

Rapid directivity detection by azimuthal amplitude spectra inversion

Simone Cesca, Sebastian Heimann, Torsten Dahm

▶ To cite this version:

Simone Cesca, Sebastian Heimann, Torsten Dahm. Rapid directivity detection by azimuthal amplitude spectra inversion. Journal of Seismology, 2010, 15 (1), pp.147-164. 10.1007/s10950-010-9217-4. hal- 00642936

HAL Id: hal-00642936

https://hal.science/hal-00642936

Submitted on 20 Nov 2011

HAL is a multi-disciplinary open access archive for the deposit and dissemination of scientific research documents, whether they are published or not. The documents may come from teaching and research institutions in France or abroad, or from public or private research centers. L'archive ouverte pluridisciplinaire **HAL**, est destinée au dépôt et à la diffusion de documents scientifiques de niveau recherche, publiés ou non, émanant des établissements d'enseignement et de recherche français ou étrangers, des laboratoires publics ou privés.

ORIGINAL ARTICLE

Rapid directivity detection by azimuthal amplitude spectra inversion

Simone Cesca · Sebastian Heimann · Torsten Dahm

Received: 24 February 2010 / Accepted: 4 November 2010 © Springer Science+Business Media B.V. 2010

1 Abstract An early detection of the presence of 2 rupture directivity plays a major role in the correct 3 estimation of ground motions and risks associated 4 to the earthquake occurrence. We present here 5 a simple method for a fast detection of rupture 6 directivity, which may be additionally used to 7 discriminate fault and auxiliary planes and have 8 first estimations of important kinematic source parameters, such as rupture length and rupture 10 time. Our method is based on the inversion of 11 amplitude spectra from P-wave seismograms to 12 derive the apparent duration at each station and 13 on the successive modelling of its azimuthal be-14 haviour. Synthetic waveforms are built assuming a 15 spatial point source approximation, and the finite 16 apparent duration of the spatial point source is 17 interpreted in terms of rupture directivity. Since 18 synthetic seismograms for a point source are calculated very quickly, the presence of directivity 20 may be detected within few seconds, once a focal 21 mechanism has been derived. The method is here 22 first tested using synthetic datasets, both for lin-23 ear and planar sources, and then successfully ap-24 plied to recent Mw 6.2-6.8 shallow earthquakes in

automated application and may be used to im- 26 prove kinematic waveform modelling approaches. 27

Keywords Directivity · Earthquake source · 29 Kinematic model · Amplitude spectra 30

1 Introduction 31

Tectonically driven shallow earthquake sources 32 are generally explained by means of shear cracks 33 occurring along a limited, almost planar region, 34 we refer as the focal region. A point source repre- 35 sentation is a common first approximation, which 36 is valid when treating far-field low frequency seis- 37 mic waveform, using wavelengths larger than the 38 rupture size. Higher frequencies seismograms and 39 spectra contain information which can be related 40 to the finiteness of the rupture process and thus 41 can be used to determine parameters describing 42 the finite source. Size and shape of the rupture 43 area, rupture velocity and preferential rupture 44 directions, an effect known as rupture directivity, 45 are some of the parameters which can be retrieved 46 by the analysis of high-frequency waveforms. In 47 particular, we are interested here in discussing the 48 problem of early detection of rupture directiv- 49 ity, distinguishing between a prominent or partial 50 unilateral rupture (a case which will be further 51 referred as asymmetric bilateral rupture), and a 52

S. Cesca (⊠) · S. Heimann · T. Dahm Institut für Geophysik, Universität Hamburg, Bundesstrasse 55, 20146 Hamburg, Germany e-mail: simone.cesca@zmaw.de

25 Peloponnese, Greece. The method is suitable for



53 bilateral one, with rupture nucleating at the centre of the rupture area and propagating toward its edges. The azimuthal dependency of amplitudes and durations of different seismic phases is a first indicator of directivity effects and is consequence of the characteristics of the finite rupture process along the fault plane, specifically the main di-59 rection and speed of the rupture front propaga-60 tion. Directivity has been often observed and has been modelled for several earthquakes in the past, with several studies treating specific earthquakes or limited datasets (e. g., McGuire et al. 2002; Warren and Shearer 2006; Caldeira et al. 2009). A quick detection of directivity effects is important towards a correct estimation of ground motions, stress field perturbations and tsunamogenic risks and consequently to mitigate earthquake effects. These considerations provide important reasons to further investigate and develop specific tools for stable, rapid and automated directivity detection, which can be used within early warning systems. 74

Several methods have been applied in the past 76 to detect and classify earthquake source directivity. A common approach is the identification of predominant unilateral ruptures from the time duration and spectral analysis of body wave pulses (e.g. Boore and Joyner 1978; Beck et al. 1995; Warren and Shearer 2006; Caldeira et al. 2009). 82 Pulse lengths at different stations are interpreted in terms of the apparent duration of the source time function (STF), and their variation in dependence on azimuth and incidence angle is interpreted to detect directivity: similarly to a Doppler effect in classical physics, shorter STFs would indicate a rupture propagating towards the considered station, while longer pulses indicate a rupture propagation in the opposite direction. Directivity effects may also be revealed based on the analysis of surface waves at different azimuths (Ben-Menahem 1961; Pro et al. 2007). Whereas time domain methods remain more common, a significant contribution within this type of inversion methods was provided by the spectral approach discussed in Warren and Shearer (2006). This method is based on the spectral estimation of the pulse broadening and accounts for the az-100 imuthal and incidence angle dependencies; it is 101 well suited for the analysis of intermediate and deep focus earthquakes and was successfully ap- 102 plied to several events. The main limits of this 103 class of methods are related to the fact that wave 104 propagation and the superposition of different 105 seismic phases are not accounted, since wave 106 propagation effects between source and receiver 107 (Green's functions) are limited to the estimation 108 of the incidence angle of given seismic phases. 109 Another possible limitation is the requirement of 110 several stations with good azimuthal coverage in 111 order to ensure reliable results. A second range 112 of applications, which on the contrary accounts 113 precisely for the effects of the earth's model on 114 the observed waveforms, is based on empirical 115 Green's functions technique (Hartzell 1978; Li 116 and Toksöz 1993; Velasco et al. 1994; Cassidy 117 1995; Müller 1985; Velasco et al. 2004; Vallée 118 2007). In this case, an aftershock with common 119 hypocenter and focal mechanism of the studied 120 event can be used to remove path effects, and iso- 121 late finite source apparent durations at different 122 stations. Evidently, the application of these tech- 123 niques is strongly limited by the availability of 124 a proper aftershock. Brüstle and Müller (1987), 125 and Imanishi and Takeo (2002) have investigated 126 the adoption of master-event techniques to detect 127 directivity: the identification of stopping phases 128 (Madariaga 1977, 1983; Bernard and Madariaga 129 1984; Spudich and Frazer 1984) at different sta- 130 tions was used there to determine the main di- 131 rection of rupture propagation, besides other 132 source properties. Stopping phases identification 133 (Imanishi and Takeo 1998, 2002) typically re- 134 quires a careful waveform analysis, which may be 135 hardly implemented within automated routines. 136 A third group of techniques are based on com- 137 plete kinematic waveform inversion, with the aim 138 of retrieving a most detailed image of the finite 139 rupture process, not limited to the identification 140 of directivity. The range of methods and appli- 141 cations is very wide, including higher order mo- 142 ment tensor analysis (Dahm and Krüger 1999; 143 McGuire et al. 2001, 2002), detailed slip map 144 approaches (e.g. Olson and Apsel 1982; Hartzell 145 and Helmberger 1982; Hartzell and Heaton 1983; 146 Beroza and Spudich 1988), and inversion meth- 147 ods adopting constrained and simplified kinematic 148 models (Dreger and Kaverina 2000; Vallée and 149 Bouchon 2004; Gallovic et al. 2009; Cesca et al. 150

171

151 2010). All these methods have a significant poten-152 tial for a stable determination of directivity but 153 their adoption towards its very fast detection is 154 limited, often requiring time consuming computa-155 tion of synthetic seismograms for several extended 156 source models. Methods developed by Dreger and Kaverina (2000) and following Cesca et al. (2010) 158 have shown a good performance and have been tested for near real-time applications, but they are still based on extended source representations and 161 thus require heavier computations with respect to 162 our method. Finally, recent results by Zahradnik et al. (2008) showed the possibility of discriminating the true fault plane on the base of spatial offsets between epicentre and centroid locations. 166 However, the method has been currently applied only to a limited number of earthquakes, with 168 variable results, and the determination of directiv-169 ity may be beyond its possibilities, for example for 170 symmetric bilateral ruptures.

We present here a simple alternative method to 172 quickly detect directivity for shallow earthquakes and discuss it with the aid of a set of applications, including both synthetic datasets and observations from recent earthquakes in Greece. Our 176 method is based on a point source representation, which drastically reduces computational require-178 ments and makes it feasible for early detection. Directivity is detected on the basis of a frequency domain inversion of the apparent duration at each station and the further interpretation of its azimuthal variation. Main strength points of the proposed method include the adoption of a common dataset and modelling tools for focal mechanism and directivity determination, the inclusion of Green's functions accounting for wave propagation through the chosen earth models without needing specific aftershocks, and the simplicity and quickness of the inversion process.

190 2 Directivity and amplitude spectra inversion

191 We make here use of the recently developed 192 Kiwi tools (Heimann 2010; Cesca et al. 2010; 193 http://kinherd.org), which provide a flexible in-194 strument to generate synthetic seismograms for 195 point and extended sources and to invert different 196 earthquake source parameters, allowing the selection of different waveform tapers, frequency 197 filters, inversion domains and misfit functions. 198 Cesca et al. (2010) showed successful applications 199 to shallow earthquake at regional distances, and 200 was able to derive both point source (best double 201 couple, DC, model, scalar moment and centroid 202 depth) and extended source (fault plane discrim- 203 ination, rupture size, rupture time, rupture nucle- 204 ation) parameters.

The first inversion step follows the approach 206 described in Cesca et al. (2010), to obtain the focal 207 mechanism, scalar moment and centroid depth: 208

- 1. Focal mechanism. We invert amplitude spec- 209 tra of full waveforms, according to Cesca et al. 210 (2010), to derive a point source focal mech- 211 anism (DC, depth and scalar moment); the 212 source epicentral location is assumed to be 213 originally known.
- Polarities. The focal mechanism presents a 215 polarity ambiguity, which can be solved by 216 comparing observed displacements and syn- 217 thetic seismograms for the two possible polar- 218 ity configurations; however, the detection of 219 the true polarity is here not strictly required, 220 as the whole inversion process is carried out 221 in the frequency domain, and only amplitude 222 spectra are involved in the fitting procedure.

The source representation through the Kiwi tools 224 allows the adoption of different rise times. For a 225 spatially extended source model, where the rup- 226 ture region is discretised into a number of spatial 227 point sources, the rise time represent the time 228 during which each point source radiates seismic 229 energy. The duration of the whole rupture process 230 is related to rise and rupture times. If we adopt 231 a point source representation, the rise time will 232 represent the duration of the source time function. 233 This parameter was used in Cesca et al. (2010) 234 to have a first, rough, estimate of rupture times 235 and to choose a proper rise time during kinematic 236 source modelling. We proceed here differently: in- 237 stead of determining true duration of the rupture 238 process, we investigate apparent durations as seen 239 by individual stations.

In detail, during the second inversion step, we 241 proceed as follows (Fig. 1 illustrates an example 242 of the main steps, relative to selected seismic 243



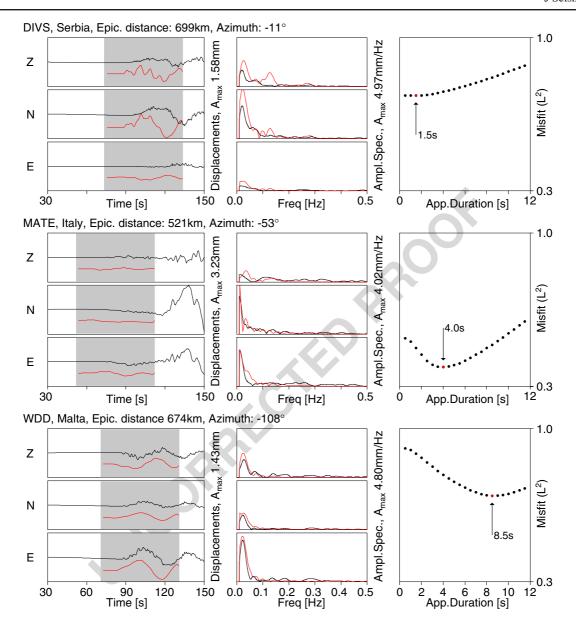


Fig. 1 Example of the procedure followed to derive the apparent source duration at different stations. Selected waveforms, spectra and amplitude spectra inversion results refer to an application to the Andravida earthquake, Greece, which is further extensively discussed in this study. Left: filtered displacements (black lines) and synthetic seismograms (red lines) for the chosen point source model are

tapered to select P waves time windows (grey intervals). Centre: amplitude spectra comparison (red lines correspond to the best fitting synthetic spectra, after comparing several source durations). Right: comparison of amplitude spectra misfit values for different source durations (best solutions for each station are identified by red circles)

244 waveforms from the Andravida earthquake, 245 which is later discussed in the text):

Waveform selection. We use all available spatial components, preferably using North, East

and vertical orientations, rather than rotated 248 traces, in order to have P wave energy on all 249 traces (which is theoretically null on transversal components); the presence of more traces 251 for each station has a smoothing effect; after 252



283

284

285

286

287

288

289

290

291

292

293

294

295

296

297

- 253 testing with different datasets, we found that more components provide more stability. We 254 perform a deconvolution of the instrumental 255 response from the data, and conversion to 256 displacements. 257
- Tapering. We limit the inversion process to P-258 2. wave time windows, which are automatically 259 selected on the base of the source-receiver 260 geometry and theoretical arrival time for the 261 earth model used during the inversion (an ar-262 rival time database is calculated in advance, to 263 reduce computational effort at the time of the 264 inversion); for the case studies here described 265 we use 60 s length time windows, starting 15 s 266 before theoretical first P arrival, and apply 267 a bandpass filter in the range 0.01-0.5 Hz 268 (these parameters may be modified depending 269 on the earthquake size, the source depth, the 270 average duration, and the range of epicen-271 tral distances where waveforms are inverted). 272 Tapers should be chosen in order to resolve 273 directivity effects. A minimum length should 274 account at least for two times the average rup-275 ture duration and for different periods at the 276 frequency range used for the inversion. For 277 stations located at small epicentral distances, 278 with minor delay between S and P phases, 279 tapers may be modified to avoid S waves. 280
- Scalar moment inversion. Since the estima-281 3. tion of the scalar moment may slightly vary depending on the inversion approach (e.g., full waveform or body waves, time domain or amplitude spectra inversion, etc.), we mention here the possibility to perform a specific inversion using an approach consistent with the following directivity inversion. Traces from all seismic stations would be used to invert the scalar moment (e.g. by amplitude spectra inversion, using a Levenberg-Marquardt approach and an L² norm misfit function). In the following applications this step is not performed, as we count with stable estimations of the scalar moments, provided by the fit of low frequency amplitude spectra from the whole waveforms.
- Apparent duration inversion. For each of the 298 4. stations, we perform an amplitude spectra in-299 version to derive the apparent source duration 300 at that station; the frequency domain inver-301

sion approach is less sensitive to unmodelled 302 structural heterogeneities; we perform here 303 a grid search for possible durations (for the 304 following case studies, tested durations varies 305 up to 30 s, with an increment of 0.5 s); in 306 general we observe smooth single-minimum 307 curves of misfit versus apparent durations, 308 and tests with different inversion approaches 309 (e.g. gradient methods) have shown very con- 310 sistent results with respect to the grid walk 311 procedure.

312

336

The apparent source time function durations can 313 be then quickly interpreted in term of simplified 314 laws for finite rupture models. With the aid of 315 synthetic tests and application to selected earth- 316 quake datasets, we will show that, often, it is not 317 necessary to have a complex rupture model to 318 fit the azimuthal distribution of apparent dura- 319 tions. For simple extended source model, such as 320 a one-dimensional linear source or a Haskell bi- 321 dimensional rupture model (Haskell 1964), the 322 effects of directivity can be treated analytically. A 323 unilateral rupture along a horizontal linear source 324 will produce theoretical P-wave pulses of shorter 325 duration for stations located toward the rupture 326 propagation, and larger duration for stations in 327 the opposite direction. Bilateral ruptures result 328 in a minor azimuthal variation of the apparent 329 source time function. A range of asymmetrically 330 bilateral rupture models exists in between. Effects 331 of oblique and vertical rupture propagations may 332 also be modelled but are more difficult to reveal 333 (Beck et al. 1995) and have been more rarely 334 observed (e.g. Eshghi and Zare 2003; Nadim et al. 335 2004).

We originally focus on the two-dimensional 337 problem, with source and observer laving on the 338 same plane. Let us assume a horizontal linear 339 source model of length L, with the rupture starting 340 at one edge (A) and propagating unilaterally till 341 the other edge (B). The rupture time t_R is the time 342 required for the rupture front to propagate along 343 the entire rupture length, from A to B, at a rupture 344 velocity $v_{\rm R}$, which is assumed to be constant. The 345 rise time t_r , defined as the duration of seismic 346 source emission from a point along the source, is 347 here assumed to be constant, according to healing 348 front theory (Nielsen and Madariaga 2003) and 349



350 will be further considered negligible with respect 351 to the rupture time. Finally, v_P is the average P 352 wave velocity at the focal region. Typically, rup-353 ture propagates with a velocity slightly below the shear wave velocity at the focal region, which also shows a common scale with compressional wave velocity in seismogenic regions. Then, according to Ben-Menahem and Singh (1981), and including 358 the rise time, for a receiver located at azimuth φ (defined with respect to the direction of rupture propagation) the apparent source duration $\Delta t(\varphi)$ 361 will be given by:

$$\Delta t(\varphi) = t_{\rm r} + \frac{L}{v_{\rm R}} - \frac{L}{v_{\rm P}} \cos(\varphi). \tag{1a}$$

362 In view of a more general formulation, also ac-363 counting for asymmetric and pure bilateral rup-364 ture, the rupture length L is divided into two 365 segments L_1 and L_2 , with the following expression 366 for the apparent source duration $\Delta t(\varphi)$:

$$\Delta t(\varphi) = \operatorname{Max} \left[t_{r} + L_{1} / v_{R} - \left(L_{1} / v_{P} \right) \cos(\varphi) , \right.$$
$$\left. t_{r} + L_{2} / v_{R} + \left(L_{2} / v_{P} \right) \cos(\varphi) \right] (1b)$$

367 We can then introduce the following non-368 dimensional variables: $\tau(\varphi)$ is the ratio between 369 the apparent source duration $\Delta t(\varphi)$ and the rup-370 ture time t_R , $t_{r/R}$ is the ratio between rise and rupture time, $v_{R/P}$ is the ratio between rupture velocity and P wave velocity at the source; L₁ and 373 L_2 ($L_1 \ge L_2$) are expressed as $(1-\chi)L$ and χL , respectively, x being the ratio between the shortest segment and the entire rupture length (χ may 376 range from 0, for a pure unilateral rupture, to 0.5, for a pure bilateral one). The azimuthal dependency of $\tau(\varphi)$, making use of the non-dimensional 379 notation is the following ($t_{r/R}$ can be in general 380 neglected):

$$\tau (\varphi) = \operatorname{Max} \left[t_{r/R} + 1 - v_{R/P} \cos(\varphi) , \right.$$

$$t_{r/R} + L_2 / L_1 + L_2 / L_1 \cos(\varphi) \right]. (2)$$

381 The radiation pattern for three significant cases (pure unilateral, pure bilateral and asymmetric 383 bilateral) is shown in Fig. 2, where we have chosen $v_{\rm R/P} = 0.5$ ($v_{\rm R/P}$ equal to 0.25 and 1.0 for the slow and fast cases respectively), χ is equal to 0, 1/3 and 1/2 for the three considered cases. We choose different rupture lengths L, in order to have a constant length of the largest rupture segment. Symmetries of apparent duration radiation pat- 389 terns can be observed, with a one lobe shape for 390 a pure unilateral rupture and a two lobe shape for 391 a pure bilateral one. The theoretical curve of the 392 apparent rupture duration (Fig. 2, right) for the 393 unilateral case range from $t_r + t_R - t_P$ (azimuth of 394 rupture direction) to $t_r + t_R + t_P$ (opposite direc- 395 tion); its average value is equal to $t_r + t_R$ (for a 396 unilateral rupture model t_R and t_P are the rupture 397 time and P wave travel time along the entire 398 rupture length). For the pure bilateral rupture, the 399 apparent duration varies between $t_r + t_R$ (perpen-400 dicular to rupture direction) to $t_r + t_R + t_P$ (paral-401 lel to rupture direction), with t_R and t_P referring 402 here to half of the rupture length. The larger vari- 403 ation of the apparent duration for the unilateral 404 case, with respect to the bilateral, explains the 405 major difficulties in observing directivity for the 406 second case. Less known is the behaviour of asym- 407 metric bilateral ruptures, although this model is 408 the most general. In this case the radiation pat- 409 tern (Fig. 2, bottom) present a deformed one- 410 lobe shape, with the minimum observed apparent 411 rupture duration at about 45° from the rupture di- 412 rection of the largest rupture length. The azimuth 413 α , where a cusp-like minimum in the apparent 414 duration may be observed, can be obtained by 415 equalizing the two right terms in Eq. 2:

$$\alpha = a \cos \left[\frac{(1 - 2\chi)}{v_{R/P}} \right]. \tag{3}$$

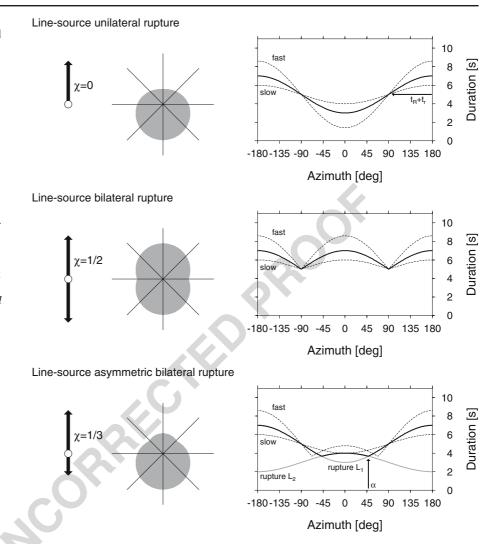
Since v_R pertains to $[0, v_P]$, it follows that for a 417 pure unilateral rupture ($\chi = 0$) we have a sin- 418 gle minimum in direction of rupture propagation, 419 while cusp-like minima are not observed. For a 420 pure bilateral rupture ($\chi = 0.5$), $\alpha = \pm \pi/2$ always. 421 For the intermediate case of asymmetric bilateral 422 ruptures (0 < χ < 0.5), two cusp-like minima are 423 observed if the following condition is met: 424

$$\chi > \frac{\left(1/v_{\text{R/P}}\right)}{2}.\tag{4}$$

This means that the observation of two minima in 425 the apparent duration curve indicate a dominant 426 bilateral rupture processes. 427

Figure 2 additionally shows how larger vari- 428 ations in the apparent duration estimations are 429 found, when increasing the rupture velocity (in 430 the figure, effects for an extreme case of $v_R = 431$

Fig. 2 Theoretical models for pure unilateral (top), pure bilateral (centre) and asymmetric bilateral (bottom) line sources. The ruptures start at a nucleation point (white circles, left plots) and propagate along segments L₁ and L₂ (black arrows), producing different apparent duration radiation patterns (grev regions, central plots) which have an azimuthal dependence. Right plots present the curves of the apparent duration versus azimuth: thick black lines represent an average behaviour $(v_R = 0.5v_P)$, while dashed lines represent fast $(v_R = v_P)$ and slow $(v_R = 0.25v_P)$ rupture cases. In the case of a partial unilateral rupture, the curve of apparent duration for the average case (thick black line) is given by the maximum of the curves associated to unilateral ruptures along the two segments (thick grey lines)



432 v_P are considered). On the other hand, a slower 433 rupture tends to behave similarly to a spatial point 434 source model (with a finite time duration); Fig. 2 435 shows the case of $v_R = 0.25v_P$, a proportion which 436 is not unrealistic and can be proper for shallow 437 earthquake in sediment layers (e.g. Selby et al. 438 2005; Dahm et al. 2007).

For a 3D earth model, directivity effects do not depend only on station azimuth but also on take-off angles. Equations 1a, 1b and 2 will then depend on the angle between the rupture direction and the ray direction at the source, instead than on station azimuth. These considerations opened space for specific detailed inversion approaches, such as the successful study by Warren and Shearer (2006), which accounted for take off angles and

required a dense station distribution both in terms 448 of azimuthal coverage and range of epicentral dis- 449 tances. We discuss here the problem for the case 450 of unilateral rupture. A first observation is that 451 even for the same source-receiver configuration, 452 different rays travelling along different paths, with 453 different take-off angles, will present different 454 directivity effects. For any considered bodywave 455 with take-off angle θ , we predict theoretically a 456 minor variability of the apparent duration with 457 respect to the one modelled for a planar case. For 458 example, the maximal apparent duration, at the 459 azimuth opposite to the direction of the rupture 460 propagation, will be equal to $t_r + t_R + t_P \cos(\theta)$, 461 which differs from Eq. 1a. Additionally, we have 462 the superposition of different P wave arrivals, with 463



518

544

464 different take-off angles. In any case, the interpre-465 tation of the maximal apparent duration following 466 Eq. 1a and neglecting take-off angles and rise time effects will lead to an overestimation of the terms $t_{\rm R}$ and $t_{\rm P}$ and consequently an overestimation of the rupture length. The value derived in such approximation can be safely considered as an upper bound to the real rupture length.

Our approach here is to model only the az-472 473 imuthal variation of directivity effects. The approximation is limited by different conditions. 475 First, we will focus on shallow earthquakes and 476 use only regional distances seismograms (epicentral distances in the range 200-1,000 km), basing our inversion on the fit of time windows centred at the first P wave arrivals. Additionally, we will consider only the case of horizontal to sub-horizontal ruptures (rupture propagating along direction dipping at most 20°). In these circumstances, we will show that the main rupture propagation can be detected even with the approximated approach 485 here described. The interpretation of additional rupture parameters is strongly limited by neglecting take-off angles and rise time effects, and the derived rupture length should be considered as an 488 upper bound to its real value. 489

On the base of the previous discussion, we can 490 491 now explain the last step of the inversion approach, where the distribution of apparent source 493 time durations is interpreted in terms of simplified 494 rupture models:

- Point source model. If we assume a spatial 495 1. point source model, with a given source time function of duration t_r , the curve of apparent duration will be a straight line, with no azimuthal dependence. This is the particular case, which can also be described by the previous equations, with L = 0, which leads to $t_{\rm R} = t_{\rm P} = 0$. The model has a unique unknown, t_r (by definition, rise time), which represents the apparent rupture time everywhere.
- 505 2. Pure/predominant unilateral rupture. In this case, the azimuthal distribution of apparent 506 durations is expected to follow a sinusoidal be-507 508 haviour. The fit is expected to be larger in case of a pure unilateral rupture, but the model can 509 still well reproduce data also for asymmetric 510 511 bilateral ruptures. In both cases, the minimum

of the curve will indicate the main rupture 512 direction (φ_0) . The model is represented by 513 the curve of the apparent duration $\Delta t(\varphi)$:

$$\Delta t(\varphi) = -A\cos(\varphi - \varphi_0) + B \tag{5}$$

with A and B positive, describing in first approx- 515 imation the travel time of P waves along the rup- 516 ture length (t_P) and the sum of rise and rupture 517 times $(t_r + t_R)$.

3. Pure bilateral rupture. In case of a pure bilat- 519 eral rupture, we will use the following curve, 520 instead: 521

$$\Delta t(\varphi) = A \left| \cos \left(\varphi - \varphi_0 \right) \right| + B. \tag{6}$$

4. Comparison of directivity models. Since 522 Eqs. 5 and 6 are dependent on three unknown 523 parameters, while the standard average rup- 524 ture time has only one (the average earth- 525 quake duration or duration of the common 526 source time function), they will always pro- 527 vide a better fit with respect to the common 528 rupture duration model. An F test can than 529 be used to evaluate the misfit improvement 530 versus the increase of degrees of freedom. F 531 values above 0.5 will be used here to prefer 532 unilateral or bilateral models with respect to 533 the point source solution. The modelling of 534 asymmetric bilateral rupture requires more 535 free parameters, being the superposition of 536 two functions as in Eq. 5. Chances for the im- 537 plementation of such a model are highly lim- 538 ited by several factors, including data quality, 539 focal mechanism, epicentral azimuthal cover- 540 age and local structural heterogeneities. We 541 suggest here the adoption of this model only 542 for specific cases rather than its implementa- 543 tion within the automated processing.

The major advantages of the presented approach 545 are that only point source synthetic seismograms 546 are used, so that forward modelling and inversion 547 are fast. A second advantage lies in the intrinsic 548 simplicity of this approach, which provides simple 549 plots of easy interpretation and suggest its imple- 550 mentation for automated routines. In the worse 551 case, if no clear pattern is detected, the retrieved 552 information can still be used to better focus a 553 more detailed full waveform kinematic inversion 554

496

497

498

499

500

501

502

503

555 (e.g., as in Cesca et al. 2010), by limiting the 556 range of rupture times and/or spatial extension 557 to be tested. A last important advantage resides in the coherency of the inversion approach: we use the same tools and the same data to derive 560 first the focal mechanism (or moment tensor) and to detect directivity effects, instead of relying on an externally calculated focal mechanism solution, which may be biased by a specific selection of stations, earth model, and processing routines. An 565 important limitation of the method applicability is represented by a specific range of earthquake magnitudes. For small earthquakes, with the rupture process occurring in less than 2 s, the variation of the apparent rupture duration could be only detected by fitting high frequency spectra (up to 0.5 Hz or above), which requires a detailed knowledge of the crustal structure, and possibly the adoption of 3D earth models. On the opposite side, large earthquakes may present more complex rupture processes, with different slip patches or asperities, breaking at different times. Large earthquakes also present significant discrepancies between hypocentral and centroid locations. Both effects are not considered in our model and can then lead to erroneous interpretations. As a rule of thumb, we believe that earthquakes with magnitudes in a range Mw 5.5-7.0 may be the best 583 suited for a successful application.

584 3 Synthetic tests for linear and planar sources

585 Before applying the method to real data, and in order to assess the method performance, we carry out a set of inversions using synthetic datasets. 588 We generate synthetic seismograms first for linear sources and consequently to planar ones. Considered source models present different significant focal mechanisms and directions of rup-592 ture propagation. Both synthetics for linear and planar sources are generated using the Kiwi tools 594 (http://kinherd.org; Heimann 2010), assuming the 595 PREM model (Dziewonski and Anderson 1981). The inversion is then carried out as described in the previous paragraph, assuming a spatial point source. Focal mechanisms, scalar moment and depth are retrieved at a first stage, by using the 600 method described in Cesca et al. (2010), which has been here specifically modified in order to include 601 the retrieval of the apparent rupture time at each 602 station. In the synthetic tests, a dense grid of 154 603 stations is considered, in order to plot apparent 604 duration contours. Stations location accomplish 605 to the chosen conditions in terms of epicentral 606 distances.

Three line source mechanisms (strike-slip (SS) 608 strike $\varphi = 30$, dip $\delta = 90$, rake $\lambda = 0$; normal fault 609 (NF) $\varphi = 30$, $\delta = 45$, $\lambda = 90$; thrust fault (TF) $\varphi = 610$ 30, $\delta = 20$, $\lambda = -90$) and three rupturing mod- 611 els (pure unilateral, asymmetric bilateral, pure 612 bilateral) are considered at first, thus providing 613 a set of nine source models. Ruptures propagate 614 horizontally for the strike-slip and normal fault, in 615 direction NNE-SSW, and toward ESE along the 616 low-angle dipping plane, for the thrust fault. The 617 source model centroid is always located at a depth 618 of 20 km, the source length is 30 km, rupture ve- 619 locity is 3.5 km/s. Pure unilateral ruptures start at 620 the southern (SS, NF) or western (TF) edge. The 621 same main rupture direction is used for asymmet- 622 ric bilateral ruptures, the two ruptured segments 623 having lengths $L_1 = 22.5$ and $L_2 = 7.5$ km. Rise 624 time is fixed to 2 s in all cases. The inversion 625 is carried out using 40 s time windows, starting 626 10 s before the theoretical arrival of P phases. A 627 frequency bandpass, between 0.01 and 0.5 Hz, is 628 used to filter Green's functions and data. Inver- 629 sion results are shown in Fig. 3. Coloured surface, 630 representing the inverted apparent source dura- 631 tion, highlight the radiation patterns of apparent 632 duration. The azimuthal distribution of apparent 633 duration clearly shows directivity effect and its 634 minor dependence on epicentral distance using 635 our approach and proof that a good quality fit 636 can be achieved using the simplified azimuthal de- 637 pendent curves. In particular, the fit of the cosine 638 curve described by Eq. 5 can be used to detect uni- 639 lateral or asymmetric bilateral ruptures, while the 640 curve from Eq. 6 to detect pure bilateral ruptures. 641 Rupture directivity is correctly detected in all 642 cases, with the exception of the bilateral rupture 643 along the low-dipping angle plane of the thrust 644 mechanism (Fig. 3, bottom right), where unilateral 645 rupture is incorrectly estimated. The reason for 646 this discrepancy can be described as follows. Since 647 the extended source is not horizontal, the seg- 648 ment toward ESE and WNW are located below 649



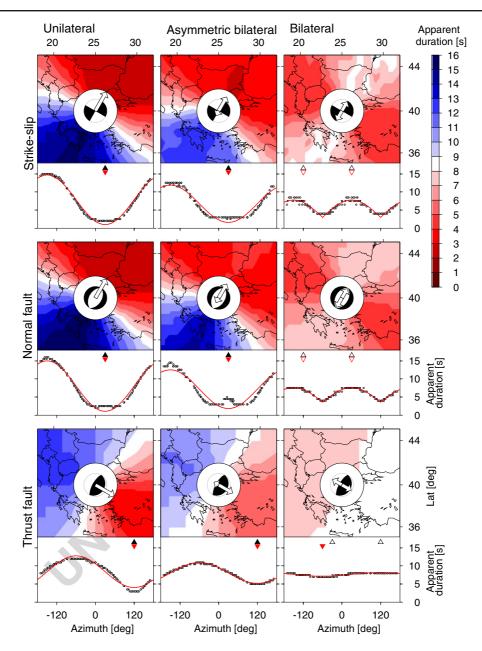


Fig. 3 Inversion results for linear sources. We consider three focal mechanisms (strike-slip, top; normal fault, centre; thrust fault, bottom) and three rupture processes (pure unilateral, left; asymmetric bilateral, centre; pure bilateral, right). For each case, we show colour plots representing the inverted apparent duration at a dense grid of station around the epicentre (red to blue scale represents increasingly longer apparent durations). The focal mechanisms are shown at the epicentral location, together with white arrows describing rupture directions (arrow

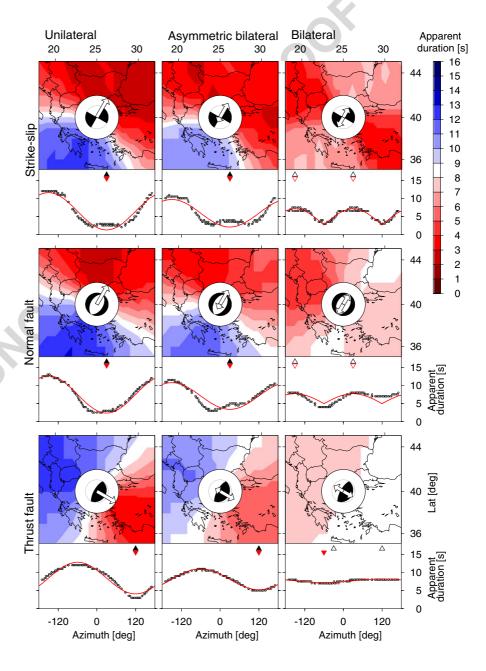
sizes are proportional to rupture lengths). Graphics below each colour plot represent the apparent duration versus azimuth (dots) and the best fitting model (red curves). Upper triangles represent the correct solution (a single black triangle is plotted at the proper azimuth for unilateral and asymmetric bilateral ruptures; two white triangles indicate rupture directions for pure bilateral rupture cases). Inverted triangles represent inversion results (single red triangle for unilateral and asymmetric bilateral ruptures, two white triangles for pure bilateral ruptures)



and above the centroid depth, respectively. Since we use a layered model, the frequency content of synthetic seismogram varies from shallower to deeper sources, as well as take off angles. These effects, which cannot be reproduced by a point source located at the centroid depth, results larger than those related to the bilateral rupture, thus explaining the detection of an apparent directivity towards dip direction.

Given the successful application to line sources, 659 we simulate now more realistic rupture processes, 660 generating synthetic seismograms for the eikonal 661 source model (Heimann 2010; Cesca et al. 2010). 662 We use here circular faults, with rupture propagating with a variable velocity, scaling by a coefficient 664 0.9 with shear wave velocity in the crustal model. 665 Given the adoption of the PREM model, the 666 source depth and its extension, rupture velocity 667

Fig. 4 Inversion results for circular eikonal sources. For each of the nine considered source models, we show colour plots representing the inverted apparent duration around the epicentre (red to blue scale represents increasingly longer apparent durations). Graphics below each coloured plot represent the apparent duration versus azimuth (dots) and the best fitting model (red curves). We use the same symbol convention as in Fig. 3





725

668 range between 2.9 and 4.0 km/s. Inversion is car-669 ried out using the same approach and parameters 670 as for the previous test and results summarized in 671 Fig. 4. The main characteristics of the azimuthal patterns of the apparent rupture duration are pre-673 served. Directivity effects can be detected and modelled for all pure unilateral and asymmetric bilateral ruptures, although the fit quality results 676 in some cases significantly poorer than for line source cases. The adoption of bi-dimensional rup-678 tures, and specifically the inclusion of sources at different depths, slightly modifies the apparent duration pattern, as can be seen by the comparison of plots in Figs. 3 and 4 relative to normal fault 682 mechanism. The modification of the apparent rupture radiation pattern has similar causes than those described for linear sources. Additionally, 685 it is here also depending on the variable rupture velocity (which, according to the crustal model, is faster for deeper sources than for shallower ones). These effects result critical for the inversion of a bilateral rupture for the case of a thrust 690 fault mechanism, which is erroneously interpreted as unilateral (with a rupture propagation in dip direction). On the other hand, the discrimination 693 between pure and asymmetric bilateral rupture is not always possible; the observation of the charac-695 teristic lobe associated to the asymmetric bilateral 696 rupture may indicate such a rupture process, but 697 could not be sufficient, alone, to distinguish this 698 case to a pure unilateral rupture. 699

A variation of the scalar moment will lead to a 700 scaling of synthetic seismograms and modify their amplitude spectra. As a consequence, the inver-702 sion of apparent durations may lead to slightly different results. In order to investigate these 704 effects, we have perturbed the scalar moments 705 used for synthetic tests, and analyse inversion 706 results. While the uni- or bilateral mode of the 707 rupture and the main rupture direction are not 708 influenced by a variation of the scalar moment and are always correctly retrieved, the rupture time 710 and the following estimation of rupture lengths 711 suffer slight changes. A perturbation of 10% of 712 the correct scalar moment always led to uncer-713 tainties below 5% in terms of rupture time and 714 rupture length. Synthetic tests suggest the imple-715 mentation of the method here proposed towards a 716 rapid detection of directivity effects. Additionally, these tests point out specific cases, where the in- 717 version approach results more critical, and where 718 a careful discussion of results is suggested. In gen-719 eral, horizontal ruptures and pure or partially uni- 720 lateral ruptures are more easily detected, whereas 721 pure bilateral sources and rupture propagating 722 along dipping directions may be more problematic 723 to resolve.

4 The Andravida 8.6.2008 earthquake

On June 8th, 2008, a magnitude Mw 6.4 726 earthquake struck NW Peloponnese, Greece. 727 The earthquake, here further referred as the 728 Andravida earthquake, produced two casual- 729 ties, about 100 injuries and several damages 730 (Chouliaras 2009). A wide number of studies cov-731 ered the earthquake source and its effects. A pure 732 strike-slip focal mechanisms was unanimously 733 provided by several institutions and catalogues 734 surveying regional and global seismicity, includ- 735 ing National Observatory in Athens (NOA), 736 Aristotle University of Thessaloniki (AUTH), 737 INGV European-Mediterranean RCMT Cata- 738 logue (INGV-RCMT), Swiss Federal Institute 739 of Technology, United States Geological Sur- 740 vey (USGS) and Global CMT Catalogue (CMT). 741 According to these models, fault planes are al- 742 most vertical and oriented NNE-SSW and WNW- 743 ESE. Source depth estimations showed some 744 variability, ranging between 10 and 38 km, and 745 magnitudes Mw ranged between 6.3 and 6.5. 746 The epicentral locations of the earthquake after- 747 shocks (Ganas et al. 2009; Gallovic et al. 2009; 748 Kostantinou et al. 2009), which are distributed 749 within a narrow strip extending NNE-SSW, pro- 750 vide a convincing image of the rupture orienta- 751 tion. The cloud of aftershocks elongates for about 752 30-35 km, providing a first rough estimation of 753 rupture size. The aftershock distribution is denser 754 towards the Northern edge. To the south, epi- 755 central locations may indicate a minor bending 756 of the rupture area to a slightly larger strike. 757 All published source models are consistent with 758 the identification of the NNE-SSW striking fault 759 plane. Sokos et al. (2008), using hypocentral- 760 centroid relative location method, identified the 761 same plane; since centroids locations are generally 762



763 located North of hypocentral locations, some indi-764 cation for a dominant propagation towards North 765 may arise from this study. Even more convincing, with respect to the detection of directivity, are the studies of Kostantinou et al. (2009), Gallovic et al. (2009) and Cesca et al. (2010). The first authors derived a finite source model, also consistent with their aftershock relocations, finding a rupture length of 22.4 km and an asymmetric bilateral rupture, with a major rupture along the 773 NE branch. Gallovic et al. (2009) used a conjugate gradient method to detect the spatio-temporal evolution of the rupture process. Results indicate a predominantly unilateral rupture, propagating along a main slip patch, with a rupture length 778 of about 20 km and a rupture velocity of about 779 3 km/s. Cesca et al. (2010), based on amplitude 780 spectra inversion of full waveform and assuming an eikonal source model, detected an asymmet- 781 ric rupture propagation, with a predominance of 782 rupture propagation towards NNE; the rupture 783 length was estimated 40 km, while the average 784 rupture velocity was fixed to about 3.2 km/s. The 785 consistency of these results (see Fig. 5 bottom), 786 obtained with different methods and datasets, 787 offer a serious reference to our study in terms of 788 fault plane identification and rupture directivity. 789

In our inversion, we assume the epicentral lo-790 cation provided on the EMSC-CSEM webpage 791 and the focal mechanism determined by Cesca 792 et al. (2010), which is based on the fit of full 793 waveform amplitude spectra. We observed that 794 a new estimation of the scalar moment, based 795 on the fit of P wave spectra only, would present 796 minor variation with respect to the assumed value 797 (5.97e18 instead of 6.07e18 Nm). Then, we invert 798

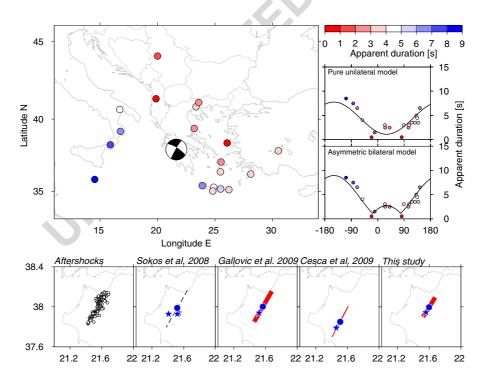


Fig. 5 Inversion results for the Mw 6.4 Andravida (NW Peloponnese) earthquake (top) and comparison with published source models (bottom). Top: coloured dots represent the inverted apparent duration at the stations used, according to the given colour scale; the azimuthal distribution of apparent durations may be fitted (top right) assuming a pure unilateral or, better, a partially unilateral rupture (thick lines). Bottom: the comparison of aftershocks distribution (after Gallovic et al. 2009) identifies

the NNE–SSW rupture plane in agreement with centroid-hypocentral technique (Sokos et al. 2008), adjoint method (Gallovic et al. 2009), full waveform kinematic inversion (Cesca et al. 2010) and our results; the last four methods consistently detect a partially unilateral rupture towards NNE. Stars and blue circles represent here nucleation points and centroids respectively; the rupture area is plotted in red



863

864

799 for the apparent duration at each station sep-800 arately and plot resulting values in function of 801 station azimuth (Fig. 5), according to the discussed 802 methodology. As a first approximation, we try to fit apparent durations by means of pure unilateral and pure bilateral rupture models, assuming both fault planes. Since fault planes are almost vertical, only horizontal directions of the rupture velocity along these planes are considered. The unilateral rupture model with rupture propagation towards NNE provide a very good fit to the apparent du-810 ration data, and is preferred to remaining models on the base of the F test. This result provides a clear indication for a rupture propagating towards 813 NNE, and thus can be used to discriminate the 814 true fault plane (NNE-SSW) from the auxiliary 815 one (WNW-ESE). Based on the good fit, we try 816 to refine our solution by investigating asymmetric 817 bilateral ruptures. Results provide an even more 818 convincing fit (Fig. 5, bottom right), when an asymmetric bilateral source model with rupture propagating mostly Northward is assumed, as the curve account for the two symmetric minima at about -21.5 and 82.5° and the internal characteristic lobe of asymmetric rupture (see Fig. 2, 824 bottom). On the other side, this result is in very good agreement with published models discussed 826 before. According to the previous discussion for a simplified bidimensional case, the maxima of the two cosine curves associated to the rupturing of two segments of the fault are equal to t_P + $t_{\rm R} + t_{\rm r}$, where these terms refer to the P wave propagation, rupture and rise time related to each segment. The maxima of the apparent duration curve are equal to 8.9 s (segment L₁ toward NNE) and 2.7 s (segment L₂ toward SSW). Assuming an average P wave velocity of 8 km/s (consistent with the used velocity model at the hypocentral depth), a rupture velocity of 3 km/s (consistent with Gallovic et al. 2009), and considering the rise time negligible with respect to rupture time, we obtain rupture lengths of about 19 and 6 km. The total length of about 25 km for the main patch is in general agreement with most of published results. We observe a discrepancy with the rupture size of about 40 km determined in Cesca et al. 845 (2010). This last value might be overestimated, 846 as the adopted full waveform kinematic inversion 847 may in some cases be affected by a trade-off between different source parameters. The observed 848 discrepancy may also indicate a different response 849 of the two approaches to the rupture process and 850 energy emission, with the full waveform inver- 851 sion detecting the largest rupture length, and the 852 directivity inversion identifying the main rupture 853 patch. The upper limit value we found here is 854 slightly larger than the length estimated by stan- 855 dard empirical relations (according to Wells and 856 Coppersmith 1994, the average rupture length for 857 a Mw 6.4 is about 14 km). We remark that the 858 interpretation of inversion results to this extent 859 should be carried out only in best conditions, 860 where the fitting of the apparent duration curve 861 is good enough to further interpret it.

5 The SW Peloponnese 14-20.2.2008 seismic sequence

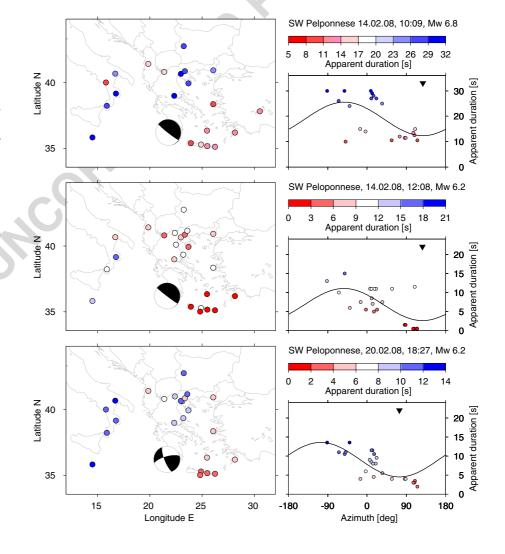
A seismic sequence struck the region offshore 865 SW Peloponnese, Greece, in the days following 866 February 14th, 2008, with three major earth- 867 quakes occurring within a week. On February 868 14th, a first Mw 6.8 (magnitude estimated by 869 EMSC-CSEM and Cesca et al. 2010) event struck 870 the region at 10:09 UTC. Two hours later, at 871 12:08, UTC, a Mw 6.2 aftershock occurred. Fi- 872 nally, on February 20th (18:27 UTC), a Mw 6.2 873 event took place. Focal mechanisms (EMSC- 874 CSEM webpage) indicate thrust faulting for the 875 first two events, while the last one has a different, 876 strike-slip mechanism, with fault planes striking 877 ENE and NNW. Source depths, according to 878 EMSC-CSEM catalogue, were 30, 20 and 25 km, 879 respectively, for the three earthquakes. Whereas 880 different institutions (e.g. NOA, AUTH, Uni- 881 versity of Patras UPSL, INGV-RCMT, USGS, 882 CMT) provided point source solutions for these 883 events, few trials has been carried out so far to 884 interpret rupture kinematics (Roumelioti et al. 885 2009; Cesca et al. 2010). A strongly uneven sta-886 tion distribution and large epicentral gaps toward 887 SW have possibly limited source modelling until 888 now. Based on full waveform inversion, Cesca 889 et al. (2010) identified the ENE-WSW plane for 890 the strike-slip earthquake of February 20th, and 891 a partial unilateral rupture towards the coast was 892 found. For the first two earthquakes, the low angle 893



planes dipping toward NW were preferred, but the inversion results were not completely satisfactorily. Roumelioti et al. (2009) adopted an empirical Green's functions approach, using the Mw 6.2 aftershock to model the finite fault of the largest earthquake; their results support the identification of the low angle dipping plane as well as directivity towards SSW. Finally, the identification of low dip angle rupture planes in this region for thrust earthquakes would agree with local tectonics, associating the earthquake occurrence to oceanic subduction (Underhill 1999).

906 Figure 6 summarizes our inversion results. 907 Differently from the application to the Andravida 908 earthquake, apparent durations for the earthquakes occurring on February 14th show a major 909 spreading and are worse fitted by the simplified 910 cosine function we associated to pure unilateral 911 ruptures. A better fit, and minor spreading, is 912 observed for the main event, with respect to its 913 aftershock. However, for both two earthquakes, a 914 general trend can be detected, indicating a minimum of the cosine curve for an azimuth of about 916 135°, which suggests a main direction of the rupture propagation towards SE. These results are 918 unable by themselves to provide further informations about the true fault plane, as they may be 920 modelled either assuming the low-angle and the 921 steep dipping plane. A better fit is obtained for the 922 February 20th earthquake: clear directivity effect

Fig. 6 Inversion results for the February 14th, 2008, Mw 6.8 (top), the February 14th, 2008, Mw 6.2 (centre) and the February 20th, 2005, Mw 6.2 (bottom) earthquakes, offshore SW Peloponnese. For each earthquake, coloured dots represent the inverted apparent duration at the stations used, according to the colour scale given for each case; the azimuthal distribution of apparent durations (right) may be fitted assuming unilateral ruptures (thick black lines). Inverted triangles indicate the retrieved rupture directions





970

980

981

990

991

997

998

1002

1003

924 is here retrieved, indicating a rupture mostly prop-925 agating towards ENE, thus along the WSW-ENE 926 fault plane. The differentiation between pure or partial unilateral rupture may be here rewarded as beneath the limit of a safe data interpretation, 928 but we observed how the general result is in well agreement with previous results by Cesca et al. 931 (2010), where the source model was derived by 932 the fit of high-frequency (up to 0.1 Hz) amplitude 933 spectra from the whole waveforms.

Finally, we investigate effects of the assumption 934 935 of imprecise point source parameters on the esti-936 mation of rupture directivity. The effects of anomalous source depth estimation are studied for the 938 February 20th aftershock, as different Institutions 939 have provided a range of different values ranging 940 from 8 to 25 km. We repeated the inversion using a different point source solution (strike 249°, dip 942 88°, rake -12° , depth 12 km), as provided by the 943 INGV European-Mediterranean RCMT Catalog; 944 this focal mechanism is similar to our solution and 945 major differences concern centroid depth, which is 946 now shallower. Directivity inversion remains very stable, showing a consistent identification of uni-948 lateral rupture direction toward SE, and indicates 949 that a source depth variation of about 10km does 950 not result in any significant variation in the radia-951 tion pattern of apparent duration. In a similar way, 952 we tested slightly different focal mechanisms for 953 both earthquakes of February 14th: even if our fo-954 cal mechanisms are in relatively good agreement 955 with other published solutions, some difference 956 can be observed. For example, the INGV-RCMT 957 catalogue indicates strike angles of 333° and 298°, 958 for these earthquakes, which differ from our so-959 lution (347° and 341° respectively). Even in this 960 case, after adopting the source parameters pro-961 vided by INGV-RCMT, the inversion results are 962 stable, with the detection of main rupture direc-963 tions pointing towards SE-ESE.

964 6 Conclusions

965 We propose here a new method for a quick detec-966 tion of directivity effects for shallow earthquakes 967 at regional distances. Among the most important features of the method, we highlight here the 968 following ones:

Rapid inversion

The assumption of spatial point source allows 971 an extremely rapid generation of synthetic seis- 972 mograms and point source parameters inversion, 973 thus offering a tool to early detect directivity; to 974 quantify such improvement, on a standard single 975 processor PC the inversion of directivity is here 976 carried out within a minute, about 20 times faster 977 than the full kinematic inversion for the same 978 event, using the approach described in Cesca et al. 979 (2010).

Coherent inversion

The use of the Kiwi tools for data processing 982 and inversions improves significantly the consis-983 tency of our methodology: for example, the same 984 dataset and the same inversion tools can be used 985 to first derive the focal mechanism, then the scalar 986 moment and finally the apparent duration; we 987 believe this consistency between data used for 988 different inversions significantly improve the co- 989 herency of the inversion approach.

Accounting for wave propagation

The method is based on amplitude spectra inver- 992 sion, using theoretical Green's functions for the 993 chosen earth model. In this way, we account for 994 wave propagation effect, an improvement with 995 respect to standard methods based on pulse length 996 estimations.

No requirements of specific aftershocks

Avoiding the use of empirical Green's function, 999 the method is not limited by the existence of 1000 a proper aftershock, nor need to wait for its 1001 occurrence.

Automation

The adoption of the Kiwi tools and the simplicity 1004 of the inversion approach made possible the im- 1005 plementation of the method as automated routine. 1006

In this manuscript we have demonstrated the 1007 method performance, both with a range of syn- 1008 thetic tests and with observed data for different 1009



1063

1066

1070

1075

1079

1082

1087

1092

1097

1098

1099

1102

1108

1113

1010 shallow earthquakes recently occurred. These ap-1011 plications offer indications about the quality and 1012 extent of inversion results. The retrieval of pure 1013 unilateral and pure bilateral ruptures is in gen-1014 eral better resolved than asymmetric ruptures, 1015 although the application to the June 8th, 2008, 1016 Andravida earthquake showed that this case can 1017 also be detected, in favourable conditions. In 1018 general, directivity effects are better resolved for 1019 strike slip earthquakes, with respect to normal or 1020 thrust faulting. Directivity detection offers often a 1021 chance to identify the rupture plane, discriminat-1022 ing it from the auxiliary one. The determination 1023 of rupture time, rise time and rupture velocity on 1024 the base of the proposed method is beyond its 1025 purposes and should require a careful supervision. 1026 We have here focused to earthquake with magni-1027 tudes of Mw 6 to 7 and shallow hypocentres. The 1028 extension of this inversion approach for the study 1029 of other range of magnitudes or deeper sources 1030 may be investigated in future.

1031 Acknowledgements We thank Prof. J. Zahradnik and 1032 two anonymous reviewers for useful comments and suggestions. The facilities of GEOFON and IRIS Data 1033 1034 Management System, and specifically the IRIS Data Man-1035 agement Center, were used for access part of the waveform 1036 and metadata required in this study. We acknowledge all 1037 institutions providing seismic data used in this research: 1038 GEOFON, MEDNET Project, Greek National Seismic 1039 Network and Aristotle University Thessaloniki Network. 1040 Maps and focal mechanisms have been plotted with GMT 1041 (Wessel and Smith 1998). This work has been funded by the 1042 German DFG project KINHERD (DA478/14-1/2) and the 1043 German BMBF/DFG "Geotechnologien" project RAPID 1044 (BMBF07/343).

1045 References

- 1046 Beck SL, Silver P, Wallace TC, James D (1995) Directivity 1047 analysis of the deep Bolivian earthquake of June 9, 1048 1994. Geophys Res Lett 22:2257–2260
- 1049 Ben-Menahem A (1961) Radiation of seismic surface 1050 waves from finite moving sources. Bull Seismol Soc 1051 Am 51:401-453
- 1052 Ben-Menahem A, Singh SJ (1981) Seismic waves and 1053 sources. Springer, New York
- 1054 Bernard P, Madariaga R (1984) A new asymptotic method 1055 for the modelling of near-field accelerograms. Bull 1056 Seismol Soc Am 74:539-557

- Beroza GC, Spudich P (1988) Linearized inversion for fault 1057 rupture behaviour: application to the 1984 Morgan 1058 Hill, California, earthquake. J Geophys Res 93:6275- 1059
- Boore D, Joyner W (1978) The influence of rupture inco- 1061 herence on seismic directivity. Bull Seismol Soc Am 1062 68:283-300
- Brüstle W, Müller G (1987) Stopping phases in seismo- 1064 grams and the spatiotemporal extent of earthquakes. 1065 Bull Seismol Soc Am 1:47-68
- Caldeira B, Bezzeghoud M, Borges JF (2009) DIRDOP: 1067 a directivity approach to determining the seismic rup- 1068 ture velocity vector. J Seismol 14:565–600. doi:10.1007/ 1069 s10950-009-9183-x
- Cassidy JF (1995) Rupture directivity and slip distrib- 1071 ution for the Ms 6.8 earthquake of 6 April 1992, 1072 Offshore British Columbia: an application of the em- 1073 pirical Green's function method using surface waves. 1074 Bull Seismol Soc Am 85:736-746
- Cesca S, Heimann S, Stammler K, Dahm T (2010) Auto- 1076 mated procedure for point and kinematic source inver- 1077 sion at regional distances. J Geophys Res 115:B06304. 1078 doi:10.1029/2009JB006450
- Chouliaras G (2009) Seismicity anomalies prior to 8 June 1080 2008, Mw 6.4 earthquake in Western Greece. Nat 1081 Hazards Earth Syst Sci 9:327–335
- Dahm T, Krüger F (1999) Higher-degree moment ten- 1083 sor inversion using far-field broad-band recordings: 1084 theory and evaluation of the method with application 1085 to the 1994 Bolivia deep earthquake. Geophys J Int 1086 137:35-50
- Dahm T, Kruger F, Stammler K, Klinge K, Kind R, 1088 Wylegalla K, Grasso JR (2007) The 2004 Mw 4.4 1089 Rotenburg, Northern Germany, Earthquake and its 1090 possible relationship with Gas Recovery. Bull Seismol 1091 Soc Am 97:691-704
- Dreger D, Kaverina A (2000) Seismic remote sensing 1093 for the earthquake source process and near-source 1094 strong shaking: a case study of the October 16, 1999 1095 Hector mine earthquake. Geophys Res Lett 27:1941– 1096
- Dziewonski AM, Anderson DL (1981) Preliminary reference earth model. Phys Earth Planet Int 25:297-356
- Eshghi S, Zare M (2003) Reconnaisance report on 26 1100 December 2003 Bam earthquake. International Insti- 1101 tute of Earthquake Engineering (IIEES)
- Gallovic F, Zahradnik J, Krizova D, Plicka V, Sokos 1103 E, Serpetsidaki A, Tselentis GA (2009) From earth- 1104 quake centroid to spatial-temporal rupture evolution: 1105 Mw 6.3 Movri Mountain earthquake, June 8, 2008, 1106 Greece. Geophys Res Lett 36:L21310. doi:10.1029/ 1107 2009GL040283
- Ganas A, Serpelloni E, Drakatos G, Kolligri M, Adamis I, 1109 Tsimi C, Batsi E (2009) The Mw 6.4 SW-Achaia (west-1110 ern Greece) earthquake of 8 June 2008: seismological, 1111 field, GPS, observations and stress modeling. J Earthq 1112 Eng 8:1101-1124
- Hartzell SH (1978) Earthquake aftershocks as Green's 1114 functions. Geophys Res Lett 5:1-4 1115



1205

1209

- 1116 Hartzell S, Heaton DV (1983) Inversion of strong ground 1117 motion and teleseismic waveform data for the fault 1118 rupture history of the 1979 Imperial Valley, California, 1119 earthquake. Bull Seismol Soc Am 83:1553-1583
- 1120 Hartzell S, Helmberger DV (1982) Strong-motion mod-1121 elling of the Imperial Valley earthquake of 1979. Bull 1122 Seismol Soc Am 72:571-596
- 1123 Haskell NA (1964) Total energy and energy spectral den-1124 sity of elastic wave radiation from propagating faults. 1125 Bull Seismol Soc Am 54:1811-1841
- 1126 Heimann S (2010) A robust method to estimate kinematic 1127 earthquake source parameters. PhD Thesis, Univer-1128 sity of Hamburg, Germany, pp 145
- 1129 Imanishi K, Takeo M (1998) Estimates of fault dimensions 1130 for small earthquakes using stopping phases. Geophys 1131 Res Lett 25:2897-2900
- 1132 Imanishi K, Takeo M (2002) An inversion method to 1133 analyze rupture process of small earthquakes using 1134 stopping phases. J Geophys Res 107:ESE2.1–ESE2.16. 1135 doi:10.1029/2001JB000201
- 1136 Kostantinou KI, Melis NS, Lee SJ, Evangelidis CP, 1137 Boukouras K (2009) Rupture process and aftershock 1138 relocation of the 8 June 2008 Mw 6.4 earthquake 1139 in Northwest Peloponnese, Western Greece. Bull 1140 Seismol Soc Am 99:3374–3389
- 1141 Li Y, Toksöz MN (1993) Study of the source process of 1142 the 1992 Columbia Ms = 7.3 earthquake with the empirical Green's function method. Geophys Res Lett 1143 1144 20:1087-1090
- 1145 Madariaga R (1977) High-frequency radiation from crack 1146 (stress drop) models of earthquake faulting. Geophys 1147 J R Astron Soc 51:625-651
- 1148 Madariaga R (1983) High-frequency radiation from dy-1149 namic earthquake fault models. Ann Geophys 1:17-23
- 1150 McGuire JJ, Zhao L, Jordan TH (2001) Teleseismic inversion for the second-degree moments of earthquake 1151 1152 space-time distributions. Geophys J Int 145:661-678 1153
- 1154 McGuire JJ, Zhao L, Jordan TH (2002) Predominance of 1155 unilateral rupture for a global catalog of large earth-1156 quakes. Bull Seismol Soc Am 92:3309-3317
- 1157 Müller CS (1985) Source pulse enhancement by deconvolu-1158 tion of empirical Green's functions. Geophys Res Lett 1159 12:33-36
- 1160 Nadim F, Moghtaderi-Zadeh M, Lindholm C, Andresen A, Remseth S, Bolourchi MJ, Mokhtari M, Tvedt T 1161 1162 (2004) The Bam earthquake of 26 December 2003. 1163 Bull Earthquake Eng 2:119–153
- 1164 Nielsen S, Madariaga R (2003) On the self-healing fracture 1165 mode. Bull Seismol Soc Am 93:2375-2388
- 1166 Olson AJ, Apsel RJ (1982) Finite faults and inverse theory 1167 with applications to the 1979 Imperial Valley earth-1168 quake. Bull Seismol Soc Am 72:1969–2001

- Pro C, Buforn E, Udías A (2007) Rupture length and 1169 velocity for earthquakes in the Mid-Atlantic Ridge 1170 from directivity effects in body and surface waves. 1171 Tectonophysics 433:65–79
- Roumelioti Z, Benetatos C, Kiratzi A (2009) the 14 1173 February 2008 earthquake (M6.7) sequence offshore 1174 south Peloponnese (Greece): source models of the 1175 three strongest events. Tectonophysics 471:272–284
- Selby N, Eshun E, Patton H, Douglas A (2005) Un- 1177 usual long-period Rayleigh wave from a vertical dip- 1178 slip source: the 7 May 2001, North Sea earthquake. 1179 J Geophys Res 110:B10304. doi:10.1029/2005JB003721 1180
- Sokos E, Serpetsidaki A, Tselentis GA, Zahradnik J (2008) 1181 Quick assessment of the fault plane, for the recent 1182 strike-slip event in the North-Western Peloponnese, 1183 Greece, (8 June 2008, Mw 6.3). EMSC-CSEM Report 1184
- Spudich P, Frazer LN (1984) Use of ray theory to calcu- 1185 late high-frequency radiation from earthquake sources 1186 having spatially variable rupture velocity and stress 1187 drop. Bull Seismol Soc Am 74:2061–2082 1188
- Underhill JR (1999) Late Cenozoic deformation of the 1189 Hellenide forelands, Western Greece. Geol Soc Am 1190 Bull 101:613-634 1191
- Vallée M (2007) Rupture properties of the giant suma- 1192 tra earthquake imaged by empirical Green's function 1193 analysis. Bull Seismol Soc Am 97:103-114 1194
- Vallée M, Bouchon M (2004) Imaging coseismic rupture in 1195 the far field by slip patches. Geophys J Int 156:615–630 1196
- Velasco AA, Ammon CJ, Lay T (1994) Empirical green 1197 function deconvolution of broadband surface waves: 1198 Rupture directivity of the 1992 Landers, California 1199 (Mw = 7.3), earthquake. Bull Seismol Soc Am 84:735–1200
- Velasco AA, Ammon CJ, Farrell J, Pankow K (2004) 1202 Rupture directivity of the 3 November 2002 denali 1203 fault earthquake determined from surface waves. Bull 1204 Seismol Soc Am 94:293-299
- Warren LM, Shearer PM (2006) Systematic determination 1206 of earthquake rupture directivity and fault planes from 1207 analysis of long-period P-wave spectra. Geophys J Int 1208 164:46-62
- Wells DL, Coppersmith KJ (1994) New empirical relations 1210 among magnitude, rupture length, rupture width, rup- 1211 ture area and surface displacements. Bull Seismol Soc 1212 Am 84:974-1002
- Wessel P, Smith WHF (1998) New improved version of 1214 the generic mapping tools released. Eos Trans AGU 1215 79:579 1216
- Zahradnik J, Gallovic F, Sokos E, Serpetsidaki A, 1217 Tselentis GA (2008) Quick fault-plane identification 1218 by geometrical method: application to Mw 6.2 Leoni- 1219 dio earthquake, January 6, 2008, Greece. Seismol Res 1220 Lett 79:653-662



AUTHOR QUERY

NO QUERY

