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Interpretation of the microwave non-thermal radiation of the Moon during impact events

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Abstract. The results of recent observations of the non-thermal electromagnetic (EM) emission at wavelengths of 2.5 cm, 13 cm, and 21 cm are summarized. After strong impacts of meteorites or spacecrafts (Lunar Prospector) with the Moon’s surface, the radio emissions in various frequency ranges were recorded. The most distinctive phenomenon is the appearance of quasi-periodic oscillations with amplitudes of 3–10 K during several hours. The mechanism concerning the EM emission from a propagating crack within a piezoelectric dielectric medium is considered. The impact may cause the global acoustic oscillations of the Moon. These oscillations lead to the crackening of the Moon’s surface. The propagation of a crack within a piezoelectric medium is accompanied by the excitation of an alternative current source. It is revealed that the source of the EM emission is the effective transient magnetization that appears in the case of a moving crack in piezoelectrics. The moving crack creates additional non-stationary local mechanical stresses around the apex of the crack, which generate the non-stationary electromagnetic field. For the cracks with a length of 0.1–1 µm, the maximum of the EM emission may be in the 1–10 GHz range.

1 Introduction

The prediction of earthquakes is an important problem, which is still far from a final solution. One of the possible methods for recognition of precursors of earthquakes is the investigation of the electromagnetic (EM) radiation of seismic origin. The EM radiation before earthquakes has been recorded in the kHz and MHz frequency ranges (Gokhberg et al., 1995). Also, the EM emission was measured in laboratory rock fracture experiments (Gershenzon et al., 1986). There are several probable models considering how the mechanic stress can be transformed into the EM radiation:

1. formation of new micro cracks in the rock,
2. charges arising at the peaks of existing cracks formed under the action of an increasing load,
3. piezoelectric and piezomagnetic effects.

The source of the EM radiation is located near the epicenter of an earthquake with depths of about 10–30 km. Thus, only the low frequency part of the total EM radiation can reach the Earth’s surface, because of the finite conductivity of the Earth’s crust, and in fact the total spectrum of the EM emission cannot be measured.

However, the non-thermal origin of the EM radiation after collisions of man-made and natural objects with the Moon has been measured. In this case, the source of the radiation is located directly on the Moon’s surface, and there exists the possibility of recording the total spectrum of the EM emission. Therefore, the investigation of non-thermal emissions from the surface of the Moon may give some insights on the specific mechanisms of the EM radiation of seismic origin in the Earth.

2 Review of radio observations of the Moon

The main source of lunar radio emission at centimeter wavelengths is the thermal emission of a brightness temperature equal to 250 K. The intensity of the thermal radio emission changes slowly with a period equal to one month. However, after strong impacts of meteorites or spacecrafts with the Moon’s surface, the quick changes of the intensity of the lunar radio emission in various frequency ranges were detected. The collision between the American spacecraft Lunar...
The high kinetic energy collisions can excite seismic waves in the lunar body. In the morning of 17 November 2000, the changes of the received lunar radio signal gradually continued for 5 h (see Fig. 1). There were at least two isolated events with valley-to-peak amplitudes of the oscillations of the antenna temperature exceeding 10 K. On 18 November 2000, the valley-to-peak amplitudes of the oscillations reached 10 K. Most of the quasi-periodic oscillations had periods from 1 to 6 min. Observed valley-to-peak oscillation amplitudes in November 2001, were larger than that in November 2000. The valley-to-peak amplitudes of quasi-periodic oscillations reached 25 K on 18 and 19 November 2001. The most frequent periods found were 2.5, 3.0, 3.7, and 6.0 min. Those periods were similar to the periods detected during the Leonid 2000 meteor shower. The times of maxima of the amplitudes of the variations of the lunar radio flux agree well with predicted times of maxima of the Leonid activity on the Moon. This may be evidence that the amplitude of the variations of the lunar radio flux increases with the increasing Leonid activity on the Moon, if the instrumental effects are negligible.

A study of fluctuations of the lunar radio flux under the Lyrid shower was carried out at 1.3 cm, 6 cm, and 18 cm on 16-20 April 2001 (Berezhnoy et al., 2002). The bandwidth was $\Delta f=150$ MHz at 1.3 cm and $\Delta f=6$ MHz at 6 cm and 18 cm. An important feature is the periodic variations of the lunar radio flux. Periods from 2 to 15 min were similar to those at other wavelengths. The amplitude of the variations of lunar flux was about 1–10 K for all wavelengths.

All these observations of the Moon were conducted with the same antenna. This means that the detected variations of the lunar flux could be caused by either atmospheric turbulence or an instability of the receiver. In order to study the possible role of instrumental effects, simultaneous observations of the center of the Moon at 6.2 cm were conducted at Pushchino radio telescope (Russia) and Simeiz radio telescope (Ukraine) during 17–19 November 2001. These observations are important because they were conducted simultaneously at similar frequencies as the Ventspils observations.

The frequency and the bandwidth of the receiver at Simeiz are 4.866 GHz and 2 MHz, respectively. The frequency and bandwidth of the receiver at Pushchino are 4.830 GHz and 6 MHz, respectively. The diameter of the antennas of both radio telescopes is 22 m. No correlation between fluctuations of the lunar radio flux at 4.866 GHz and 4.83 GHz was found. Thus, the detected periodicities of the lunar flux at other wavelengths should be checked again by simultaneous observations at two isolated antennas. This means that the intensity of the lunar radio emission of seismic origin is less than the sensitivity of both telescopes (3 K or 5000 Jy at 6.2 cm). We cannot estimate the upper level of the intensity of the lunar radio emission of seismic origin at 2.46 cm, because the lunar radio flux was not calibrated at these observations. Radio flashes caused by collisions of kg meteoroids with the Moon were not detected in the telescopes (Volvach et al., 2004).

Notice that the absolute values of the lunar radio brightness temperature were not measured during these observations. The calibration of the lunar radio flux was conducted.
every 30–60 min. For this reason, only the information concerning the temporal variability of the lunar radio flux at 1 s time resolution was obtained. The values of the observed periods of the lunar radio flux at 2.46 cm and 6.2 cm during the Leonid meteor shower in 2001 were different. This means that the excitations of free oscillations of the Moon cannot explain the periodicities of the lunar radio flux. The Leonid meteor shower activity shows the periodicities in the minute range (Singer et al., 2000). Probably, the periodicities of the Leonid meteor shower activity may be responsible for the appearance of periodicities in the lunar radio flux.

3  Piezoelectric mechanism of electromagnetic emission

In this section we will propose a mechanism explaining the origin of the microwave lunar radio emission. The absolute values of intensity of the radio emission produced by this mechanism can also be estimated. The observed EM emission can be explained by various mechanisms, such as the motion of charged dislocations, piezoeffect, and plasma emission. In this work, the mechanism concerning the EM emission from a propagating crack within a piezoactive dielectric medium is considered. The impact can cause global acoustic oscillations of the Moon. These oscillations can lead to cleaving mechanically weak rocks on the whole Moon’s body, and the shock necessarily leads to the cracking of a local part of its surface. The piezoelectric SiO$_2$ is a constituent of the lunar minerals; therefore, it can radiate after cracking.

We are considering a piezoelectric plate to which a mechanical stress of a value $P_0$ is applied along the OY-axis and normal to the optical OZ-axis, as shown in Fig. 2. For our estimations, all sizes of the plate ($L$, $H$, $h$) are assumed to be equal and we take quartz (SiO$_2$) as a piezoelectric material. The piezoefffect produces an electric field along the piezoaxes OX and OY. It is important that the fracturing of crystals, known as cleavage (Shuvalev et al., 1981), can occur in distinctive crystallographic planes. According to this, the crack caused by $P_0$ can expand along the axis OX, as shown in the Fig. 2. When the crack exists inside the plate, it begins to move through the crystal only if the applied mechanical stress exceeds certain value. The velocity of the crack in quartz is not greater than the velocity of the transverse acoustic waves. It is necessary to analyze the non-stationary problem because of the following reasons. First of all, an EM emission could exist as a result of the relaxation of the electron system, on the created surfaces in any materials. However, the time of this relaxation is very short and the frequency of the corresponding EM is very high. Also, the motion of the unloading wave from the crack occurs, and non-stationary local mechanical stresses around the crack in the piezoelectric appear.

The equations of motion are obtained by using a free energy $\Psi$ of a crystal. The free energy $\Psi=\Psi(u_{kl}, T, E_k, B_k, \alpha_1, ..., \alpha_N)$ is a function of the deformation tensor $u_{kl}$, temperature $T$, external electric and magnetic fields $E$, $B$, and, perhaps, of some $N$ internal parameters $\alpha_k$. If a dielectric is under the influence of small deformations and small electric fields, the equations of motion can be derived in the linear theory approach. This assumption gives us the possibility to analyze the emission phenomenon in a linear approach. We use the Maxwell’s equations for the description of electromagnetic waves

$$\nabla \times E = -\frac{1}{c} \frac{\partial D}{\partial t}, \nabla \times H = 0,$$  

$$\nabla \times D = \rho,$$  

and the equations of motion for a piezoelectric system

$$\frac{\rho}{c^2} \frac{\partial^2 U_i}{\partial t^2} = \frac{\partial \sigma_{ij}}{\partial x_j},$$

$$\sigma_{ij} = c_{ijkl} U_{kl} - e_{kij} E_k,$$  

$$D_i = \varepsilon_{ij} E_j + e_{ijk} U_{jk},$$

where $D$ is the electric induction, $c$ is the velocity of light, $\sigma_{ij}$, $U_{ij}$ are the components of stresses and mechanic displacement, respectively; $c_{ijkl}$, $e_{kij}$, $\varepsilon_{ij}$ are the components of the tensors of the elastic constants, the piezoelectric constants, and the dielectric permittivity, respectively (Makarets et al., 2002).

Substituting $D$ from the Eqs. (2) into the Maxwell’s equations and taking into account the definition of the linear deformation tensor $U_{ij}=(1/2) (\partial U_i / \partial x_j + \partial U_j / \partial x_i)$, one can obtain the wave equation for the components of the electric field $E_i(r, t)$, which can be presented as:

$$\frac{\partial^2 E_i}{\partial x_i} - \nabla \times \nabla \times E = \frac{1}{c^2} \frac{\partial^2 E_i}{\partial t^2} + \frac{1}{2c^2} e_{kl} \left( \frac{\partial}{\partial x_l} \frac{\partial^2 U_k}{\partial t^2} + \frac{\partial}{\partial x_k} \frac{\partial^2 U_l}{\partial t^2} \right) = S_i(r, t).$$

The right hand side $S(r, t)$ of the Eq. (3) is the source of the EM emission, and it is connected with the mechanical displacement vector $U$. As our calculations have shown (Makarets et al., 2002), this source is equivalent to the effective transient magnetization vector under the above mentioned conditions.

It is possible to analyze three basic cases of fracturing (see Fig. 3), which determine three configurations of the mechanical displacement vector $U$. In the static case, if the crack

![Fig. 2. Model of a moving crack.](image-url)
Consider the possible sources of the microwave radiation.

The micro cracks (L \sim 1 \mu m) can be considered as possible sources of the microwave radiation.

4 Frequency dependence of the electromagnetic emission

The dependence of the intensity of the radio emission of a seismic origin versus frequency was measured before strong earthquakes in Greece (Eftaxias et al., 2003). Based on observations in the kHz and MHz frequency ranges, it was established that the intensity of the radio emission of seismic origin increased as the square of the wavelength. This dependence may not be so strong in the case of meteoroid impacts on the Moon due to the absence of the skin effect. Unfortunately, based on the performed radio observations of the Moon, it is impossible to estimate the frequency of the maximum intensity of the lunar radio emission of seismic origin. In Makarets et al. (2002) it was shown that the calculated intensity of the EM emission from the cracks possesses an infinite number of periodically located maxima, and its first maximum value appears at the frequency \( \omega_{\text{max}} \approx \pi (V/L) \), and the corresponding wavelength is \( \lambda_{\text{max}} \approx 2L (c/V)(\varepsilon_{11})^{-1/2} \). The next maxima of the intensity of the EM radiation correspond to shorter waves, but they have no physical sense. An explanation is as follows. In simulations (Makarets et al., 2002), it was assumed that the velocity \( V \) of the moving crack is constant: \( V = V_0 \) during a time \( T = L/V_0 \). Therefore, at the beginning of the motion and at its end, the acceleration is infinite. A more realistic model for the frequency dependence of the intensity of the EM radiation should be applied. Namely, the acceleration of the crack can be approximated by a part of a sinusoid:

\[
\frac{dV}{dt} = a_0 \sin(\omega_0 t), \quad 0 < t < T; \quad \omega_0 = 2\pi/T; \quad V_{\text{max}} = V_0; \quad T = 2V_0/L.
\]

or step-like:

\[
\frac{dV}{dt} = \begin{cases} a_0, & 0 < t < T/2, \\ -a_0, & T/2 < t < T. \end{cases}
\]

A comparison of the dependencies of the intensity of the EM radiation on frequency for various cases of approximation of the velocity of the moving crack is presented in Fig. 4. One can see that the curves near the first (real) maximum are the same for all kinds of approximations. The corresponding value of the frequency is \( \omega_{\text{max}} \approx \omega_0 \approx \pi V_0/L \). The nonsingular approximations also give the same result for the frequency dependence at higher frequencies. If an approximation of the
velocity of the moving crack is discontinuous, then the non-
physical higher maxima possess the same value as the first
true maximum. Therefore, we need to use here the smooth
sine-like approximation for the velocity (Eq. 4). The value
of the maximum velocity $V_0$ is $V_0=2 \text{ km/s}$.

The frequency dependencies of the intensity of the EM ra-
diation for the single micro cracks of different lengths are
presented in Fig. 5. They are plotted for the following con-
ditions: $i$) the direction of the EM emission is $\theta=\psi=\pi/2$, $ii$) the crack-type is III. The angular distribution of the radi-
ation possesses a sharp orientation in the piezactive plane
$\theta=\pi/2$. It should be noted that under real conditions the ampli-
tudes of the maximum are sensitive to the crystal shape and,
therefore, some directions of the maximum EM radia-
tion can be changed. One can see that under the motion of
micro cracks the maximum of the EM radiation is in the mi-
crowave range.

It should be noted that the low frequency EM emissions
can be registered during volcano eruptions, when a fracture
is located at the surface of the volcano. In the case of a deep
fracture, the EM waves at higher frequencies dissipate, and
only ELF (extremely low frequency) and VLF (very low fre-
quency) waves are registered at the Earth’s surface. Hence,
it is necessary to use very sensitive sensors to register the
higher frequencies, because the high frequency EM waves
attenuate very quickly with distance in the lithosphere.

**5 Conclusions**

The frequency range of the EM radiation due to crackening in
piezorocks, which may be recorded experimentally, depends
on the depth of the localization of the source. If the source
is located within the depth of the planetary crust, the higher

**Fig. 4.** Frequency dependencies of the EM intensity emitted in the
direction $\theta=\psi=\pi/2$ by a crack of the type III for various accel-
erations of the crack. 1 is for sine-like approximation (Eq. 4), 2 is
for step-like approximation (Eq. 5), 3 is for the approximation
with constant velocity (singular acceleration). The final length of
the moving crack is $L=0.1 \mu\text{m}$.

**Fig. 5.** Frequency dependencies of EM intensity for non-singular
dependence of the velocity of the crack on time for cracks with
different lengths: 1 is for the length of the crack $L=0.1 \mu\text{m}$, 2 is for
$L=0.3 \mu\text{m}$, 3 is for $L=1 \mu\text{m}$.

frequency components are suppressed during their propaga-
tion, due to the skin effect. However, when the source is
located at the surface of a planet, all the spectral compo-
ents of EM radiation may be measured. Therefore, the pos-
sible origin of non-thermal emission from the Moon’s sur-
face may be explained by the EM emission under micro frac-
turing of piezorocks during the propagation of the acoustic
surface waves. For the cracks of lengths $L=0.1 \mu\text{m}–10 \mu\text{m}$,
the maximum intensity of the EM emission takes place at
$\sim0.1–10 \text{ GHz}$. But the tail of the emission spreads up to $10–
30 \text{ GHz}$.

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