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Cosmic ray-induced stratospheric aerosols: A possible connection to polar ozone depletions

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Abstract. The model calculations of altitude distribution of CN (condensation nuclei), plausible centers of sulfate aerosol formation after the occurrence of GLE, are presented. Events with relativistic solar protons (i.e. protons with energies >450 MeV) are observed at ground level by neutron monitors and called ground-level events (GLEs) (Shea and Smart, 2001). Analysis of experimental data and model calculations permits us to explain some distinctions observed in ozone total content (OTC) variations during several GLEs. For example, model simulations show a significant CN concentration enhancement during the May 1990 GLEs of relatively “moderate” magnitude, when polar ozone “mini-holes” (OTC depletions up to 20%) have been observed, while no OTC variations and considerable aerosol enhancements were seen during more powerful GLEs (4 August 1972, 2 May 1998, 14 July 2000) (Reagan et al., 1981; Shumilov et al., 1995, 2003). Our results demonstrate that “moderate” GLEs may increase aerosol content significantly and cause ozone “mini-hole” creation.

Keywords. Atmospheric composition and structure (aerosols and particles) – Solar physics, astrophysics and astronomy (energetic particles) – Space plasma physics (ionization processes)

1 Introduction

Since the discovery of the Antarctic ozone “hole” by Farman et al. (1985) a number of reports concerning ozone research at high latitudes have appeared (Solomon, 1988; Evans, 1990; Larsen and Henriksen, 1990; Taalas and Kyrö, 1992). According to these reports no ozone decreases comparable with that in the Antarctic have been observed in the Northern Hemisphere, and only small (up to 30%) and short-lived (several days) ozone total content depletions or so-called ozone “mini-holes” have been observed in the Arctic (Solomon,

1988; Evans, 1990; Taalas and Kyrö, 1992). The main sources of Arctic ozone “mini-holes”, as well as Antarctic ozone “hole” creation, seem to be polar stratospheric clouds (PSC) and sulfate aerosols, and as a consequence the heterogeneous reactions on the surface of aerosol particles (Solomon, 1988; Larsen and Henriksen, 1990). Dynamical processes in the stratosphere have also been considered as another reason for ozone “mini-hole” creation (Taalas and Kyrö, 1992). The main sources of sulfate aerosols in the atmosphere are volcanic eruptions, anthropogenic emissions and biological activity. The contribution of each component to aerosol layer formation in the two stratosphere differs in two hemispheres and depends on latitude. Additional aerosols can also be created in the polar atmosphere through the ion nucleation mechanism. The mechanism of the ion nucleation through the ionization process has been developed by Arnold (1982) as a potential source of condensation nuclei (CN) on which stratospheric sulfate aerosols can form (Arnold, 1982; Hofman and Rosen, 1983).

Shumilov et al. (1995, 1996) reported that high-energy (150–300 MeV) solar protons can contribute to aerosol layer and initiate ozone “mini-holes” at high latitudes. Taking into account that the low (less than -78°) stratospheric temperatures are rare in occurrence at northern polar latitudes (10% days in a year) (Taalas and Kyrö, 1992), we consider the solar protons as one of the possible agents influencing the ozone layer and aerosols in Arctic.

In the present paper the results of model calculations of aerosol enhancement after some GLEs in the frame of the ion nucleation mechanism are presented.

2 Model simulations of stratospheric aerosol increases after occurrence of GLEs

The main sources of sulfate aerosols in the atmosphere are volcanic eruptions, anthropogenic emissions and phytoplankton activity. Besides these, another factor responsible for aerosol creation in the atmosphere seems to exist. This

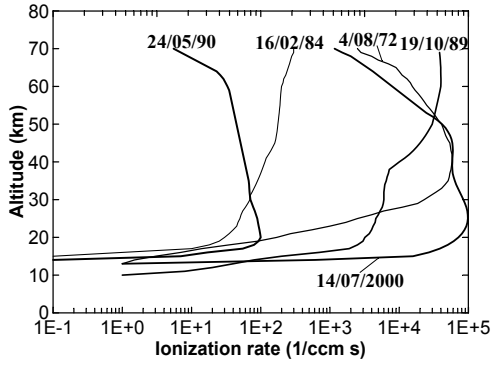


Fig. 1. Ionization rate profiles calculated for GLE events at the time of their maximum.

factor is a ion nucleation process initiated by high-energy solar (SCR) and galactic cosmic rays (GCR) penetrating below the 30-km altitude (Arnold, 1982; Hofman and Rosen, 1983; Yu and Turco, 2000; Yu, 2002). The process implied in the ion nucleation is the growth of ions by association of sulfuric acid molecules to a critical size at which the resulting molecular cluster would be stable against ion-ion recombination.

The rate of ion nucleation is related to the ion production rate (Arnold, 1982; Hofman and Rosen, 1983) and also depends on sulfuric acid vapor concentration $[H_2SO_4]$, temperature, relative humidity, and air pressure (Yu, 2002). A detailed scheme illustrating the ion-catalyzed nucleation of new ultrafine particles from trace condensable vapors (H_2SO_4 and H_2O) in the atmosphere, which may then grow into new cloud condensation nuclei (CCN), is performed in Carslaw et al. (2002); Yu, (2002). We propose that ion nucleation and aerosol enhancement may play a significant role in ozone “minihole” creation at high latitudes after GLEs (Shumilov et al., 1995, 2003). This result seems to correlate with the detection of large increases of stratospheric aerosols in the Arctic (~ 17 km altitude) after the occurrence of GLEs (Shumilov et al., 1996).

Figure 1 shows ionization profiles calculated for maximum intensity of some GLEs (4 August 1972, 16 February 1984, 19 October 1989, 24 May 1990 and 14 July 2000). For the ion production rate Q ($cm^{-3}s^{-1}$) calculation the following equation has been used:

$$Q(h) = \frac{1}{\varepsilon} \int_0^{\infty} \int_{\varphi=0}^{2\pi} \int_{\theta=0}^{\pi/2} D(E) E_i(E) \sin \theta d\theta d\varphi dE, \quad (1)$$

where $E_i(E)$ is the energy loss in the i -th altitude slab in MeV, $\varepsilon=35$ eV is the energy for one ion-electron pair production, $D(E)$ is the differential energy spectrum of incident solar protons, φ is the azimuthal angle and θ is the polar angle.

For calculations we used real GOES solar proton data that allowed us to evaluate a real differential spectrum in an energy interval from 1 to 850 MeV (<http://spidr.ngdc.noaa.gov/spidr/>). Spectra for 4 August 1972 and 16 February 1984

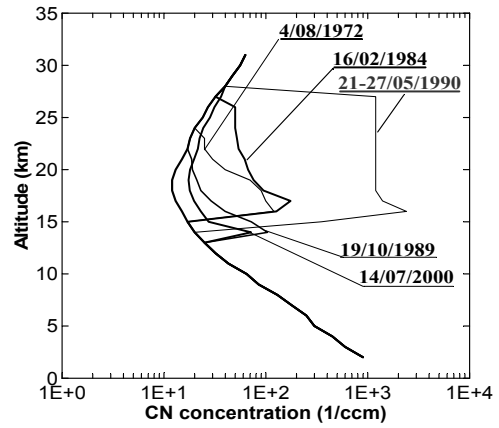


Fig. 2. Model simulations of condensation nuclei (CN) initiated by GLEs.

were estimated using solar proton parameters taken from Reagan et al. (1981) and Hargreaves et al. (1987), respectively.

To calculate aerosol enhancement during GLE events considered we used the ion nucleation theory according to which aerosol particles are formed by ions and H_2SO_4 molecules (Arnold, 1982). The ion nucleation rate J ($cm^{-3}s^{-1}$) is related to ion production rate Q by equation (Hofman and Rosen, 1983):

$$J = Q \cdot (1 + (\alpha Q)^{1/2} / K [H_2SO_4])^{-N_c}, \quad (2)$$

where α is the ion recombination coefficient ($10^{-7} cm^3 s^{-1}$), K is the association rate coefficient ($10^{-9} cm^2 s^{-1}$), $[H_2SO_4]$ is the molecular concentration of sulfuric acid, and N_c is the number of H_2SO_4 in a critically sized embryo. In the stratosphere only H_2SO_4 and water molecules can form stable clusters, and N_c is equal to 3 ($h \leq 27$ km) and 10 ($h > 27$ km) respectively (Arnold, 1982). Sulfuric acid concentration profiles were taken from Arnold (1982), with a maximum of $10^5 cm^{-3}$ between $h=15$ km and $h=27$ km. Note that additional sources, such as volcanic eruptions and anthropogenic emissions, can increase the $[H_2SO_4]$ level at high latitudes by two orders (Hofman and Rosen, 1983; Langner et al., 1992). The eruption of El Chichon on 14 April 1984, for example, has created a high level of sulfuric acid supersaturation ($> 10^7$ molecules cm^{-3}) (Hofman and Rosen, 1983). Typical sulfuric acid vapor concentrations according to some measurements and estimates lie in the range of perhaps $10^6 - 10^7 cm^{-3}$ (Eisele and Tanner, 1993; Weber et al., 1996).

We assumed that nucleation event durations roughly coincide to solar proton event periods (113 h for May 1990 GLEs, 4 h for 16 February 1984 GLE and 4 August 1972 GLE, 10 h for 19 October 1989 GLE; 14 July 2000 GLE). Of course, we have to take into account condensation and coagulation processes. But it's only a simple model simulation (most likely some physical estimates). There are also other nucleation processes in the troposphere and stratosphere, such as binary

homogeneous nucleation of sulfuric acid (H_2SO_4) and water vapor (H_2O) and heteromolecular condensation of (H_2SO_4) on the particles. These processes, as well as some other ones, including coagulation and sedimentation, are taken into account in more complicated models (Timmreck and Graf, 2000), and their application is beyond the scope of the paper.

Figure 2 demonstrates condensation nuclei (CN) profiles calculated using Eq. (2) for GLEs shown in Fig. 1.

3 Discussion and conclusions

As illustrated in Fig. 1, the vertical profiles shown have considerable discrepancies above 20 km: ionization rates on 14 July 2000, 4 August 1972 and 19 October 1989 exceed other two by two orders of magnitude. These discrepancies are connected to higher proton fluxes during 4 August 1972, 19 October 1989 and 14 July 2000 GLE events. We also may add that no OTC depletions have been observed after these events (Shumilov et al., 2003). On the other hand, the two largest GLEs (19 October 1989 and 14 July 2000) caused significant NO_x productions and ozone decreases at stratospheric altitudes on the data of rocket UARS spacecraft measurements, and also according to the modelling based on the ordinary gas-phase photochemistry (Zadorozhny et al., 1992; Jackman et al., 2001; Krivolutsky, 2001, 2003; Shumilov et al., 2003).

Previously it was shown (Shumilov et al., 1995) that OTC depletions observed during 29 September 1989 and May 1990 GLEs could not be explained using only the homogeneous photochemical model introduced by Crutzen et al. (1975). Our calculations of OTC depletions based on additional HO_x and NO_x production during solar proton events (Crutzen et al., 1975; Jackman et al., 1980) were one order of magnitude less than the observed ones (Shumilov et al., 1995). This fact by no means implies that gas-phase photochemical theory only cannot be used to interpret ozone “mini-hole” creation during GLEs. For example, homogeneous theory describes very well the NO_x variations and ozone depletions caused by solar protons above the 40-km altitude, where model calculations and experimental results coincide (Jackman et al., 1980, 2001; Zadorozhny et al., 1992; Krivolutsky, 2001, 2003; Shumilov et al., 1995, 2003). As it is evidenced from Fig. 1, the 24 May 1990 GLE event was not a powerful one, however, during this event, ozone “mini-holes” have been observed (Shumilov et al., 1995).

In our work we consider the ion nucleation mechanism (Arnold, 1982; Hofman and Rosen, 1983) as one of the possible sources of enhanced concentration of ultrafine particles or condensation nuclei (CN) in the high-latitude atmosphere after some proton events of GLE type. Such ultrafine particles can grow by condensation to the size of typical cloud condensation nuclei (CCN) $\sim 30\text{--}80$ nm within 1 day (Kulmala et al., 1998; Birmili and Wiedensohler, 2000). We suppose this mechanism to be responsible for ozone “minihole” creation after GLEs. It should be noted that according to another hypothesis (“electrofreezing” mechanism), the variations in

cosmic rays change the charge status of aerosols that affect the ice production rate and cloud microphysics (Tinsley and Dean, 1991; Tinsley, 2000). On the surface of these particles a complex set of heterogeneous processes takes place which lead to significant ozone depletions (Prather, 1992).

Figure 2 gives a possible explanation why ozone “mini-holes” have been observed only after May 1990 GLEs in spite of the fact that these events were the “moderate” ones (see Fig. 1). Note that there were 3 GLEs from 21 to 27 May, 1990 (21, 24, 26 May) that, as would be expected have an accumulated effect on the ozone depletion observed. Large proton fluxes during powerful GLE events (4 August 1972, 19 October 1989 and 14 July 2000) have effectively reduced the nucleation rate, because for large Q (more than $100\text{ cm}^{-3}\text{ s}^{-1}$) the ion recombination rate predominates. Limiting Q values when further increase of Q leads to a J decrease are $\sim 100\text{ cm}^{-3}\text{ s}^{-1}$ and $1000\text{ cm}^{-3}\text{ s}^{-1}$ for $[\text{H}_2\text{SO}_4]$ values equal to $3 \times 10^6\text{ cm}^{-3}$ and 10^7 cm^{-3} , correspondingly (Hofman and Rosen, 1983). We would thus expect J to increase nearly linearly with Q only for moderate solar proton events (see Fig. 2 from Hofman and Rosen, 1983). The largest GLEs (4 August 1972, 19 October 1989, and 14 July 2000) could effectively reduce the nucleation rate.

We can compare these calculations with experimental results only for one (16 February 1984) GLE event, when a stratospheric aerosol enhancement has been observed at auroral latitudes (Shumilov et al., 1996; Kasatkina et al., 1999). Recent aircraft measurements have demonstrated evidence for cosmic ray-induced aerosol formation in the upper troposphere (Eichkorn et al., 2002). Furthermore, ions induced by solar protons can account for observation of unusual mid-stratospheric aerosol layers in the high-latitude stratosphere (Gerding et al., 2003), although the authors don't share this point of view. In a series of recent papers it was noted as well that air ions produced by galactic cosmic rays may play an important role in the production of new aerosol particles that may affect the cloud properties (Yu and Turco, 2000, 2001; Yu, 2002; Carslaw et al., 2002; Marsden and Lingenfelter, 2003). These model simulations seemed to support experimental data indicating the presence of a significant positive correlation between GCR intensity and the frequency of low clouds, below ~ 3.2 km (Marsh and Svensmark, 2000; Palle and Butler, 2000).

It is seen from Fig. 2 that CN concentration for May 1990 GLEs exceeds by about two orders the background level. According to Prather (1992), such a large additional CN amount can lead to catastrophic ozone total content depletions (more than 10% per day).

The results of model simulations based on the ion nucleation mechanism have shown that moderate GLEs may contribute to aerosol content significantly and initiate ozone “mini-hole” creation. The goal of our simulations was to demonstrate that the solar protons may cause, under specific combinations of some atmospheric parameters ($[\text{H}_2\text{SO}_4]$, temperature, relative humidity, pressure, and as well solar proton spectral features), considerable (up to 25%) ozone total content depletions. Therefore, the simple (linear)

correlation between solar proton flux and the values of ozone total content depletions is hardly consistent from the physical point of view.

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