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Spatial and temporal variations of small-scale plasma turbulence parameters in the equatorial electrojet: HF and VHF radar observational results

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Abstract. The spatial and temporal variations of various parameters associated with plasma wave turbulence in the equatorial electrojet (EEJ) at the magnetic equatorial location of Trivandrum (8.5° N, 77° E; dip 0.5° N) are studied for the first time, using co-located HF (18 MHz) and VHF (54.95 MHz) coherent backscatter radar observations (day-time) in the altitude region of 95–110 km, mostly on magnetically quiet days. The derived turbulence parameters are the mean electron density irregularity strength ($\delta n/n$), anomalous electron collision frequency (ν_e^*) and the corrected east-west electron drift velocity (V_{ey}). The validity of the derived parameters is confirmed using radar data at two different frequencies and comparing with in-situ measurements. The behaviour of $\delta n/n$ in relation to the backscattered power during weak and strong EEJ conditions is also examined to understand the growth and evolution of turbulence in the electrojet.

Keywords. Ionosphere (Equatorial ionosphere; Ionospheric irregularities) – Space plasma physics (Turbulence)

1 Introduction

The field-aligned ionisation irregularities, namely, the type I and type II irregularities, that are generated by the plasma instabilities in the equatorial electrojet region, have been studied extensively during the past several years using coherent HF and VHF radars (Fejer and Kelley, 1980 and references therein). These (type I and type II irregularities) are generated by the action of the two-stream and gradient drift instability mechanisms, respectively. The type I irregularities are characterized by phase velocities nearly equal to the ion-acoustic velocity (C_s) of the medium and appear more or less simultaneously at all radar elevation angles when the threshold condition is exceeded. For type I waves the observed phase velocity is independent of the zenith angle and for type

II waves, it shows a cosine θ dependence (where θ is the elevation angle). Type II irregularities have phase velocities proportional to the electron drift velocity and are observed even for very small electron drifts, whenever the electron density gradient and the electric field are in the same direction.

The characteristic features of the type I and type II irregularities present in the equatorial electrojet (EEJ) are being increasingly understood, due to several theories, both linear and nonlinear, which have been developed, taking into account the effect of plasma wave turbulence (e.g. Sudan, 1983a,b; Farley, 1985; Kudeki et al., 1987). The dominant role of large-scale waves in determining the altitude structure and dynamics of the EEJ was confirmed by the radar interferometry technique (Farley et al., 1981) and later by a combination of ground-based and in-situ measurements (Pfaff et al., 1987). Ronchi et al. (1990, 1991) carried out theoretical investigations on the modifications of large-scale electrojet structure by small-scale plasma wave turbulence. Sudha Ravindran and Reddy (1993) investigated the effect of plasma turbulence on the phase velocity of type I waves from VHF backscatter radar observations.

The high spatial resolution observations of the EEJ at Jicamarca by Farley et al. (1994) have revealed wave-like structures, which are in agreement with the theoretical investigations of Ronchi et al. (1990). The latter included background short wavelength turbulence in the long wavelength equations through a turbulent collision frequency operator and studied its effects on the large-scale dynamics of the EEJ. By this, they could explain the discrepancy between the altitude of the maximum of the experimental and theoretical EEJ current density profiles.

In another approach, Krishna Murthy and Sudha Ravindran (1994) studied experimentally the effect of plasma wave turbulence mainly on the altitude structure of the EEJ, based on limited VHF radar observations at Trivandrum. But the time evolution of the turbulence parameters, their variation in relation to the signal strength; etc have not been examined in their study. In the present study, the parameters

Table 1. Experimental specifications of VHF and HF radar systems at Trivandrum.

Parameter	VHF	HF
Operating frequency (MHz)	54.95	18
Pulse width (μs)	20	20
Pulse repetition frequency (s^{-1})	500	250
Peak power (kW)	20	50
Average power-aperture product (Wm^2)	2.2×10^5	1.25×10^7
Beam direction	Fixed beam 30° off zenith towards west	3 beam orientations zenith, 30° off zenith towards west and east
Half power two-way beamwidth (E-W)	$\sim 3.5^\circ$	$\sim 5^\circ$

$\delta n/n$, v_e^*/v_e and V_{ey} (corrected to include turbulence) have been estimated mainly for quiet days at different times, and the time evolution of these parameters has also been studied. These temporal variations are examined in association with the backscatter signal power and ΔH variations (deviations of the horizontal component of the Earth's magnetic field from the nighttime level). The behaviour of $\delta n/n$ at different altitudes in the EEJ, in relation to the backscattered power during weak and strong EEJ conditions is presented in detail. In addition, the turbulence parameters have been estimated, using both HF and VHF radar observations at several altitudes in the EEJ, for the first time, at Trivandrum on the magnetic equator.

2 Experimental setup

The VHF and HF radars at Trivandrum are pulsed, monostatic, coherent, Doppler systems operating in the backscatter mode. The scattered wave vector follows the condition for Bragg scatter, given as

$$\lambda = 2\lambda_{irr} \sin(\theta/2), \quad (1)$$

where λ is the radar wavelength, λ_{irr} is the irregularity scale size and θ is the scattering angle. For backscattering $\theta=180^\circ$; hence,

$$\lambda_{irr} = \lambda/2. \quad (2)$$

These radar systems (used in this study) operate at 54.95 and 18 MHz, respectively, corresponding to irregularity scale sizes of 2.7 and 8.3 m for backscattering. Detailed specifications, data recording and analysis procedures for VHF backscatter radar have been described in Reddy et al. (1987). The mode of data acquisition and analysis procedure is given in Manju (1997) and Janardhanan et al. (2001). The system details for both the radars are given in Table 1. The parameters estimated from the radar Doppler spectra are the power (zeroth moment), mean Doppler frequency, \bar{f}_D (first moment) and the width of the spectrum (second moment). The power is a measure of the strength of the scattering centres present in the volume illuminated by the radar antenna beam; the mean Doppler frequency gives the mean drift velocity of the

irregularities and the width of the spectrum is a gross indicator of the strength of turbulence.

HF radar employs the Doppler beam swinging (DBS) technique to point the antenna beam at three angles, viz., vertical and 30° off zenith in the east and west directions. A single 12×6 antenna array (12 in the magnetic east-west and 6 in the magnetic north-south direction) is employed for both transmission and reception. A centre-fed dipole of full wavelength is used as the antenna element. The entire array is configured in such a manner that there are six sub arrays and the beam positions at three angles are obtained by proper phasing of these sub arrays. For the present study, data recorded in the west beam position only are used, since the VHF radar beam is along the west at zenith angle 30° .

3 Data and method of analysis

The present study is carried out using a database obtained partly using VHF radar observations and partly using simultaneous HF and VHF radar observations during daytime only. Both the radars point towards west and hence they measure in the same volume at the same time (for simultaneous observations using the two radars). Section 4.1 presents the results for those days when VHF radar data sets alone are available. These data sets are obtained at 5-min intervals and six data sets are averaged for each half hour to estimate the turbulence parameters (Krishna Murthy and Ravindran, 1994). For the comparative study using both HF and VHF radar data, presented in Sect. 4.2, thirty data sets at one-minute intervals are averaged for each half hour. The averaging is carried out to minimize the effects of winds in the EEJ region with periods less than 25 min on the deduced altitude structure of the electrojet.

The radar spectral data used for the present study have been selected so that they are purely type II (uncontaminated by type I). This assures the reliability of electron drift velocity and the electric field derived from the irregularity drift velocity (Viswanathan et al., 1987). This is due to the fact that the phase velocity of type II irregularities is directly related to the electron drift velocity and hence to the driving electric field, whereas the phase velocity of type I waves becomes saturated to the ion-acoustic velocity (C_s)

irrespective of the electric field magnitude. Since the generation of type I waves at higher scale sizes is less probable, the HF radar observations are generally free from type I contamination. With the VHF radar observations, significant contribution from type I irregularity is mostly present close to noontime (Viswanathan et al., 1987) in the altitude range of 101–112 km.

The method of estimation of turbulence parameters used here involves the matching of the experimental and theoretical drift velocity profiles of the EEJ by including the turbulence contribution in the drift velocity equations, thereby removing the discrepancy between the altitude of maximum of the two profiles (Krishna Murthy and Ravindran, 1994). The radar signal is required to be present at six or more altitudes without type I contamination, for the estimation of turbulence parameters using this method. This method is evolved based on the theoretical work presented by Ronchi et al. (1990), who suggested that the main features of the large-scale observations and the dynamics of EEJ can be successfully explained when the additional electron mobility and diffusion due to background short wavelength plasma turbulence are also considered. The electric fields of the large amplitude short-scale waves perturb the electron orbits randomly, thus causing anomalous diffusion of electrons across the magnetic field. This anomalous electron diffusion is interpreted mathematically as an additional anomalous electron collision frequency (v_e^*) over and above the normal collision frequency (v_e). In the above theory the inclusion of a turbulent mobility operator in the equilibrium equation acts as an enhanced collision frequency (v_e^*). The Pedersen and Hall mobilities of electrons, μ_{Pe} and μ_{He} , respectively, are given as

$$\mu_{Pe} = q_e v_e / m_e (\Omega_e^2 + v_e^2) \quad (3)$$

$$\mu_{He} = q_e \Omega_e / m_e (\Omega_e^2 + v_e^2), \quad (4)$$

where, q_e and m_e are the electronic charge and mass, respectively, Ω_e is the electron gyrofrequency and v_e is the collision frequency of electrons with neutrals.

According to the non-linear short scale turbulence theory of Ronchi et al. (1990), the expressions for the turbulent Pedersen (μ_{Pe}^*) and Hall (μ_{He}^*) mobility parameters are given as

$$\mu_{Pe}^* = 1/2(\Psi/1 + \Psi)(\mu_{He}/\mu_{Pe})^2 \left\langle \left| \frac{\delta n(h)}{n} \right|^2 \right\rangle \mu_{Pe} \quad (5)$$

$$\mu_{He}^* = 1/2(\Psi/1 + \Psi) \left\langle \left| \frac{\delta n(h)}{n} \right|^2 \right\rangle \mu_{He}, \quad (6)$$

where

$$\Psi = v_i v_e / \Omega_i \Omega_e, \quad (7)$$

where Ω_i is the ion gyrofrequency and v_i is the collision frequency of ions with neutrals. $\left\langle \left| \frac{\delta n(h)}{n} \right|^2 \right\rangle$ are the small-scale electron density irregularity strength which is altitude dependent. $\delta n(h)$ are the fluctuations in electron density for small-scales less than 100 m as a function of altitude h , and n is the

large-scale electron density. The brackets indicate that the mean of the quantity for a range of shorter scale sizes less than 100 m is considered.

For a typical value of $\delta n(h)/n$ ($\sim 10\%$) (Pfaff et al., 1987), it is seen that $\mu_{He}^* \ll \mu_{He}$ and $\mu_{Pe}^* \geq \mu_{Pe}$. Therefore, μ_{Pe}^* is the turbulent mobility parameter of interest in the present study.

V_{ey} , the experimentally observed electron drift velocity in the E-W direction, is given by (Reddy et al., 1987):

$$V_{ey} = V_p(1 + \Psi), \quad (8)$$

where V_p is the phase velocity of type II irregularities in the E-W direction.

$$V_p = \lambda \frac{\bar{f}}{D} = 5.4 \frac{\bar{f}}{D} \quad (\text{Reddy et al., 1987}), \quad (9)$$

where \bar{f} is the mean Doppler frequency and λ is the radar wavelength.

The expression for the theoretical electron drift velocity $(V_{ey})_T$ (Reddy et al., 1987) is given by

$$(V_{ey})_T = [\rho_i / (1 + \Psi)] [E_y / B], \quad (10)$$

where E_y is the global E-W electric field and B is the geomagnetic field.

$$\rho_i = (v_i / \Omega_i) \quad (11)$$

Here the estimation of turbulence parameters is based on the assumption that the discrepancy between the observed and theoretical electron drift velocity profiles in the EEJ, is due to the non-inclusion of the effect of plasma wave turbulence, in the respective equations (for details, refer to Krishna Murthy and Ravindran, 1994). This is taken care of by including height dependent Ψ^* in the equations for V_{ey} and $(V_{ey})_T$ and equating them.

$$V_{ey} = V_p(1 + \Psi + \Psi^*) \quad (12)$$

$$(V_{ey})_T = [\rho_i / (1 + \Psi + \Psi^*)] (E_y / B). \quad (13)$$

From Eq. (5),

$$(v_e^* / v_e) \sim \mu_{Pe}^* / \mu_{Pe} = 1/2(\Psi/1 + \Psi)(\mu_{He}/\mu_{Pe})^2 \left\langle \left| \frac{\delta n(h)}{n} \right|^2 \right\rangle \quad (14)$$

where $\left\langle \left| \frac{\delta n(h)}{n} \right|^2 \right\rangle$ is assumed to be a second degree polynomial in height of the form $a+bh+ch^2$, h is the altitude and a , b , c are constants to be determined.

$$(v_e^* / v_e) \sim \mu_{Pe}^* / \mu_{Pe} = (\Psi^* / \Psi) = 1/2(\Psi/1 + \Psi) (\mu_{He}/\mu_{Pe})^2 (a + bh + ch^2) \quad (15)$$

$$\Psi^* = A(a + bh + ch^2), \quad (16)$$

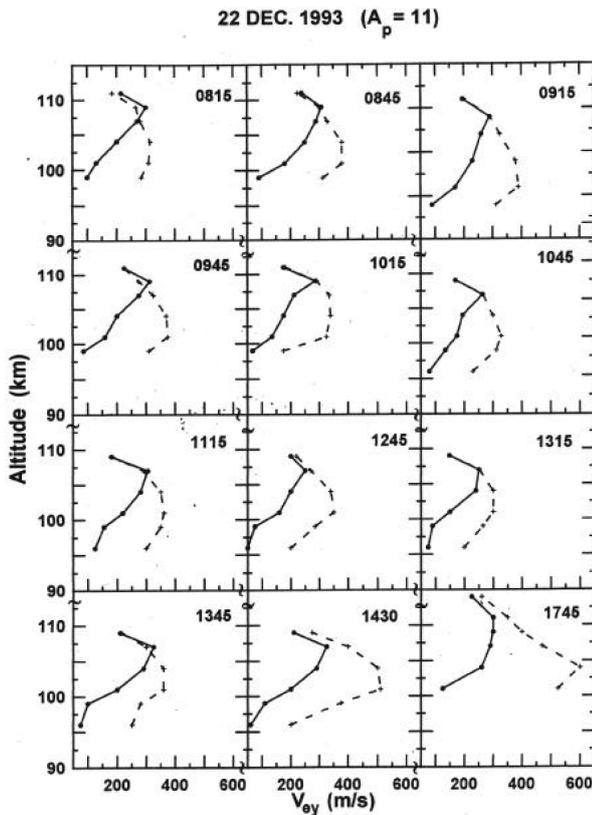


Fig. 1. Daytime altitude profiles of V_{ey} on 22 December 1993. The solid curves correspond to the uncorrected V_{ey} values (without including turbulence effects), while the dashed curves represent the corrected V_{ey} values obtained after including the effects of turbulence.

where,

$$A = 1/2[\Psi^2/(1 + \Psi)](\mu_{He}/\mu_{Pe})^2, \quad (17)$$

A is calculated at different heights using values obtained from MSIS 86 (Hedin, 1987) model for neutral densities and temperatures.

Equating (12) and (13) and substituting for Ψ^* results in

$$V_p = \rho_i/[1 + \Psi + A(a + bh + ch^2)]^2[E_y/B]. \quad (18)$$

Taking the ratio of V_p (at 2 altitudes) for 3 sets of data at six altitudes, E_y is eliminated and the equations are solved to obtain the values of a , b and c . Thereafter, the values of E_y , V_{ey} (with turbulence parameter Ψ^*), v_e^* and $\delta n/n$ are estimated.

The spatial and temporal variations of the parameters such as V_{ey} (including Ψ^*) and $\delta n/n$ (obtained using the above method) are presented in Sect. 4. The behaviour of $\delta n/n$ viz-a-viz the backscattered power is also investigated.

4 Results

The data used in the present study are relatively free from contamination by type I irregularities, in view of the fact that winter and summer months have been selected. It should be

mentioned here that based on our experience with VHF radar observations, it is known that for low solar activity conditions during winter and summer months, the type I contamination is at a minimum (for details, refer to Vikramkumar et al., 1984). The days selected are relatively magnetically quiet (as is evidenced by the geo-magnetic indices), except for 21 December 1993 ($A_p=20$), which is somewhat disturbed. On 21 December 1993, type I contamination is present in the Doppler spectrum between 09:15 and 10:45 IST (Indian Standard Time corresponding to 82.5° E longitude leads UT by 5.5 h). All times given here correspond to IST. On 15 June 1995 ($A_p=6$) data are available during 08:30–16:30, with type I contamination from 10:00–12:00. Hence, turbulence parameters are not estimated during the 10:00–12:00 period. The results presented for 22 December 1993 and 15 June 1995 are obtained using VHF radar observations while those presented for 25 and 26 October, 1995 (equinoctial period) are based on simultaneous HF radar and VHF radar observations.

4.1 Altitude profiles of V_{ey} and $\delta n/n$ in the EEJ – VHF radar observations

As an example, Fig. 1 shows V_{ey} profiles for the daytime on 22 December 1993 ($A_p=11$), which is a relatively magnetically quiet day without type I presence, even at higher altitudes. Hence, data were available at the required six altitudes for estimating turbulence parameters reasonably well. As mentioned earlier, the profiles shown here are the average profiles obtained mostly at each half hour. The solid curves correspond to the deduced V_{ey} profiles without turbulence effects, while the dashed curves correspond to the corrected V_{ey} profiles obtained after including the effects of turbulence. The time in hours corresponding to each panel is given at the top.

It is seen from Fig.1 that in all cases the altitude of the maximum of the corrected profile is lower than the deduced profile (without turbulence effects). Also, it should be mentioned that the theoretical V_{ey} profile (without considering the effect of turbulence) peaks around 97 km (not shown in the figure) (Richmond 1973a,b), which is lower than the corrected profile. Thus, the effect of including v_e^* is to shift the peak of the theoretical profile to a higher level. As can be seen, the magnitudes of the corrected V_{ey} are larger than the deduced values (without turbulence effects), especially below and around the altitude of maximum V_{ey} . This is due to the inclusion of Ψ^* for the corrected profile. However, the V_{ey} values (including turbulence) are larger mostly in the late afternoon hours in comparison with those in the pre-noon hours, indicating an increase in turbulence strength from pre-noon to afternoon.

The altitude profiles of $\delta n/n$ obtained, which is directly related to Ψ^* , refer to Eq. (16), corresponding to these times are shown in Fig. 2. The value is less than 6% up to 13:45 but later increases to $\sim 10\%$. It is seen that on occasions, it spans a very narrow range and at other times it shows marked altitude variations. Figure 3 shows the altitude profiles of V_{ey}

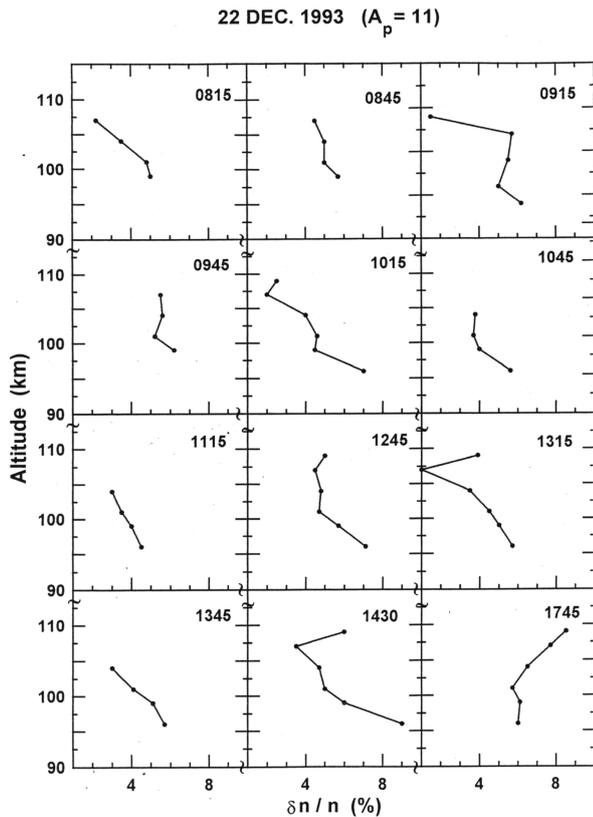


Fig. 2. Deduced altitude profiles of $\delta n/n$ corresponding to the times shown in Fig. 1.

(deduced without Ψ^* and corrected with Ψ^*) and $\delta n/n$ in the left and right panels, respectively, for the pre-noon, close to noon and afternoon times, of 15 June 1995. In Fig. 3, the V_{ey} profiles exhibit similar behaviour as in Fig. 1. The $\delta n/n$ values are found to lie in the range of 2–6% for the pre-noon and noon times with higher values at 15:45. This pattern of variation in $\delta n/n$ is seen for all the days examined. There is a consistent increase in $\delta n/n$ in all cases during the late afternoon hours.

4.2 Comparison of turbulence parameters estimated from HF and VHF radar observations

We have operated the HF and VHF radars simultaneously on a number of days and compared the turbulence parameters estimated from the two sets of observations. Figure 4 shows the altitude profiles of v_e^*/v_e , $\delta n/n$, V_{ey} and relative power at 11:15 on 25 October 1995. The solid line corresponds to the parameters estimated from the HF radar observations and the broken line to those from VHF radar observations. It is seen that there is a good agreement between the values estimated from the two radars, though VHF radar is sensitive to 2.7-m scale size irregularities while HF radar returns are from 8.3-m scale size irregularities. The $\delta n/n$ values are in the range of 2–4%. The $\delta n/n$ values are nearly the same for the two radars, although the scale sizes probed are different.

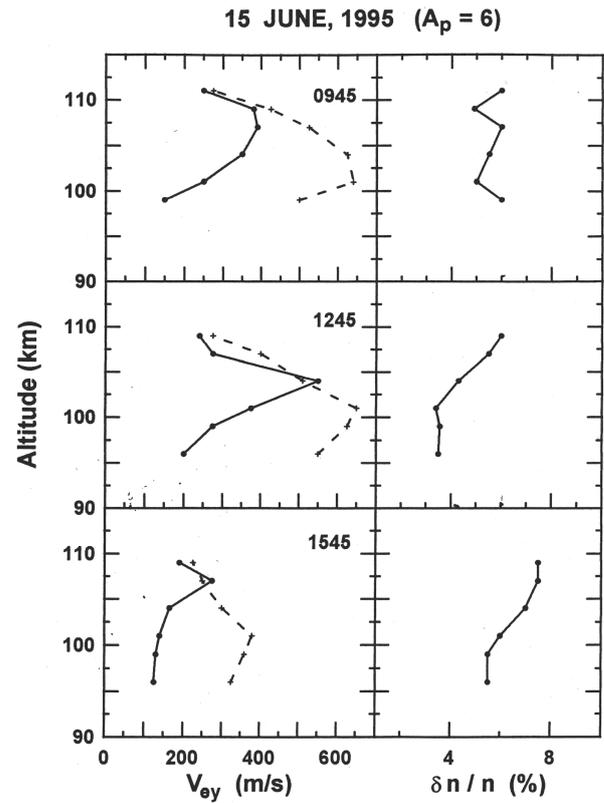


Fig. 3. Uncorrected (solid) and corrected (dashed) altitude profiles of V_{ey} and $\delta n/n$ for the pre-noon, close to noon and afternoon on 15 June 1995.

The relative power and V_{ey} are seen to peak at 101 km in both cases, though the magnitudes differ slightly.

4.3 Time variation of different turbulence parameters – VHF and HF radar observations

First we present the observations for 22 December 1993 and 15 June 1995, based on VHF radar observations. Figure 5 shows the time variation of corrected V_{ey} , relative power, $\delta n/n$ (at 104 km, close to the peak altitude of EEJ) and ΔH at Trivandrum for 22 December 1993 from 08:00–18:00. The relative power is obtained from the backscatter signal power of the Doppler spectrum for a given range bin. It is seen from the figure that V_{ey} and the relative power show an increase from $\sim 14:00$ and reach their maxima around 16:00. Correspondingly the $\delta n/n$ values also increase from 3–6% and continue to increase beyond 16:00. A similar pattern is obtained at other altitudes, too. The ΔH values at Trivandrum, shown in the top panel, also reveal a significant increase during this period and thereafter a decrease to near zero by 18:00. The increase in ΔH around 16:00 is in contrast to the typical ΔH variations observed on a quiet day, where the peak value is attained close to noon.

The temporal variations of the estimated east-west electric field E_y (taking into consideration v_e^*) on this day are shown in Fig. 6. The values increase sharply between 14:00

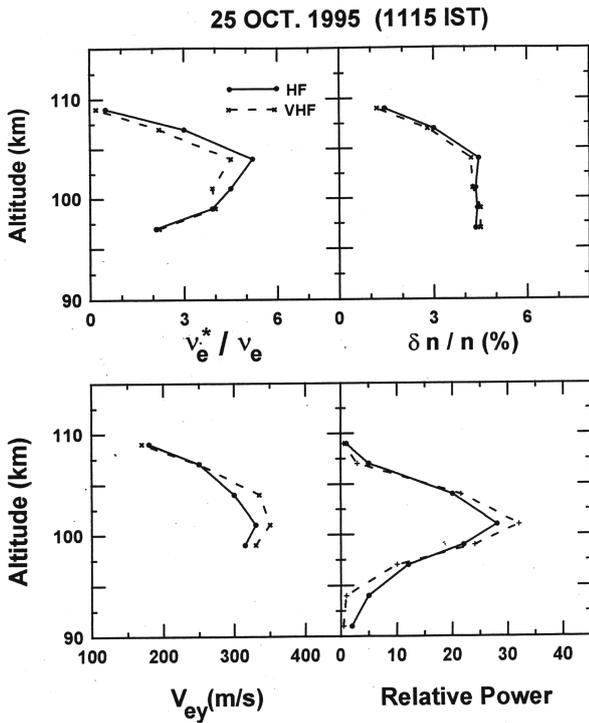


Fig. 4. Altitude profiles of v_e^*/v_e , $\delta n/n$, corrected V_{ey} and relative power (in arbitrary units) corresponding to 11:15 on 25 October 1995, from HF and VHF radar observations.

and 16:00, attaining a peak value of ~ 2.6 mV/m at 16:00. The height integrated Cowling conductivity Σ_3 generally decreases during the afternoon hours. It is known that ΔH is a function of Σ_3 and E_y , and hence the observed increase in ΔH may quite reasonably be attributed to the large magnitude of E_y . Another important aspect is that after 16:00, V_{ey} , relative power and ΔH values decrease, while $\delta n/n$ continues to increase to a value of $\sim 8\%$ at 18:00.

The time variations of these parameters at 99 km for the same day are shown in Fig. 7. It is seen that in the post-noon period, the power decreases from 13:00 onwards up to 17:45. However, V_{ey} increases up to 15:45 and then decreases. Thus, on this day $\delta n/n$ values are also the relatively high, mostly independent of the variations in V_{ey} and power in the late afternoon and pre-noon hours. Similar is the case at 101 and 107 km. The time variations of V_{ey} , relative power, $\delta n/n$ (for 104 km) and ΔH for 15 June 1995 are shown in Fig. 8. The gap in the figure corresponds to the presence of type I in the radar spectrum. It is seen that $\delta n/n$ values show similar variations, as shown by V_{ey} and power from 12:45 to 13:45. Beyond 13:45, power and V_{ey} decrease, while $\delta n/n$ shows an increase from $\sim 6\%$ to 8% by 16:30. ΔH values do show similar variations with V_{ey} and $\delta n/n$, especially when EEJ is strong. Again it is seen that in the pre-noon time the $\delta n/n$ values are high while the power and V_{ey} have not yet built up.

Time variations of V_{ey} , relative power, $\delta n/n$ (all obtained from HF radar observations) and ΔH for 25 and 26 October

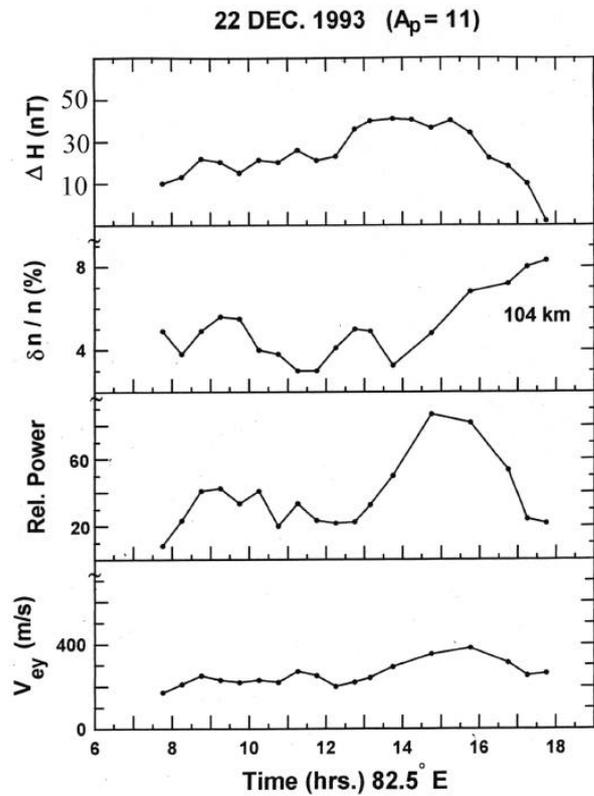


Fig. 5. Daytime variation of corrected V_{ey} , relative power, $\delta n/n$ at 104 km and ΔH at Trivandrum on 22 December 1993.

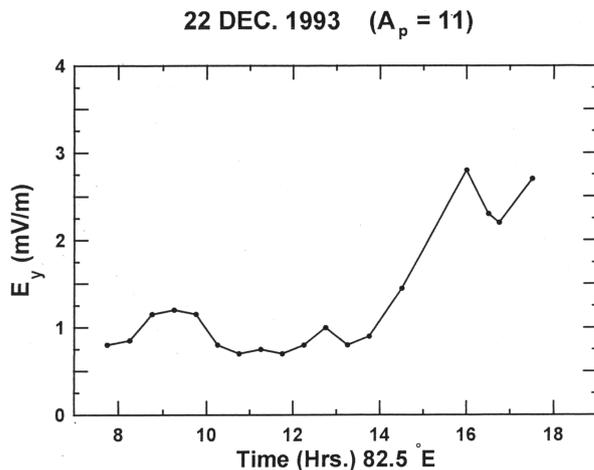


Fig. 6. Daytime variation of the estimated E_y (taking into account v_e^*) on 22 December 1993. The sharp increase in the electric field magnitude between 14:00 and 16:00 is to be noted in particular.

1995 are shown in Figs. 9 and 10, respectively. In Fig. 9, ΔH values attain their peak at $\sim 10:30$, nearly close to the time of maximum in V_{ey} (this feature can also be noticed in Fig. 10 for 26 October 1995). Prior to 09:00, the relative power is less, even while $\delta n/n$ values are relatively high. Similarly we see high $\delta n/n$ values in the late afternoon hours, while the power and V_{ey} are low or decreasing. This behaviour in

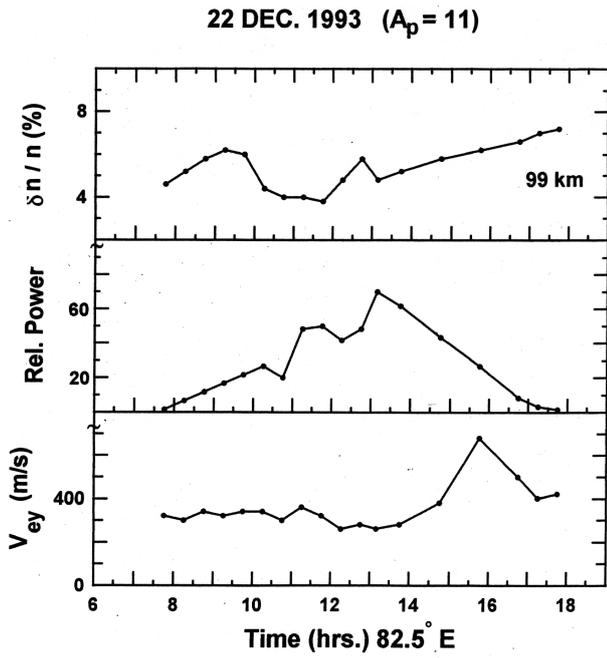


Fig. 7. Daytime variation of corrected V_{ey} , relative power and $\delta n/n$ at 99 km on 22 December 1993.

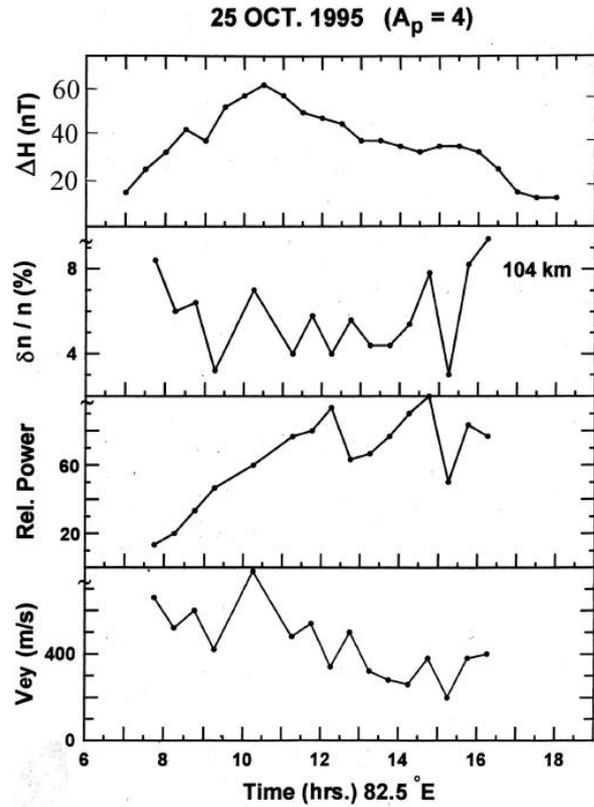


Fig. 9. Same as Fig. 5, but for 25 October 1995.

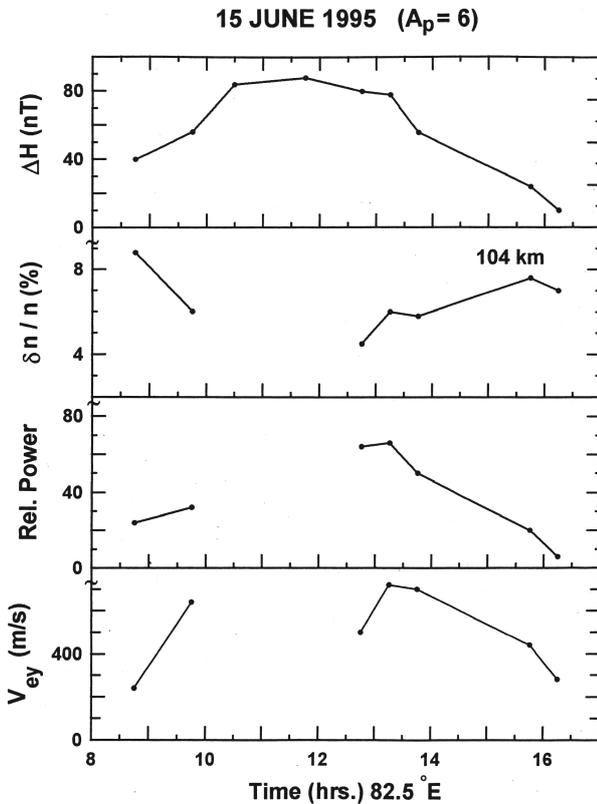


Fig. 8. Same as Fig. 5, but for 15 June 1995.

the pre-noon and post-noon periods is a consistent feature for all the days examined. It is to be noted from Fig. 10 that the parameters V_{ey} and $\delta n/n$ are shown only from 11:15, in

view of the fact that slight type I contamination was present in the HF radar data prior to that time. In such cases, the data give rise to unrealistically high values of V_{ey} , $\delta n/n$ and v_e^*/v_e . Further, in Fig. 10, the parameters V_{ey} , power, $\delta n/n$ and ΔH show similar variations up to $\sim 13:30$. Beyond that time, $\delta n/n$ shows an increase, while V_{ey} and power show a decreasing trend. The importance and the implications of the above observations are discussed below.

5 Discussion

In the present study, the parameters $\delta n/n$, v_e^*/v_e and V_{ey} (corrected to include turbulence) have been estimated mainly for quiet days at different times. The method of deducing these parameters used in the present study is based on the non-linear short scale turbulence theory developed by Ronchi et al. (1990) for scale sizes < 100 m and hence the values deduced using this method correspond to a range of turbulent scale sizes below 100 m. The deduced $\delta n/n$ values using both HF and VHF radars lie in the range of 0–10%. For similar scale sizes Pfaff et al. (1987) have reported $\delta n/n$ values in the range 0–10% using in-situ measurements in Peru for electrojet altitudes. Prakash et al. (1972) have shown using in-situ rocket measurements at Trivandrum that irregularity amplitudes for scale sizes < 15 m are 0–2% while that for scale sizes in the range 30–300 m are 0–20% for electrojet altitudes. Thus, the values deduced in the present study seem to

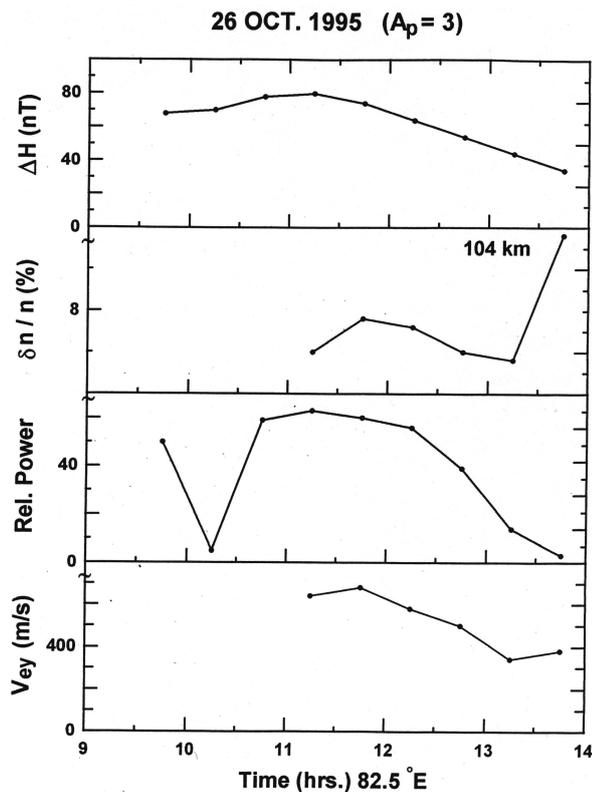


Fig. 10. Same as Fig. 5, but for 26 October 1995.

agree well with different in-situ measurements, as the scale sizes probed by the VHF and HF radars are 2.7 m and 8.3 m, respectively.

The use of HF radar data, in addition to VHF radar data for estimating turbulence parameters yields $\delta n/n$ values in the same range, thereby further confirming the validity of the method used in the present study. Moreover, the approach of using HF and VHF radar observations has an added advantage that VHF radar data is contaminated by the presence of type I irregularities during strong electrojet times (Vikramkumar et al., 1984), whereas during such times turbulence parameters can be estimated from HF radar data, as it is less susceptible to type I contamination.

The phase velocity of type II irregularities, which indicates the strength of the electrojet current, is proportional to the radar measured mean Doppler frequency (\bar{f}_D), as mentioned earlier (Reddy et al., 1987). Vikramkumar et al. (1984) have presented the typical daytime variation of the radar measured mean Doppler frequency. They have shown that the electrojet current and signal strength (power) show a rapid increase in the morning (building-up phase), a broad maximum around 11:00 (strong EEJ phase), and a decrease in the afternoon hours (decaying phase). The present study shows that there is a direct relation between $\delta n/n$ and power when the EEJ is strong. During those times when the EEJ is building up (before 09:30) and decaying (after 13:30) the $\delta n/n$ values are high, while the radar-observed backscattered power is

low. During such times (morning building up and afternoon falling off times of the EEJ), the background electron density (n) values are relatively low. These low n values result in an increase in $\delta n/n$ for a given δn value. This could possibly explain the high $\delta n/n$ observed during weak EEJ times. Moreover, the backscattered power variations obtained using the VHF radar correspond to irregularities in the wavelength range of ~ 3 m. During non-steady states of the EEJ, probably energy has not been fully transferred to the very small-scale sizes of ~ 3 m from relatively larger scale sizes. This can also result in high $\delta n/n$ (which corresponds to all scale sizes below 100 m) while power (corresponding to 3 m scale) is low. When the EEJ gains strength, energy fully reaches the lowest of scale sizes “seen” by the radar and observed power becomes high. As the EEJ gains strength by about 09:30 the background electron density (n) increases and $\delta n/n$ values reduce with respect to the earlier values.

6 Conclusions

The study carried out using co-located HF and VHF radar observations at Trivandrum brings out the following aspects.

1. The deduced values of $\delta n/n$ agree well with the values obtained by other workers using in-situ observations.
2. The $\delta n/n$ values obtained using simultaneous HF and VHF radar observations lie in the same range, thus validating the present method of estimation of turbulence parameters.
3. A direct relationship is observed between $\delta n/n$ and backscattered power during strong EEJ times. During building up and decaying phases of EEJ, the $\delta n/n$ values are high while the power is low. This shows that during the strong phase of the EEJ energy is carried to the lowest of scale sizes. During the weaker phases of the EEJ, the power values are low while $\delta n/n$ is high, as the energy has not been carried to the lowest of scale sizes and a steady state has not been reached. Thus, the comparison of $\delta n/n$ and backscattered power sheds light on the growth and evolution of turbulence in the electrojet plasma.
4. The absence of type I contamination in the HF radar data facilitates the estimation of turbulence parameters even when VHF radar data is contaminated by type I irregularities.

The study brings out the utility of the present method in estimating plasma turbulence parameters using HF and VHF radar data. Further, the behavior of $\delta n/n$ and backscattered power at different times of the day shows the diurnal evolution of plasma turbulence in the EEJ. This aspect can in the future be studied further to understand the differences in the growth and development of plasma turbulence under different geophysical conditions, such as magnetically highly disturbed, moderately disturbed, and quiet times. Another aspect that can be pursued is the study of the seasonal variations

of the plasma turbulence parameters, if any. The database generated from such studies can be potentially used to improve existing electrojet models.

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