myr old incoming lithosphere subducting only for very weak crust, which leads to fore-arc cooling and cold crust diapirism (green area). Fore-arc cooling by z_{coup} deepening occurs if $\Delta E_a \geq 155$ kJ/mol whatever subducting plate age, and can be intense enough to prevent mantle melting in wet conditions (blue area). Lastly, partial kinematic coupling at fore-arc base is modelled in a narrow range of crust strength ($135 < \Delta E_a < 155$ kJ/mol if $A_t = 100$ Myr), and moderately cools the fore-arc. The depth where partial kinematic coupling starts, $z_{partial-coup}$, **is mainly a function of** the upper plate thickness. As partial kinematic coupling is systematically associated with Rayleigh-Taylor crustal instabilities when $A_t = 20$ Myr, the domain of partial kinematic coupling for $A_t \leq 60$ Myr is assumed to merge with the domain of cold crust plumes. To conclude, no universal subduction plane strength is able to simulate alone realistic subduction, whatever subducting plate age, if the crust is significantly buoyant and the overlapping plate thick (~ 100 km).

6. Discussion and concluding remarks

6.1. Study limitations

Two main shortcomings can be underlined: (1) the numerical scheme used to interpolate strength from lithological markers' properties, and (2) modelling the subduction channel by a rheological contrast constant with depth. The viscosity at Eulerian nodes can be computed from markers' values (1) using an arithmetic averaging, assuming deformation at constant strain rate (such as pure shear parallel to compositional layering, Schmeling et al., 2008). On the contrary, a harmonic averaging simulates simple shear (at constant stress) parallel to the channel surface. While arithmetic averaging yields a numerical strengthening, the harmonic one reproduces well a thin channel low viscosity, and is more adapted to simulate a subduction channel (Schmeling et al., 2008; Deubelbeiss and Kaus, 2008). As both kinds of shear are expected in a subducting slab, the averaging scheme should change depending on deformation pattern, which is not easy to handle numerically (Schmeling et al., 2008). Geometric averaging leads to a in-between averaged strength. In the present study, brittle and ductile rheologies are combined through harmonic averaging (section 3.2) while viscosity at nodes is estimated from compositional markers by an arithmetic averaging. The subduction channel strengthening simulated for a coarse mesh (section 4) directly results from arithmetic averaging. To better model viscosity layering associated with lithological boundaries, resolution may be improved by either (i) increasing the number of Eulerian nodes or thickening the low viscosity layer when arithmetic averaging is used (as in this study), (ii) decreasing the interpolation volume around Eulerian nodes (Duretz et al., 2011), (iii) limiting irregularities in marker distribution by implementing a conservative velocity interpolation scheme (Wang et al., 2015), or (iv) modeling compositional interface during marker advection, through for instance the particle level set method (Samuel and Evonuk, 2010).

Imposing a ductile strength contrast ΔE_a constant with depth along the subduction channel (2) supposes that lithological variations are small. It is more likely that the medium composing the interplate changes with depth, from, possibly, sediments at shallow depth, upper then lower crusts, to more mafic composition at great depth. This lithological sequence could be simply modelled by a strength increase along the subduction channel, i.e., by a ΔE_a decrease with depth (Fig. 5a). As low ΔE_a yields z_{coup} shallowing (Fig. 5b and 9), a ΔE_a lowering with depth should shallow z_{coup} . If softening minerals locally form, such as serpentine at high depth (Duarte et al., 2015), a localized ΔE_a increase might favor z_{coup} deepening, but this would have to remain limited, otherwise volcanic arc extinction by mantle wedge cooling might occur (Fig. 12).

6.2. Partial kinematic coupling

Partial kinematic coupling next to the mantle wedge tip would explain the extent of the "cold fore-arc nose" detected in seismic tomographies (e.g., Zhao et al., 1992; Zhao, 2001; Stachnik et al., 2004; Abers et al., 2006, Fig. 3). Partial kinematic coupling would result from ductile strain localized within weak materials, such as sediments or hydrated phases (serpentine, talc, Peacock and Hyndman, 1999), or brittle deformation occurring deeper than the brittle-ductile transition. In thermo-kinematic models, partial kinematic coupling is modelled by imposing a given percent of kinematic coupling over a specified length at the subduction plane extremity. Coupling abruptly passes from partial to full to mimic the effect of a dehydration reaction on viscous coupling (Kneller et al., 2005, 2007; Syracuse et al., 2010). A low coupling (2-3%) between 40 and 70 km depth cools the fore-arc while sustaining a sharp sub-vertical boundary between low-temperature fore-arc and hot asthenospheric wedge, consistent with seismic cross-sections and heat flow profiles. An increased coupling ($\sim 10\%$) leads to a cooling of the subduction plane down-dip extent while a "hot tongue" of asthenosphere rises up to the location where full decoupling between converging plates is imposed (~ 40 km depth). This odd thermal structure is actually not consistent with seismic tomographies. The present study shows that partial kinematic coupling may form according to the incoming plate age and the subduction plane strength (135 -145 kJ/mol if $A_t = 100$ Myr, enlarged to $\Delta E_a > 120$ kJ/mol if $A_t = 20$ Myr, Fig. 7). The transition to full kinematic coupling is then progressive (from 0 to $\sim 7\%$) and occurs on a large distance (> 100 km extent), while it is pinpointed and thus very sharp in thermo-kinematic models. The thermal structure modelled at slab surface is then much more smoothed than in thermo-kinematic modellings.

6.3. Diapirism-induced fore-arc cooling, numerical resolution, and mantle melting prediction

In this modelling, z_{coup} is mesh-independent if the weak layer core is discretized by more than 3 computational nodes. Otherwise, the artificial strengthening of the subduction channel triggers a hot mantle upwelling to soften the interplate, ending in a numerical shallowing of z_{coup} . When crust plumes are

modelled, they reduce the crustal channel thickness. As a consequence, the asthenospheric corner flow frequently rises up to the root of crust instabilities. The sporadic upwards jumps of z_{coup} are thus probably a numerical effect that could be limited using a more refined mesh. The thermal state of the fore-arc would then be even colder and wet mantle melting even more hindered than simulated for $A_t = 20$ Myr. This is consistent with the general fore-arc cooling modelled for older incoming plates at high crust strength reduction.

6.4. *Can nominally hydrated phases be stable at the slab surface?*

Two main layers of hydrous phases are assumed to release the water responsible for partial melting in the mantle wedge. The deeper one is located within the subducting slab, inside the oceanic crust and/or the underlying serpentinized mantle. The second one would be the viscous blanket covering the downgoing slab (Fig. 3b), hydrated by the underlying dehydrating layers. Amphibole and serpentine have been assumed to be stable at the slab surface (Iwamori et al., 2007; Wada et al., 2008; Hirauchi and Katayama, 2013), but chlorite would be the hydrated phase stable at the hottest temperature at high depth ($\sim 860^\circ\text{C}$ at 60-100 km depth, Schmidt and Poli, 1998; Grove et al., 2004, 2009; Till et al., 2012). Chlorite destabilization would provide the water triggering melting at the bottom of the mantle wedge (Till et al., 2012, see section 6.5). Still, chlorite stability was assessed based on thermal models performed with a temperature-dependent Newtonian viscosity. If the stress-dependence of olivine rheology is included, the increase in thermal gradient at the slab surface (van Keken et al., 2002; Syracuse et al., 2010) prevents the formation of stable chlorite. Similarly, the simulations performed in this study predict that chlorite cannot form at the base of the mantle wedge if $\Delta E_a < 135$ kJ/mol (Fig. 13). Chlorite formation at high depth is predicted when the simulated subduction zones are the coldest, notably when z_{coup} is extremely deep (> 160 km depth, Fig. 6a3 and 10a,b). However, since hydrous minerals may be softening materials, their formation would lower the asthenosphere viscous drag at the slab surface, emphasizing wedge cooling and melting hindrance. Finally, as the crust mainly dehydrates before reaching the asthenosphere (Fig. 6a3), this study suggests that the deserpentinization of the subducting mantle may be the major source of the water triggering melting in the mantle wedge.

6.5. *Solidus assumed to assess wet mantle melting*

When the subduction channel is particularly weak ($\Delta E_a \geq 135$ kJ/mol, or a bit more depending on $\Delta\rho_c$ if $A_t = 20$ Myr), the water released by slab dehydration is predicted to be absorbed in the mantle wedge at temperature below the wet solidus, switching off the volcanic front activity. The model assumes the minimum temperature of the wet solidus to be 980°C , as estimated by Kushiro et al. (1968) and Green et al. (2010). However, other experiments performed on longer durations (Mysen and Boettcher, 1975; Grove et al., 2004; Till et al., 2012), and/or based on peridotite samples less depleted in iron (Grove et al., 2012) suggest a lower wet mantle solidus, related to the chlorite stability field:

810±10°C between ~2 and 4 GPa. Therefore, the inhibition of mantle melting induced by an extreme z_{coup} deepening might not occur if the relatively cold wet solidus of Till et al. (2012) was taken into account. Nonetheless, the strong fore-arc cooling (modelled for instance in simulation S1-15, Fig. 6a3) is artificially limited by the numerical mesh coarsening imposed at 250 km depth. The fore-arc would thus be even more cooled if mesh coarsening was shifted to the base of the simulation box. Finally, I assume that, at least on long-time scales (>30 Myr), the fore-arc conductive cooling would prevent mantle partial melting whatever the considered wet solidus.

6.6. Slab surface temperature

To discuss the range of realistic subduction channel strengths, let us consider the thermal state modelled along the subducting slab surface, defined as the top of the downgoing oceanic crust, for $A_t = 100$ Myr. (Fig 13d). Slab surface temperature (SST) are compared with pressure-temperature ($P - T$) conditions estimated from geothermometry derived from H₂O/Ce measurements on melt inclusions (Plank et al., 2009; Cooper et al., 2012). Data are compiled from very different subduction zones, such as slow subduction of young plates (e.g., Cascadia) or fast subduction of old plates (NE Japan or Tonga). They cannot not be directly compared with modelled SST, but rather considered as upper and lower boundaries. Modelling an overlapping plate thinner than simulated would also probably shift SST to lower depth. The modelling subduction setup simulates the one prevailing in the Kamchatka (Heuret and Lallemand, 2005). Imposing a stiff oceanic crust ($\Delta E_a = 20$ kJ/mol, Fig 13a) yields a slab surface much hotter than observed, while the strong fore-arc cooling associated with a weak subduction channel ($\Delta E_a \geq 135$ kJ/mol) produces too cold SST (Fig 13b and d). Imposing $\Delta E_a = 80$ kJ/mol reproduces the hotter observed SST, whereas SST obtained with $\Delta E_a = 105$ kJ/mol correctly simulate the $P - T$ conditions observed in the Kamchatka. We should expect the modelled SST to be colder than estimated along several young and/or slow subducting plates (South Chile, Nicaragua or Lesser Antilles), but it is the inverse that is simulated. This would argue for increasing ΔE_a to ~120 kJ/mol to moderate SST increase. I conclude that the ΔE_a range that mimics best worldwide SST estimates from thermometry is 80-120 kJ/mol.

Moreover, the shallowest and hottest nose in the mantle wedge may be constrained by conditions of last equilibration with the mantle of near-primary arc magmas (Kelemen et al., 2003, Fig. 13a), considered as the signature of hydrated melts reaching the upper plate Moho. Still, hot conditions at the arc Moho may derive from the heat advected by melt migration, which is here neglected. The hottest mantle wedge is modelled for the stiffest subduction channel, but the hot nose is not as shallow as observed ($\Delta E_a = 20$ kJ/mol, lines with stars in panel 13d). To reproduce observations, an additional ingredient must be included to enhance the corner flow. In simulation S1-8b, first, water transfers associated with slab dehydration/mantle wedge hydration are simulated (section 5.1.2); second, an hydrous strength reduction whose amplitude increases with water content $[\text{OH}^-]$ is imposed for the hydrated mantle: low

reducing factor (increasing from 5 to 32) when water is dissolved in nominally anhydrous minerals ($200 \text{ ppm} \leq [\text{OH}^-] \leq 0.2 \text{ wt\%}$), high and maximum (set to 350) when hydrous minerals form (e.g., amphibole, $[\text{OH}^-] \geq 0.5 \text{ wt\%}$, Arcay et al., 2005, Fig. 13c). ΔE_a is set to 80 kJ/mol. Viscosity is then low in the hydrated mantle wedge ($\sim 10^{18} \text{ Pa.s}$ at $\sim 50 \text{ km}$ depth, Fig 13c), which shallows z_{coup} to $\sim 48 \text{ km}$ and simulates the shallow and hot mantle nose inferred from near-primary arc magmas (1250°C at $\sim 50 \text{ km}$ depth, Fig. 13d). The drawback of this water-related weakening is that Rayleigh-Taylor instabilities initiate at the base of the hydrated upper plate (Arcay et al., 2005, Fig 13c), thickening the viscous blanket. Compared to simulation S1-8 without water transfers, SST is cooled by $\sim 150^\circ\text{C}$, mimicking the coldest $P - T$ conditions inferred from $\text{H}_2\text{O/Ce}$ measurements. As the convective destabilization of the overriding lithosphere is impeded if the maximum hydrous strength reduction is lower than $20^{3/2} \simeq 90$ (Arcay et al., 2006), a compromise could be to limit the maximum water-related weakening to 90, instead of 350. The asthenospheric viscosity would then be lowered by a factor varying from 4 to 19.

6.7. *Interplate strength favouring subduction viability and reproducing cold fore-arc nose extent*

This study shows that there is no universal value of the subduction channel strength able to model realistic subduction whatever subducting plate age (i.e., successful initiation, limited compressive stresses and occurrence of wet mantle melting), if crust buoyancy is -380 kg.m^{-3} (Fig. 12). Since the best ΔE_a interval modelling reasonable SST is 80-120 kJ/mol (section 6.6), crust buoyancy should not exceed $-80/-200 \text{ kg.m}^{-3}$ to allow young plate subduction (Fig. 7). Nevertheless, the simulation of a hot and shallow mantle wedge consistent with conditions of last equilibration of near-primary arc magmas requires a quite different range of crust strength in anhydrous conditions: $\Delta E_a \leq 20 \text{ kJ/mol}$. Moreover, the sharp transition between low to high seismic velocities delineating the cold fore-arc nose is often observed at shallow depth: $\sim 65 \text{ km}$ in NE Japan and Alaska (Zhao et al., 1992; Stachnik et al., 2004; Abers et al., 2006); $\sim 45 \text{ km}$ in Hikurangi (Davey and Ristau, 2011) and in the Cascadia (Zhao, 2001) and $< 40 \text{ km}$ in S Japan (Miura et al., 2003). If seismic anomalies image thermal contrasts, to model such shallow z_{coup} when $A_t \simeq 100 \text{ Myr}$, ΔE_a should be set to $\sim 30\text{-}40 \text{ kJ/mol}$ in Alaska and NE Japan, and lower than 20 kJ/mol in Hikurangi (Fig. 5b); in the Cascadia and S Japan where $A_t < 25 \text{ Myr}$, ΔE_a should be $20\text{-}40 \text{ kJ/mol}$ and at most set to 20 kJ/mol , respectively (Fig. 9). The constraints from seismic tomography lead to $10 \lesssim \Delta E_a < 40 \text{ kJ/mol}$, which is consistent with the range inferred from hot mantle wedge-derived melts. Finally, two different ranges of ΔE_a may be underlined: 80-120 kJ/mol and 10-40 kJ/mol, that are not compatible although constraints from petro-geochemistry and geophysics were obtained on the same subduction zones. How to infer a reconciling value? Again, an extra process regulating the mantle wedge structure could be invoked, such as a decrease in mantle strength depending on $[\text{OH}^-]$ (section 6.6). A moderate weakening ($\sim$ one order of magnitude) in the hydrated

wedge should be sufficient to simulate realistic SSTs and a shallow hot wedge core, keeping for the interplate strength the median value $\Delta E_a = 80$ kJ/mol (simulation S1-8b, Fig. 13c-d).

6.8. Is the down-dip extent of the subduction plane likely to be identical for worldwide subduction zones?

In conclusion, three sets of rheological parameters are pointed out: (1) $\Delta E_a \sim 80$ -120 kJ/mol to simulate SST derived from H₂O/Ce measurements; (2) $\Delta E_a \sim 10$ -40 kJ/mol to model the hot and shallow mantle wedge inferred from seismic tomographies and near-primary arc magmas; and (3) a median crust strength ($\Delta E_a \sim 80$ kJ/mol) combined to a water-related weakening. Only rheological sets (1) and (3) are compatible. In case (2), z_{coup} is strongly controlled by the corner flow. As the corner flow dynamics basically depends on the incoming slab characteristics and subduction parameters (such as convergence rate), if $\Delta E_a = 25 \pm 15$ kJ/mol then the subduction plane down-dip extent very probably varies from a subduction zone to another. In case (1), z_{coup} equals more or less the upper plate thickness: a bit shallower if $\Delta E_a = 80$ kJ/mol (82 km if $A_t = 100$ Myr), or a bit deeper if $\Delta E_a = 120$ kJ/mol (135 km, Fig. 5b). Hence, the down-dip extent of the subduction channel may be almost independent from subduction parameters if $\Delta E_a = 120$ kJ/mol, and moderately dependent if $\Delta E_a = 80$ kJ/mol. In case (3), z_{coup} strongly depends on subduction velocity (as shown for $\Delta E_a = 110$ kJ/mol by Arcay et al., 2008, using a mesh too coarse implying crust strengthening: the effective ΔE_a might have been ~ 80 kJ/mol). Since rheology (3) enables to verify all the listed constraints on the mantle wedge thermal state, it may be the best choice to simulate the subduction channel strength. This implies to limit crust buoyancy to $\Delta \rho_c \sim 80$ kg.m⁻³ to allow young plate subduction (Fig. 7), assuming crust eclogitization at ~ 17 kbar (Bousquet et al., 1997). Besides, if $\Delta E_a \sim 80$ kJ/mol, then partial kinematic coupling at the fore-arc base should not occur. Lastly, the rheological properties of the subduction channel estimated in this study imply that the subduction plane down-dip extent varies as a function of the subduction set-up, as suggested by the wide range of slab sub-arc depths (England et al., 2004; Syracuse and Abers, 2006, section 1). This conclusion disagrees with a single value of z_{coup} universally equal to 80 km depth (Wada and Wang, 2009).

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Figure 1: Sketch of the subduction interplate and definitions of the main terms used to describe the mantle wedge structure (in brown). Green ascending arrows represent fluids expelled by the dehydrating slab. The red area depicts the mantle wedge part hot enough for melting in wet conditions. In the bottom panel, two kinds of kinematic coupling between the subducting slab and the overlying mantle are considered: In case of full coupling, velocities at the slab surface are close to the subduction speed as soon as the down-dip extent of the subduction channel is crossed, defining z_{coup} there. In case of partial kinematic coupling, velocities at the slab surface are low at the subduction channel end and increase downwards. $z_{partial-coup}$ is defined as the shallowest depth where velocities at slab surface are low but different from zero, while z_{coup} is shifted deeper at the level where overlying mantle velocities reach the subduction speed.

Figure 3: Close-up on the subduction structure obtained in simulation S1-8 (Table 3). (a) Thermal state (one white isotherm every 200°C), composition (oceanic crust outlined by the green contour) and velocity field at different time steps. The trench is located at the abscissa $x = 0\text{ km}$. (b) Thermal structure and strain rate distribution at steady-state (26 Myr of convergence). One red outline every $8 \times 10^{-14}\text{ s}^{-1}$. (c): Close-up on the subduction interface. The interplate sampling line (orange line) joins points of maximum dissipation energy rates (see the text for details). Outlines of dissipation energy rate (purple lines) are depicted every $20\text{ }\mu\text{W.m}^{-3}$. Panel(d): Profiles interpolated along the interplate sampling line of, from left to right, (1) second invariant of strain rate tensor, (2) temperature, and (3) viscosity.

a. Low resolution close to the coupling depth

b. High resolution

c. Medium resolution

Figure 4: Influence of grid resolution and weak layer thickness on the down-dip extent of the decoupling interface: Zooms on the subduction structure at steady-state. Isotherms (black lines) every 200°C. The rectangular Eulerian mesh is coloured as a function of crust content at computational nodes, by highlighting very high contents in crust material (>99.0 %). Within the white line the content in weak oceanic crust at Eulerian nodes is greater than 49.0 %. Panel (a): Simulation S1-10a (Table 3), performed with the coarsest mesh LR (Table 1) and simulation S1-10b performed with the finest mesh HRLim. Panel (b) Simulation S1-10c, computed with the finest mesh HR. Panel (c): Simulation S1-10d, performed with mesh MHR of intermediate resolution.

Figure 5: (a) Simulated interplate strength for different values of ΔE_a . Stress profiles are computed with the interplate geotherm depicted in Fig. 3c1 and with a constant strain rate of 10^{-14} s^{-1} . "B&K80": Brace and Kohlstedt (1980), "Kal95": Kohlstedt et al. (1995), "R&M87": Ranalli and Murphy (1987), "K83": Kirby (1983), "Mal98": Mackwell et al. (1998), "W&C90": Wilks and Carter (1990), "G&T95": Gleason and Tullis (1995), "R95": Ranalli (1995), "Kal86": Karato et al. (1986). (b) Down-dip extent of the interplate channel, z_{coup} , as a function of ΔE_a , for A_t set to 100 Myr and 60 Myr. The ΔE_a range for which partial kinematic coupling is modelled is depicted in grey, and the range over which z_{coup} and $z_{\text{partial-coup}}$ vary through time by dashed and solid lines, respectively. In hatched areas, z_{coup} is mesh-dependent and located at 250 km depth, yielding an excessively cold fore-arc structure.

Figure 6: Influence of the crustal strength reduction, ΔE_a , on subduction mode. One black isotherm every 200°C. The crust content at Eulerian nodes is $>49.0\%$ within the white outline. The numerical mesh is colored as a function of crust content, by focusing on very high contents ($>99\%$). Panel (a): Subduction mode with $A_t = 20$ Myr depending on ΔE_a if the oceanic crust is very buoyant ($\Delta \rho_c = -380 \text{ kg.m}^{-3}$, a1: simulation S2-8a, and a2: simulation S2-15a). Panel (b): Subduction of an old lithosphere for high ΔE_a . Stream function (green outlines) is depicted every $0.2 \text{ cm}^2/\text{yr}$ counted from $-1.7 \text{ cm}^2/\text{yr}$. Left (b1): $\Delta E_a = 145 \text{ kJ/mol}$ (simulation S1-14, Table 3). In the inset, the velocity color scale is saturated to 0.65 cm/yr . Right (b2): $\Delta E_a = 155 \text{ kJ/mol}$ (simulation S1-15). (b3): same simulation but including water transfer computation. The pink line is the 980°C isotherm, i.e., the lowest solidus temperature for a water-saturated peridotite. "Eclg": slab dehydration triggered by the gabbroic crust eclogitization. "Chl out": water expelling by chlorite destabilisation within the initially serpentised mantle layer. "Stable phase A": water retention in the subducting mantle as a result of phase A formation when serpentine breaks down. In the bottom inset, the logarithmic water content color scale highlights very low water amounts in the mantle wedge.

Figure 7: Subduction modes for a 20 Myr old incoming lithosphere as a function of the subducting crust buoyancy, $\Delta\rho_c$, and crust ductile weakening, ΔE_a . Experiments are depicted by small solid black circles (Table 3). In insets (1) to (7) illustrating the different subduction regimes, isotherms are displayed every 200°C, and the subducting crust by the grey-dashed area. (1) simulation S2-12a in Table 3; (2) simulation S2-15c; (3) simulation S2-22c; (4) simulation S2-10h; (5) simulation S2-20j; (6) simulation S2-2c; (7) simulation S2-6g.

Figure 8: Evolution of the velocity field and thermal structure for a buoyant and weak subducting crust, when crust plumes form (panel a, simulation S2-13g, Table 3), or not (panel b, simulation S2-17b). One isotherm (white outline) every 200°C . Within the thick black outline, crust content is higher than 49%. The color scale refers to velocity magnitude, arrows depict velocity vectors.

Figure 9: Maps of z_{coup} simulated at different time steps for $A_t = 20$ Myr as a function of crustal weakness, ΔE_a , and buoyancy, $\Delta \rho_c$, with respect to mantle. Depth outlines are plotted every 10 km. (a) z_{coup} modelled after 10 Myr of convergence. Experiments are depicted by grey-filled black circles. Passive water transfers and wet mantle melting estimate are computed in experiments depicted by purple symbols (see Fig. 10). Crosses: wet mantle welting extinction; solid triangles: stable hydrous melting; empty triangles: intermittent wet mantle melting. (b) Shallowest z_{coup} depth obtained between 10 and 13.5 Myr. Experiments showing partial kinematic coupling are depicted by solid black circles. See Fig. 7 caption for the meaning of green, red, blue, purple and black-dashed thick lines.

Figure 10: Prediction of water amounts (labelled $[\text{OH}^-]$) available in the mantle wedge to sustain mantle melting, for a 20 Myr old downgoing plate. Purple symbols in headings refer to Fig. 9a. (a) Water absorbed in the mantle wedge at temperature exceeding the minimum wet solidus temperature (980°C , pink isotherm) in amounts always $> 0.5\text{wt}\%$ sufficient to allow continuous melting. (b) Water absorbed at $T > 980^\circ\text{C}$ in amounts strongly varying through time, temporarily very low and yielding discontinuous magmatism in the mantle wedge. (c) Progressive decrease through time of the water amount absorbed at $T > 980^\circ\text{C}$, predicting mantle melting to vanish off.

Strength contrast modelling the subduction channel

Figure 11: Schematic diagram illustrating the main mechanism controlling kinematic decoupling along the plate boundary, as a function of slab weight and the ductile strength of the subduction channel.

Figure 12: Subduction mode as a function of the incoming plate age and the contrast in ductile strength between crust and mantle, for a buoyant subducting crust ($\Delta\rho_c = -380$ kJ/mol) and a 100 km thick overriding plate. Numerical experiments are depicted by solid black circles (Table 3). The boundaries between the different subduction regimes are arbitrarily extended by dashed lines (unexplored domains). The red coloured shading stands for the fore-arc heating by corner flow decreasing as ΔE_a is increased.

Figure 13: (a-c) Subduction thermal structures at steady-state as a function of ΔE_a . One white isotherm every 200°C . The black outline delimits the subducting crust. Thermal states modelled with $\Delta E_a = 80$ kJ/mol and $\Delta E_a = 105$ kJ/mol are displayed in Fig. 3a and 4c. Right panels display viscosity profiles along the dotted line depicted in left panels. (d) $P-T$ conditions modelled along the subducting crust surface (solid lines) compared to slab surface temperature estimated by geothermometry (Plank et al., 2009; Cooper et al., 2012; Watt et al., 2013) and sub-arc slab surface depth from Syracuse and Abers (2006) (blue solid squares, blue-dotted box). Lines with stars depict the hottest geotherm modelled in the mantle wedge, compared with conditions of last equilibration with the mantle of near-primary arc magmas (Tatsumi, 1981; Tatsumi et al., 1983; Bartels et al., 1991; Draper and Johnston, 1992; Falloon and Danyushevsky, 2000; Elkins Tanton et al., 2001; Grove et al., 2003; Hesse and Grove, 2003; Watt et al., 2013, red solid squares, red-dotted box).

Table 1: Characteristics of numerical grids.

Grid	LR	HRlim	HR	MHR
Resolution $x \times z$	332×90	482×129	482×151	407×119
Δx_{min} (km)	2.846	1.267	1.267	1.72
imposed between (km)	1012 < x < 1397	1004 < x < 1382	1004 < x < 1382	996 < x < 1389
Δx_{max} (km)	9.51	9.51	10.4	10.6
Δz_{min} (km)	2.268	1.266	1.266	1.656
imposed for (km)	$z > -116.1$	$z > -102.0$	$z > -136$	$z > -142^a$
Δz_{max} (km)	10.2	10.3	10.3	10.1
Oceanic crust thickness (km)	6.94	6.94	6.94	8.33
Total tracer number	1.2544×10^6	5.76×10^6	5.76×10^6	3.24×10^6

^a for $A_t = 60$ Myr and $A_t = 100$ Myr, Δz_{min} is imposed for depth shallower than 250 km when $\Delta E_a > 120$ kJ/mol.

Table 2: Parameter names and values.

Parameter name	Symbol	Value
Box height	H_0	555 km
Bottom temperature	T_b	1888 K
Surface temperature	T_s	273 K
Mantle density	ρ_m	3300 kg.m ⁻³
Density offset between crust and mantle	$\Delta\rho_c$	$\Delta\rho_c = \rho_c - \rho_m$
Crust density	ρ_c	varying
Mantle radiogenic heat production	A	9.20×10^{-8} W.m ⁻³
Adiabatic gradient	$(\frac{\partial T}{\partial z})_{adiab}$	0.45 K.km ⁻¹
Thermal diffusivity	κ	0.8×10^{-6} m ² .s ⁻¹
Thermal expansion coefficient	α	3.5×10^{-5} K ⁻¹
Heat capacity	C_p	0.971×10^3 J.(K.kg) ⁻¹
Activation energy for the mantle	E_a^m	465 kJ/mol
Activation energy for the crust	E_a^c	varying
Activation energy offset crust/mantle	ΔE_a	$\Delta E_a = E_a^m - E_a^c$
Activation volume	V_a	1.7×10^{-5} m ³ /mol
Pre-exponential factor in non-Newtonian rheology	A_0	339428.7 MPa ⁻³ .s ⁻¹
Dissipation number	$Di = \frac{\alpha g H_0}{C_p}$	0.196
Gravity acceleration	g	9.81 m.s ⁻²
Weak layer thickness	h_c	6.9 km or 8.3 km
Convergence rate	v_c	6.5 cm/yr
Subducting plate age at trench	A_t	varying
Cohesive strength at $z = 0$	τ_0	1 MPa
Stress exponent in the viscous rheology	n	3
Stress exponent in the brittle rheology	n_p	30
Yield stress increase with depth (mantle)	γ_m	1.6
Yield stress increase with depth (crust)	γ_c	0.065

Table 3: Simulation list. (*continued on next pages*)

Simulation	A_t Myr	ΔE_a kJ/mol	E_a^c kJ/mol	$\Delta \rho_c$ kg/m ³	ρ_c kg/m ³	Numerical grid	z_{dec} ^a km
S1-2	100	20	445	-380	2920	MHR	59.8
S1-4	100	40	425	-380	2920	MHR	63
S1-6	100	65	400	-380	2920	MHR	74
S1-8	100	80	385	-380	2920	MHR	81.7
S1-8b ^e	100	80	385	-380	2920	MHR	48
S1-10a	100	105	360	-380	2920	LR	80 km ^b
S1-10b	100	105	360	-380	2920	HRlim	102 ^b
S1-10c	100	105	360	-380	2920	HR	114
S1-10d	100	105	360	-380	2920	MHR	109
S1-12	100	120	345	-380	2920	MHR	135.3
S1-13	100	135	385	-380	2920	MHR	166
S1-14 ^c	100	145	155	-380	2920	MHR	162-200 ^b
S1-15 ^c	100	155	310	-380	2920	MHR	250 ^b
S6-2	60	20	445	-380	2920	MHR	- ^d
S6-4	60	40	425	-380	2920	MHR	65.8
S6-10	60	105	360	-380	2920	MHR	118.5
S6-13 ^c	60	135	385	-380	2920	MHR	124-219 ^b
S6-15	60	155	310	-380	2920	MHR	250 ^b
S2-2a	20	20	445	-95	3205	MHR	- ^d
S2-2b	20	20	445	0	3300	MHR	- ^d
S2-2c	20	20	445	120	3420	MHR	- ^d
S2-2d	20	20	445	180	3480	MHR	- ^d

Simulation	A_t	ΔE_a	E_a^c	$\Delta \rho_c$	ρ_c	Num. grid	z_{dec}^a
S2-2e	20	20	445	250	3550	MHR	_ ^d
S2-4a	20	40	425	-140	3160	MHR	_ ^d
S2-4b	20	40	425	-95	3205	MHR	_ ^d
S2-4c	20	40	425	-50	3250	MHR	_ ^d
S2-4d	20	40	425	0	3300	MHR	_ ^d
S2-4e	20	40	425	60	3360	MHR	_ ^d
S2-4f	20	40	425	120	3420	MHR	_ ^d
S2-4g	20	40	425	180	3480	MHR	58.1
S2-4h	20	40	425	220	3520	MHR	55
S2-4i	20	40	425	250	3550	MHR	50
S2-6a	20	60	405	-240	3060	MHR	_ ^d
S2-6b	20	60	405	-140	3160	MHR	_ ^d
S2-6c	20	60	405	-95	3205	MHR	_ ^d
S2-6d	20	60	405	-50	3250	MHR	_ ^d
S2-6e	20	60	405	0	3300	MHR	84
S2-6f	20	60	405	120	3420	MHR	72.9
S2-6g	20	60	405	220	3520	MHR	53.6
S2-6h	20	60	405	250	3550	MHR	50.5
S2-7	20	70	395	0	3300	MHR	83.7
S2-8a	20	80	385	-380	2920	MHR	_ ^d
S2-8b	20	80	385	-350	2950	MHR	_ ^d
S2-8c	20	80	385	-320	2980	MHR	_ ^d
S2-8d	20	80	385	-300	3000	MHR	_ ^d
S2-8e	20	80	385	-270	3030	MHR	_ ^d
S2-8f	20	80	385	-240	3060	MHR	_ ^d

Simulation	A_t	ΔE_a	E_a^c	$\Delta \rho_c$	ρ_c	Num. grid	z_{dec}^a
S2-8g	20	80	385	-195	3105	MHR	$_{-d}$
S2-8h	20	80	385	-140	3160	MHR	$_{-d}$
S2-8i	20	80	385	-120	3180	MHR	83.4
S2-8j	20	80	385	-95	3205	MHR	89.7
S2-8k	20	80	385	-60	3240	MHR	89.1
S2-8l	20	80	385	0	3300	MHR	87.6
S2-8m	20	80	385	60	3360	MHR	85.6
S2-8n	20	80	385	120	3420	MHR	83.4
S2-8o	20	80	385	195	3495	MHR	81.6
S2-8p	20	80	385	250	3550	MHR	75.7
S2-9a	20	90	375	0	3300	MHR	92.9
S2-10a	20	105	360	-380	2920	MHR	$_{-d}$
S2-10b	20	105	360	-350	2950	MHR	$_{-d}$
S2-10c	20	105	360	-320	2980	MHR	$_{-d}$
S2-10d	20	105	360	-300	3000	MHR	$_{-d}$
S2-10e	20	105	360	-270	3030	MHR	$_{-d}$
S2-10f	20	105	360	-240	3060	MHR	111.3
S2-10g	20	105	360	-190	3110	MHR	101.6
S2-10h	20	105	360	-140	3160	MHR	103.6
S2-10i	20	105	360	0	3300	MHR	101
S2-10j	20	105	360	120	3420	MHR	93.8
S2-10k	20	105	360	250	3550	MHR	88.2
S2-12a	20	120	345	-380	2920	MHR	$_{-d}$
S2-12b	20	120	345	-350	2950	MHR	$_{-d}$
S2-12c	20	120	345	-325	2975	MHR	119

Simulation	A_t	ΔE_a	E_a^c	$\Delta \rho_c$	ρ_c	Num. grid	z_{dec}^a
S2-12d ^c	20	120	345	-60	3240	MHR	113.3
S2-12e	20	120	345	60	3360	MHR	106.4
S2-12f	20	120	345	195	3495	MHR	98.8
S2-13a	20	135	330	-380	2920	MHR	- ^d
S2-13b	20	135	330	-350	2950	MHR	- ^d
S2-13c	20	135	330	-325	2975	MHR	145
S2-13d ^c	20	135	330	-300	3000	MHR	192
S2-13e	20	135	330	-270	3030	MHR	209
S2-13f ^c	20	135	330	-240	3060	MHR	183
S2-13g ^c	20	135	330	-190	3110	MHR	194
S2-13h	20	135	330	-140	3160	MHR	128.7
S2-15a ^c	20	155	310	-380	2920	MHR	133
S2-15b	20	155	310	-350	2950	MHR	142
S2-15c	20	155	310	-300	3000	MHR	189
S2-15d	20	155	310	-240	3060	MHR	207
S2-15e	20	155	310	-190	3110	MHR	200
S2-15f	20	155	310	0	3300	MHR	126.2
S2-15g	20	155	310	120	3420	MHR	116
S2-15h	20	155	310	250	3550	MHR	111
S2-16a	20	160	305	-60	3240	MHR	143.2
S2-16b	20	160	305	60	3360	MHR	122.7
S2-17a ^c	20	175	290	-380	2920	MHR	201.4
S2-17b ^c	20	175	290	-300	3000	MHR	137.2
S2-17c ^c	20	175	290	-240	3060	MHR	250 ^b
S2-17d	20	175	290	-195	3105	MHR	207

Simulation	A_t	ΔE_a	E_a^c	$\Delta \rho_c$	ρ_c	Num. grid	z_{dec}^a
S2-17e	20	175	290	-140	3160	MHR	175
S2-17f	20	175	290	120	3420	MHR	121.9
S2-17g	20	175	290	250	3550	MHR	117.6
S2-20a	20	200	265	-380	2920	MHR	117.2
S2-20b	20	200	265	-300	3000	MHR	115.1
S2-20c	20	200	265	-240	3060	MHR	189
S2-20d ^c	20	200	265	-195	3105	MHR	174
S2-20e	20	200	265	-140	3160	MHR	190
S2-20f ^c	20	200	265	-95	3205	MHR	167
S2-20g	20	200	265	-60	3240	MHR	174.1
S2-20h	20	200	265	0	3300	MHR	150.9
S2-20i	20	200	265	60	3360	MHR	132.2
S2-20j	20	200	265	195	3495	MHR	135.7
S2-22a	20	220	245	-380	2920	MHR	109
S2-22b	20	220	245	-300	3000	MHR	99
S2-22c	20	220	245	-190	3110	MHR	170.9
S2-22d	20	220	245	-95	3205	MHR	156
S2-22e	20	220	245	0	3000	MHR	170.2
S2-22f	20	220	245	120	3420	MHR	153.4
S2-22g	20	220	245	250	3550	MHR	146

^aCoupling depth at steady-state ($t > 8$ Myr) if $A_t > 20$ Myr; z_{coup} at 10 Myr of convergence if $A_t = 20$ Myr.

^bCoupling depth dependent on numerical mesh.

^cRun includes the computation of passive water transfers.

^dFlat and aborted subduction.

^eRun includes water transfer computation and a mantle strength reduction depending on water content.